

# New Techniques and Technologies for Remote Sensing of Volcanic Ash and SO<sub>2</sub> Gas

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# Overview

Types of volcanic activity

Remote sensing of “volcanic clouds”

Ash detection – IR techniques

Examples

SO<sub>2</sub> gas detection – IR techniques

*The emphasis will be on the use of remote sensing to measure and assess these effects*







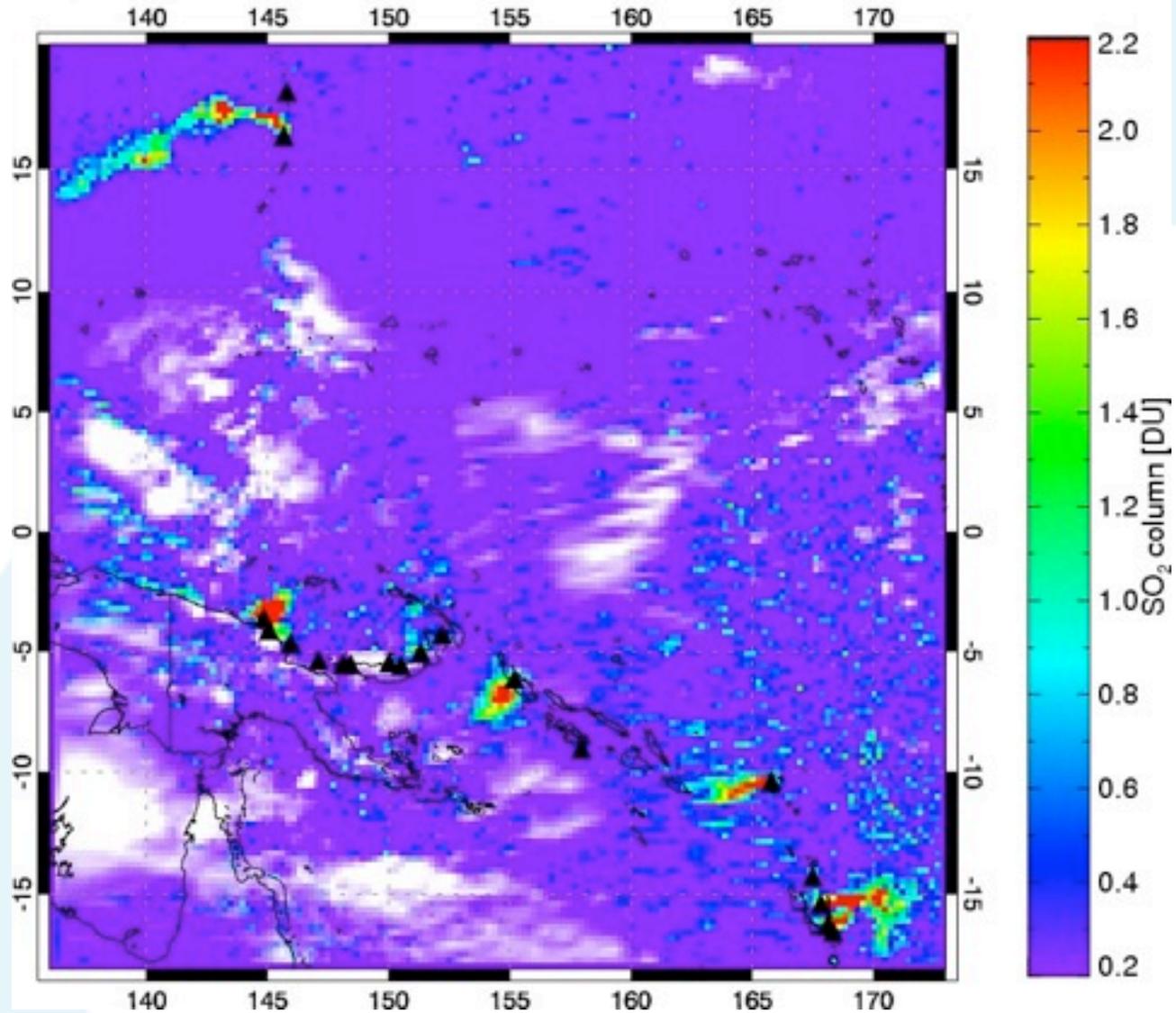






Aura/OMI - 04/23/2006 02:20-05:38 UT

SO<sub>2</sub> mass: 21.646 kt; Area: 740221 km<sup>2</sup>; SO<sub>2</sub> max: 7.36 DU at lon: 145.21 lat: -3.41



# The Volcanic Explosivity Index (VEI)

<u>VEI</u>	<u>Description</u>	<u>Volume</u> (m <sup>3</sup> )	<u>Height</u> (km)	<u>Type</u>	<u>Number</u>
0	Non-explosive	10 <sup>4</sup>	0.1	Hawaiian	699
1	Small	10 <sup>6</sup>	1	Strombolian	845
2	Moderate	10 <sup>7</sup>	5	Vulcanian	3477
3	Moderate-large	10 <sup>8</sup>	15	Vulcanian	869
4	Large	10 <sup>9</sup>	25	Plinian	278
5	Very large	10 <sup>10</sup>	>25	Ultra-Plinian	84
6		10 <sup>11</sup>	>25	Ultra-Plinian	39
7		10 <sup>12</sup>	>25	Ultra-Plinian	4
8		10 <sup>13</sup>	>25	Ultra-Plinian	1?

VEI~2  
Every 2 weeks

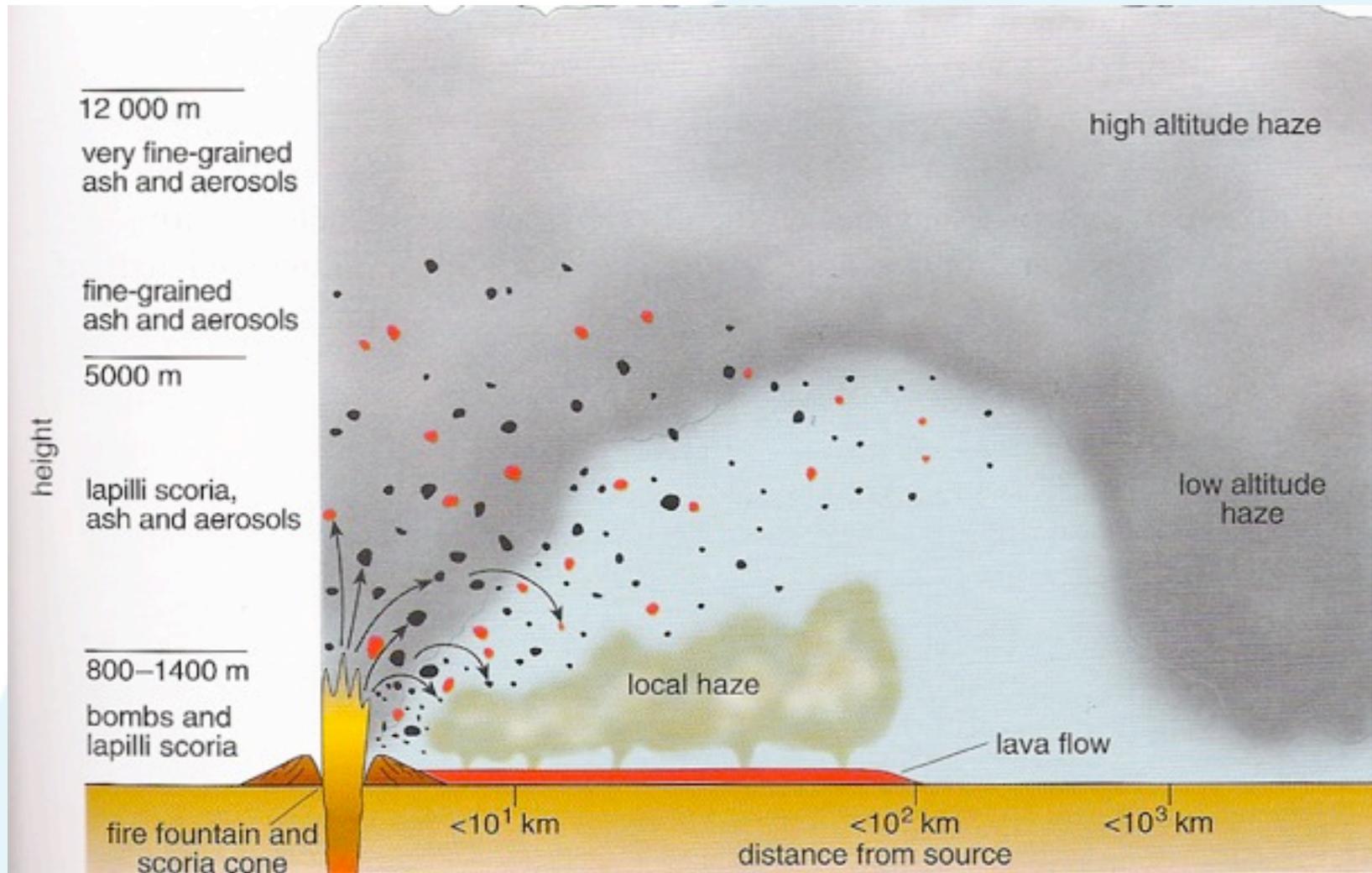
VEI~6  
Krakatau

VEI~7  
Tambora

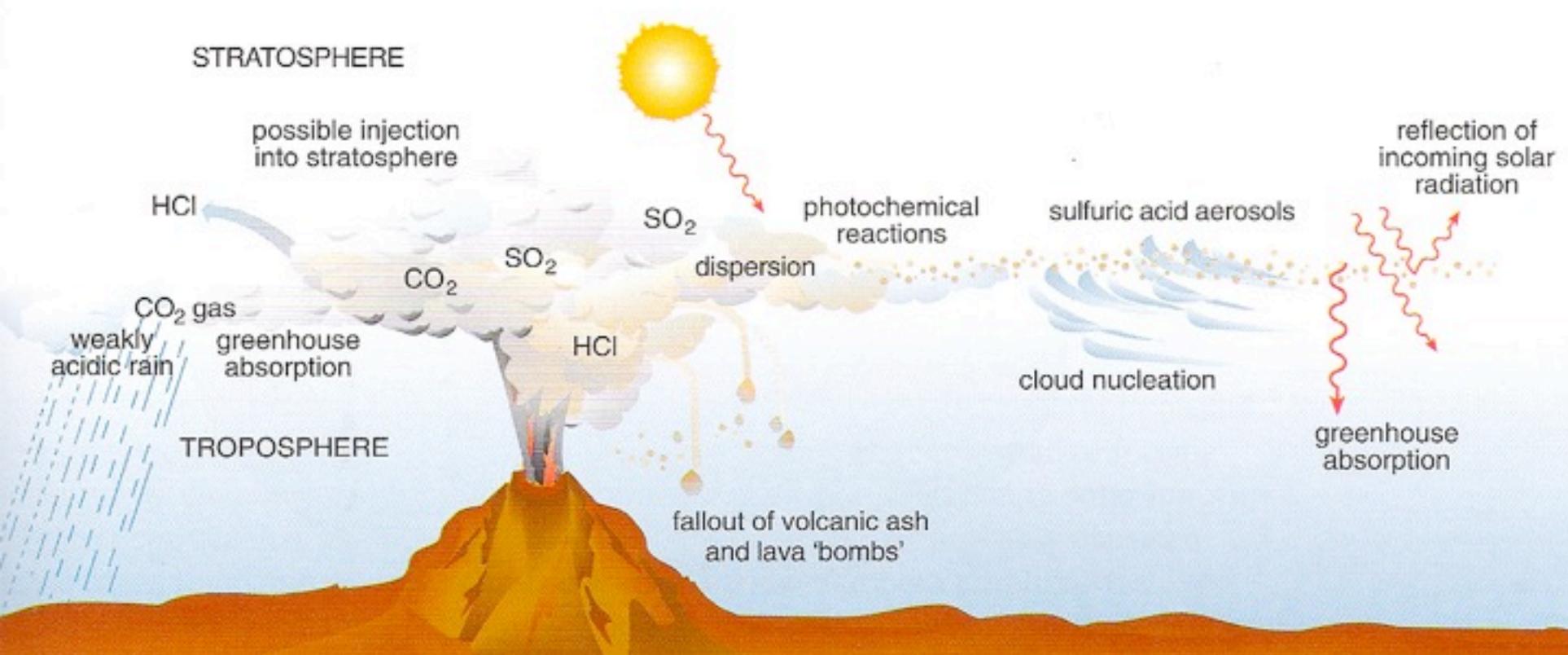
VEI~5  
MtStHelens



# Local effects

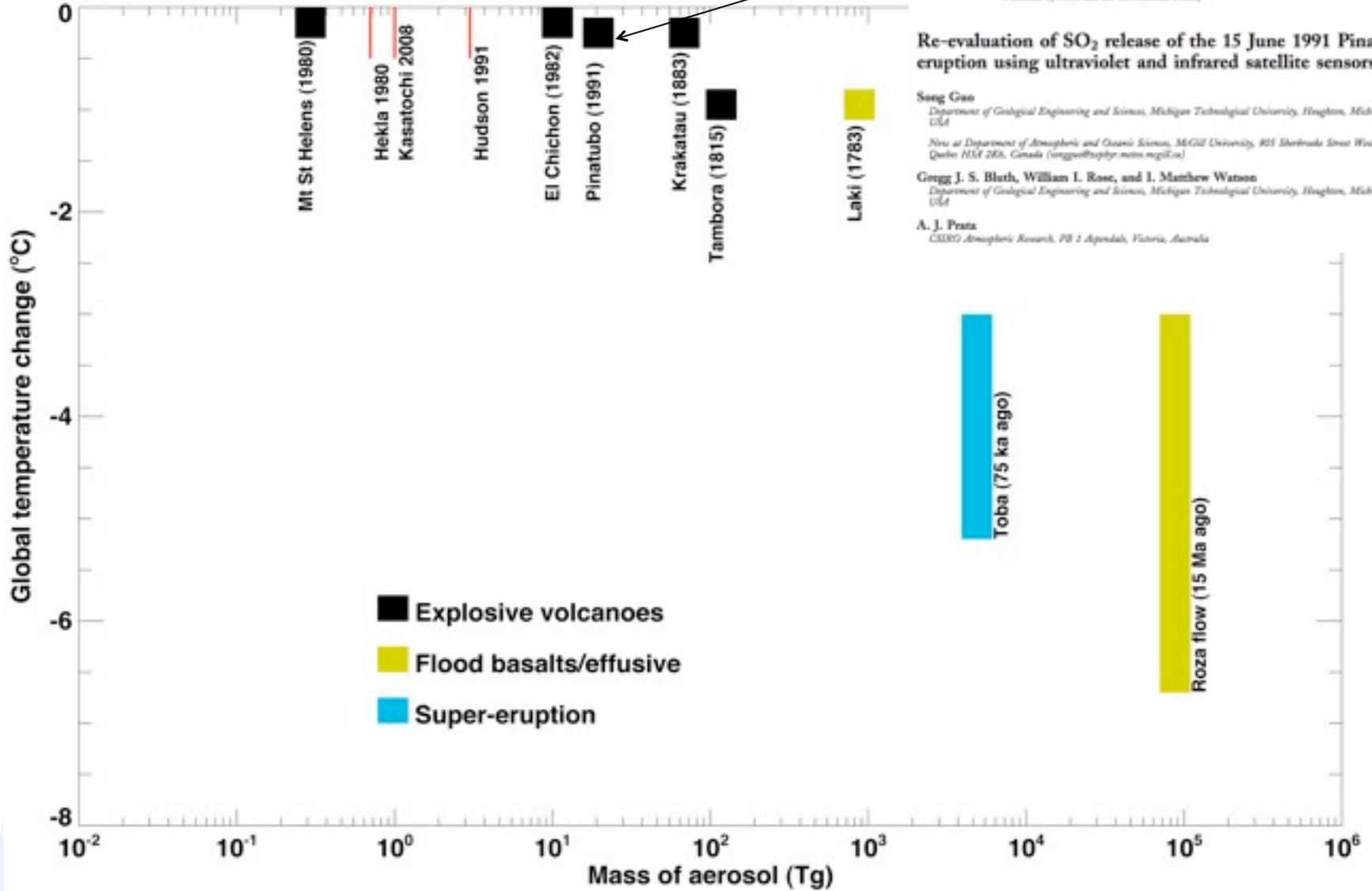


# Volcanoes and Climate



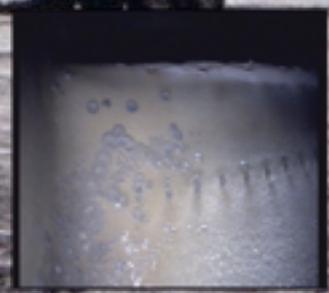
# Cooling Effects

~18 Tg



# AVIATION HAZARDS

Quito airport



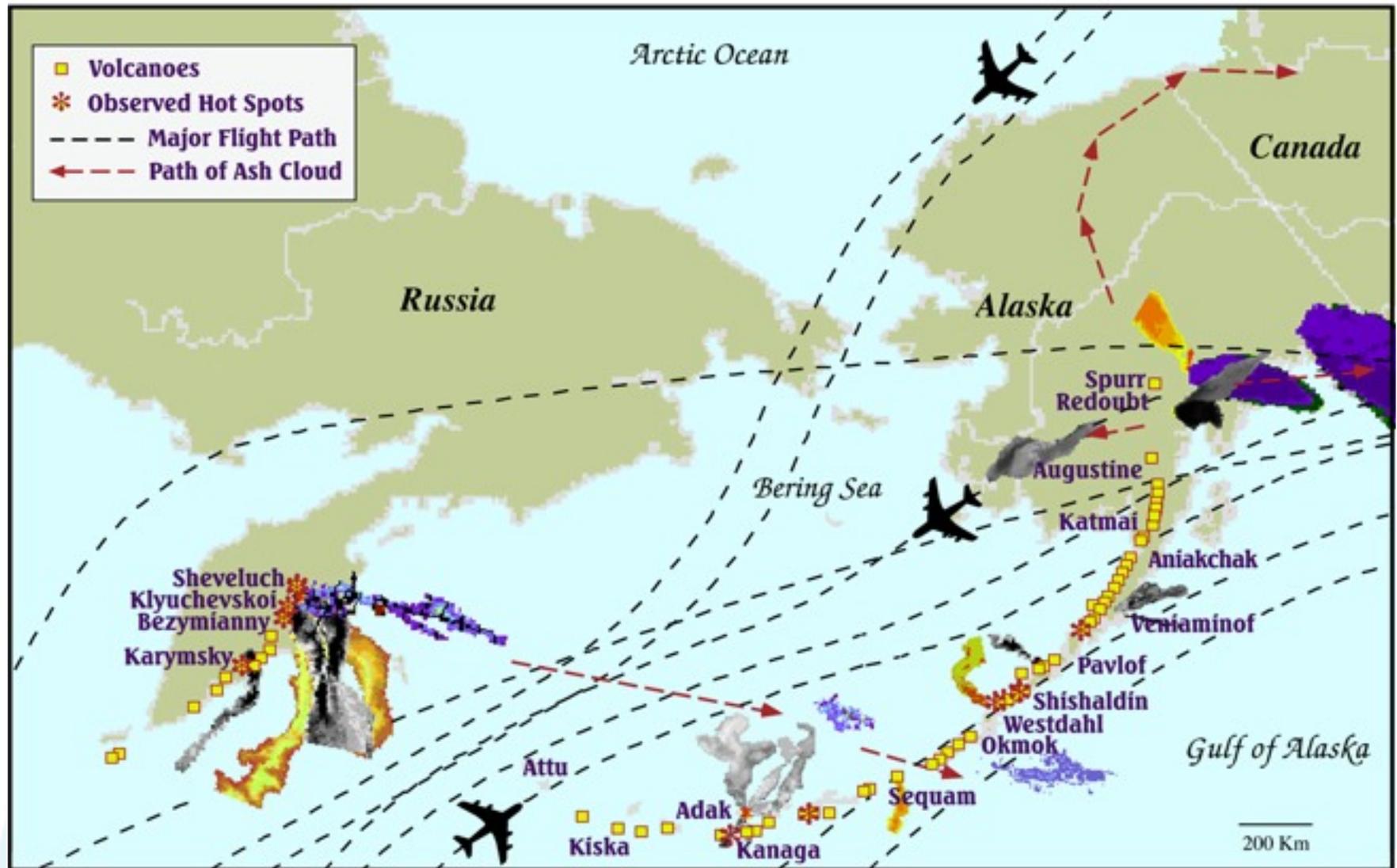
● Plugged cooling holes

● Erosion of edges

● Ash build-up

PASI, Costa Rica

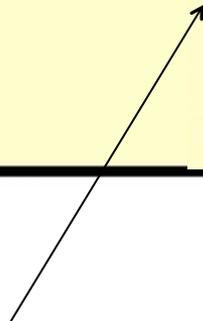
# Eruptions Affecting North Pacific Air Routes During the 1990's



# Growth in aviation traffic

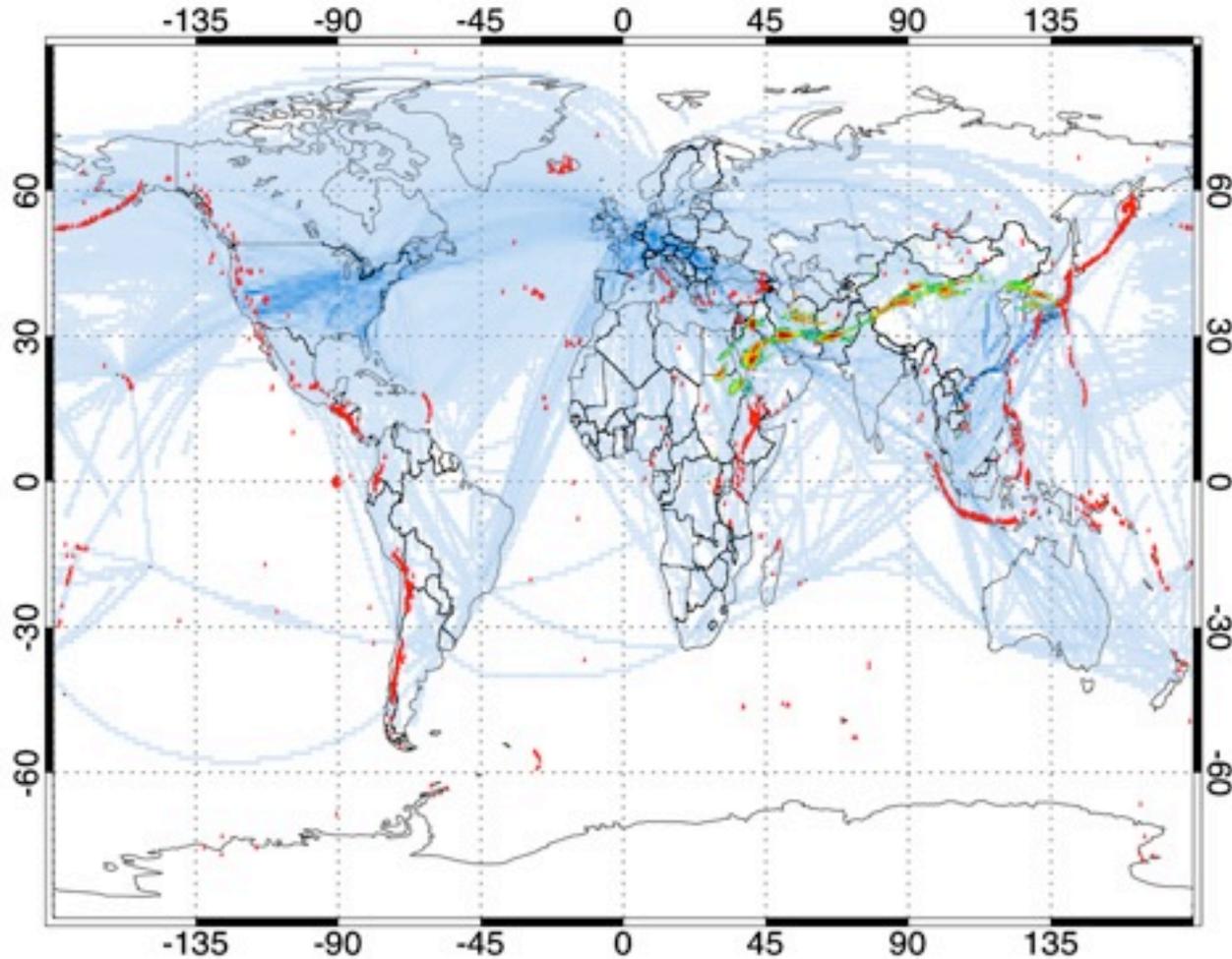
Region	Annual average growth rate of passengers (percentage)	
	2001-2005	
Europe	2.3	4.5
Africa	3.3	5.0
Middle East	6.4	4.0
Asia and Pacific	4.1	7.5
North America	1.8	5.0
Latin America and Caribbean	2.9	4.5
World	2.7	6.0

Freight & mail  
(% change in ton-km)



Source: *International Civil Aviation Organisation*

# Global aviation threat



# Remote sensing of volcanic clouds

**OMI** -  $\text{SO}_2$ , aerosols, BrO

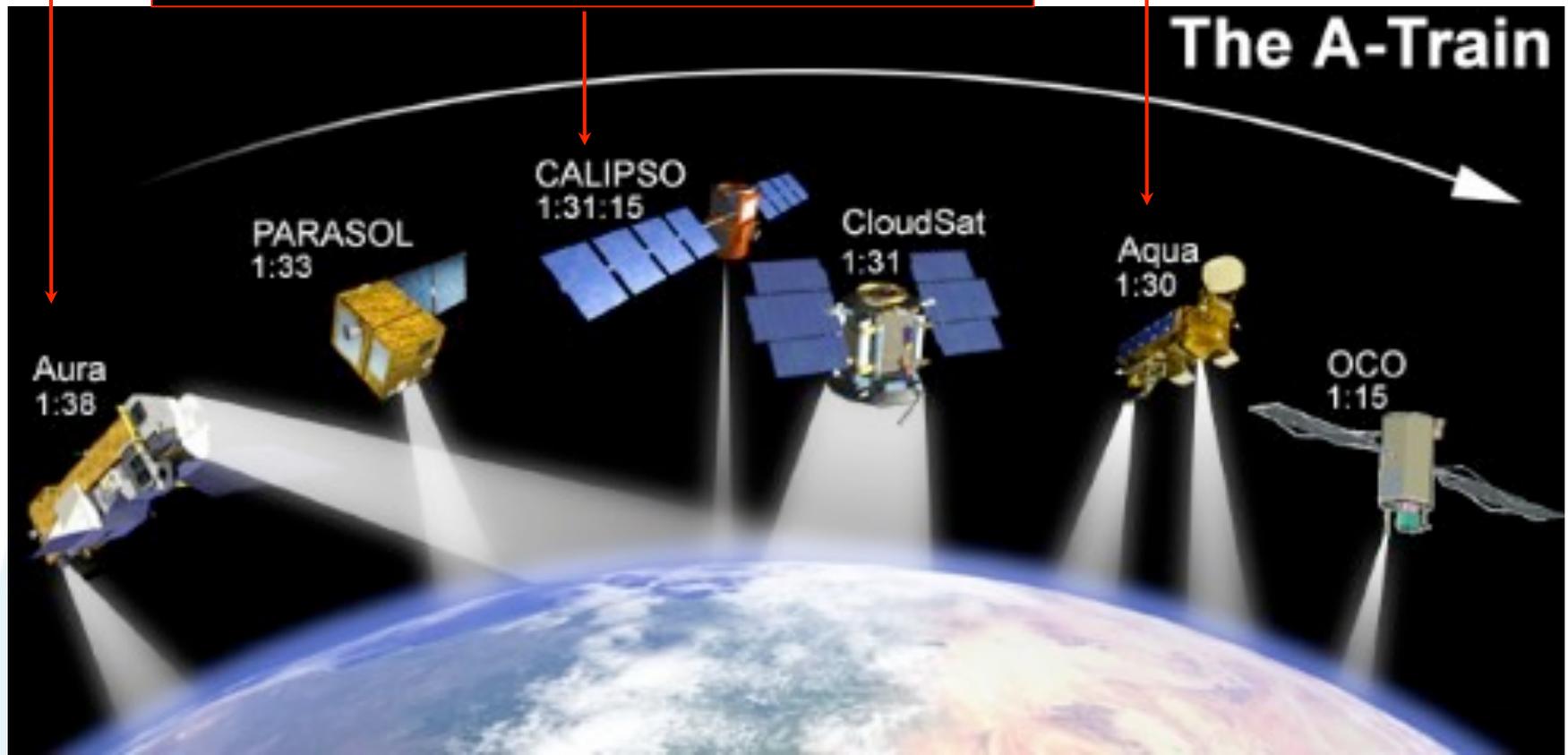
**TES** -  $\text{SO}_2$ , HCl

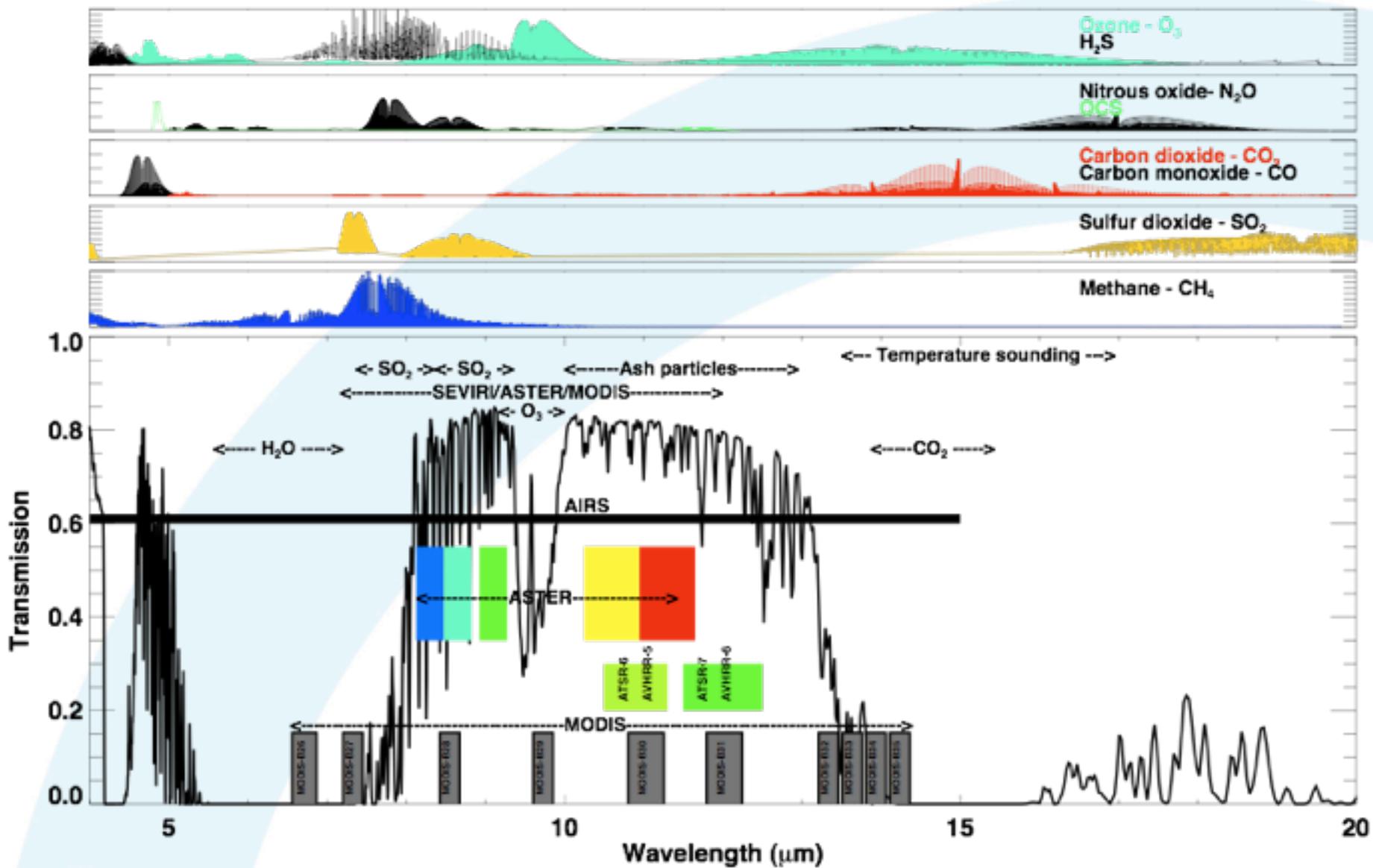
**MLS** - strat.  $\text{SO}_2$ , HCl

**MODIS** -  $\text{SO}_2$ , ash, sulfate

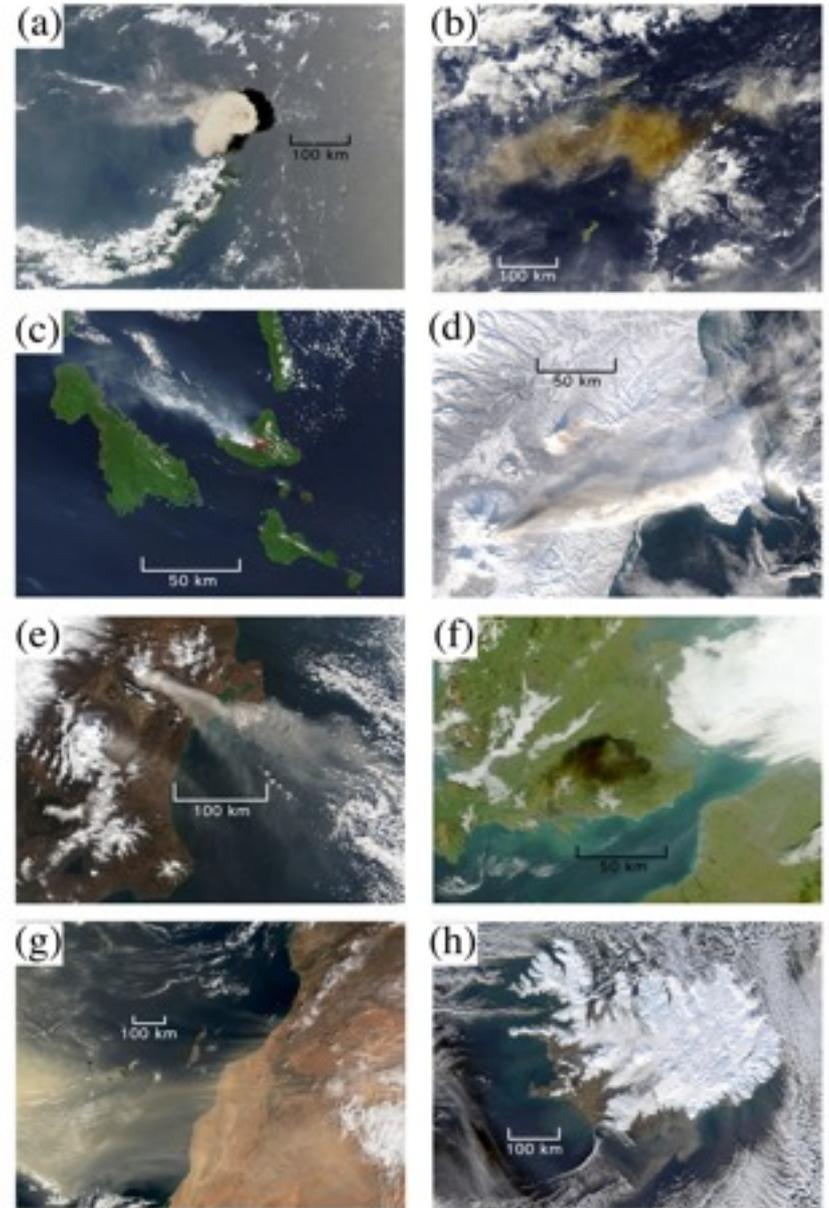
**AIRS** - UTLS  $\text{SO}_2$ , aerosols,  $\text{SO}_2$  profile?

**CALIPSO** - cloud height, aerosol type





# Detection



# Volcanic ash – Detection and discrimination

Prata, A. J., 1989, Infrared radiative transfer calculations for volcanic ash,  
 Geophys. Res. Lett., 16(11), 1293–1296

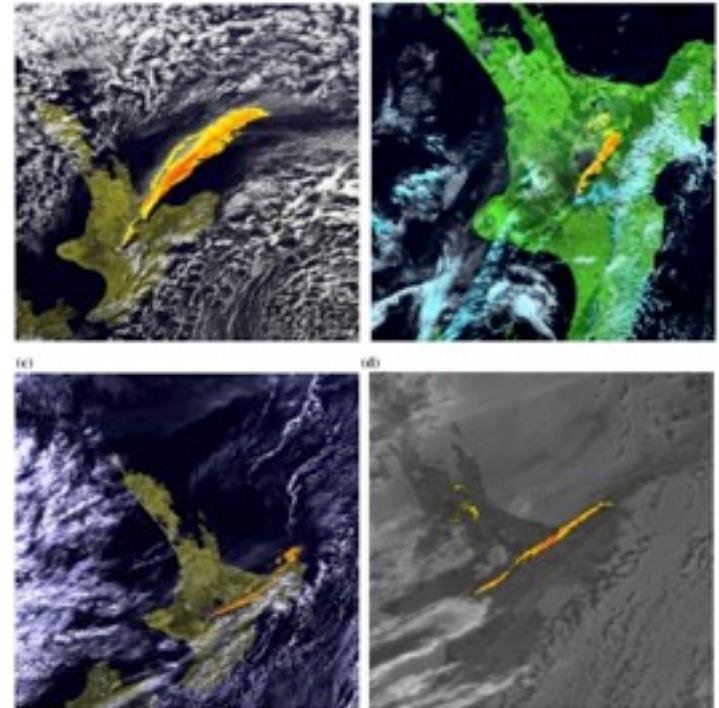
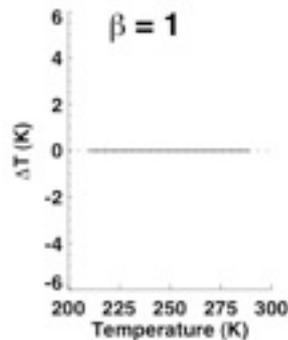
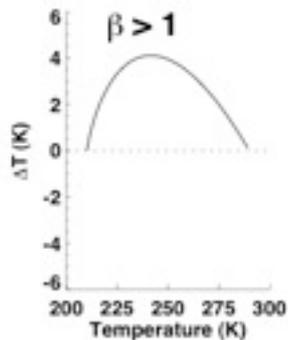
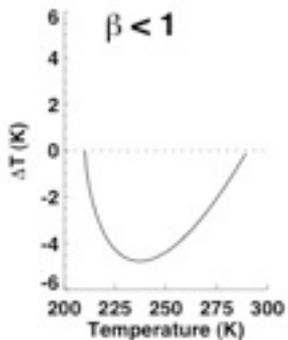
“reverse absorption” algorithm

$$T_{11} - T_{12} < 0$$

$$\Delta T = \Delta T_c [X - X^\beta],$$

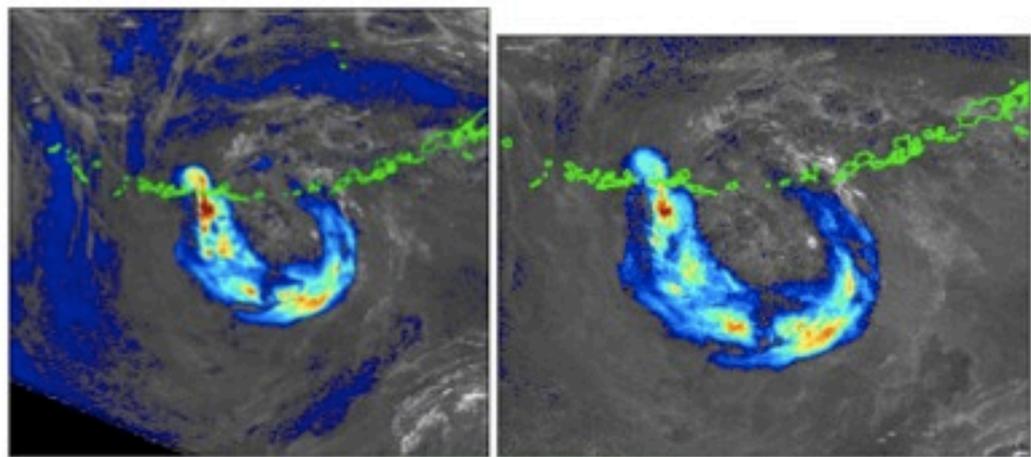
$$X = 1 - \frac{\Delta T_4}{\Delta T_c},$$

$$\Delta T = T_4 - T_5, \quad \Delta T_c = T_s - T_c, \quad \Delta T_4 = T_s - T_4.$$

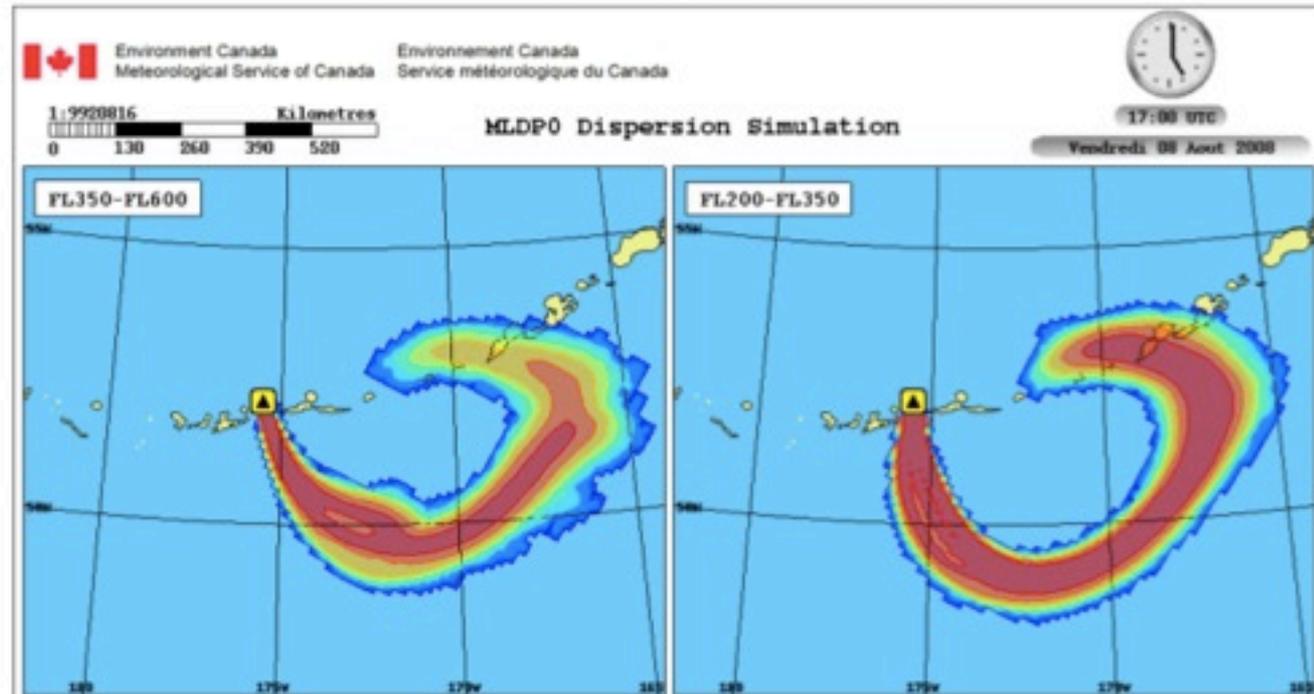


Operational use of the  
“reverse absorption” algorithm

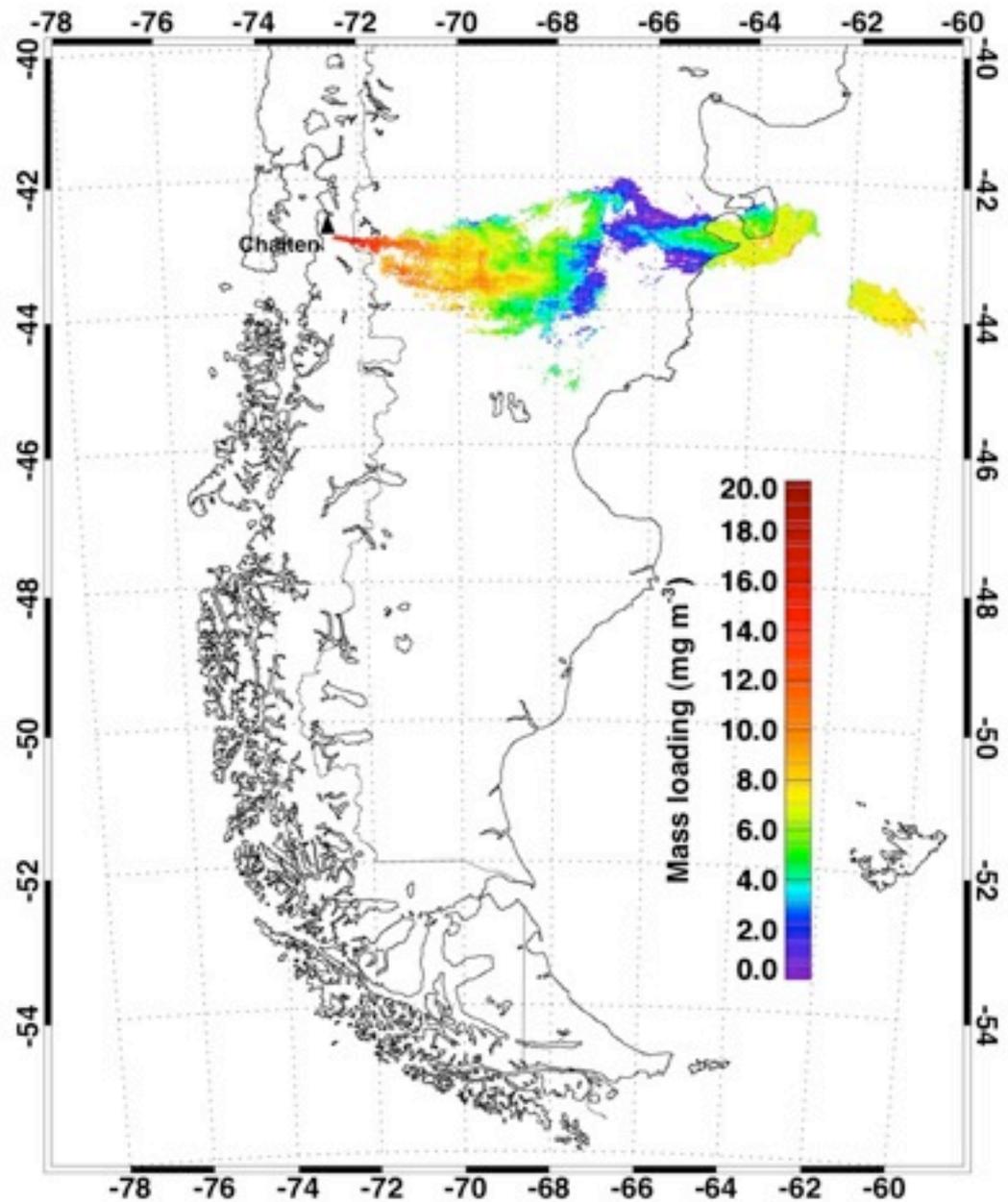
Kasatochi eruption



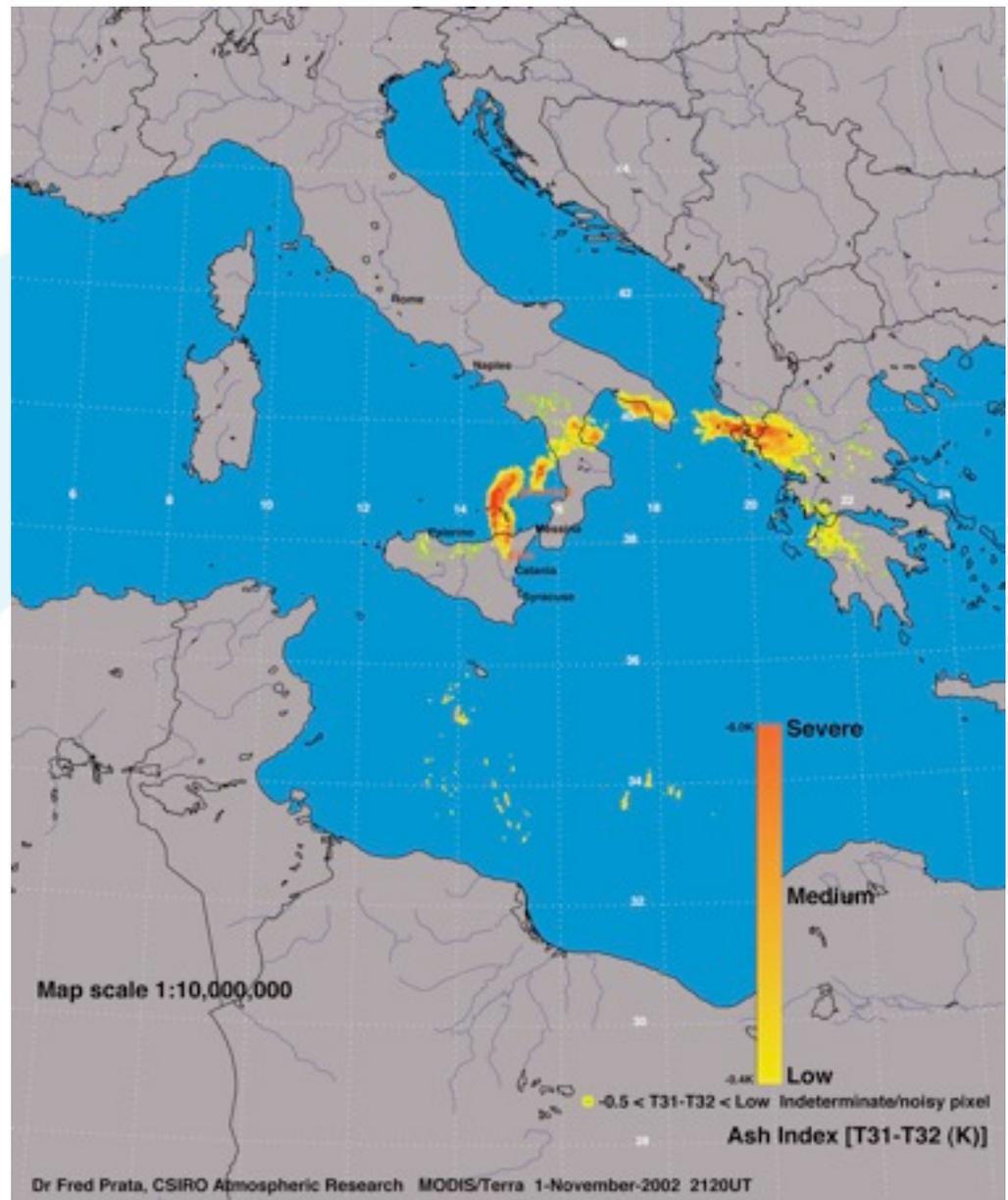
AVHRR 4m5 – Left: NOAA-15 2008/08/08 16:41 UTC  
AVHRR 4m5 – Right: NOAA-16 2008/08/08 17:09 UTC



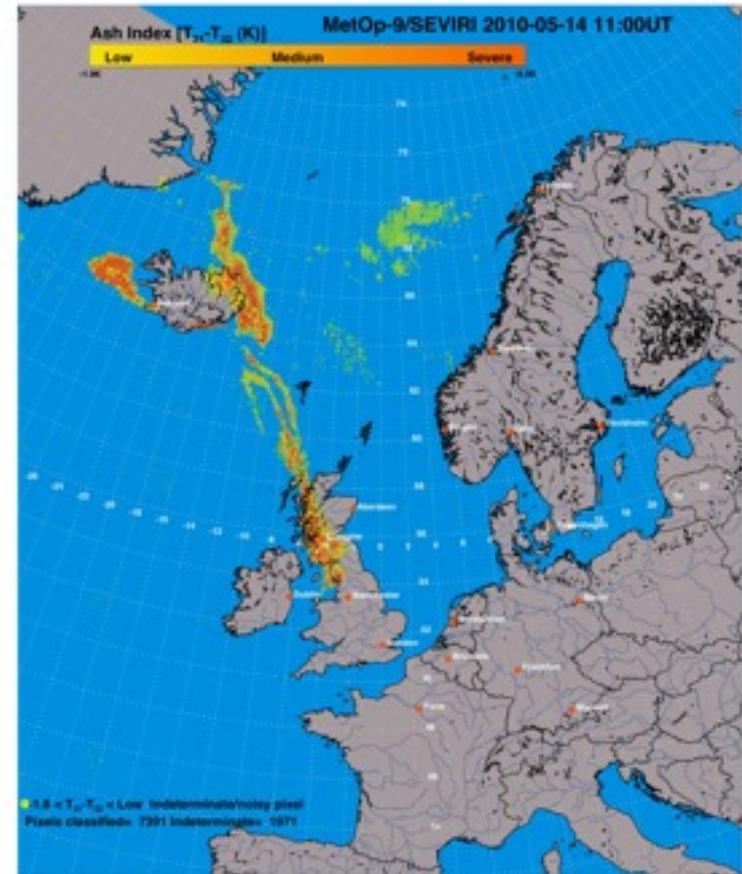
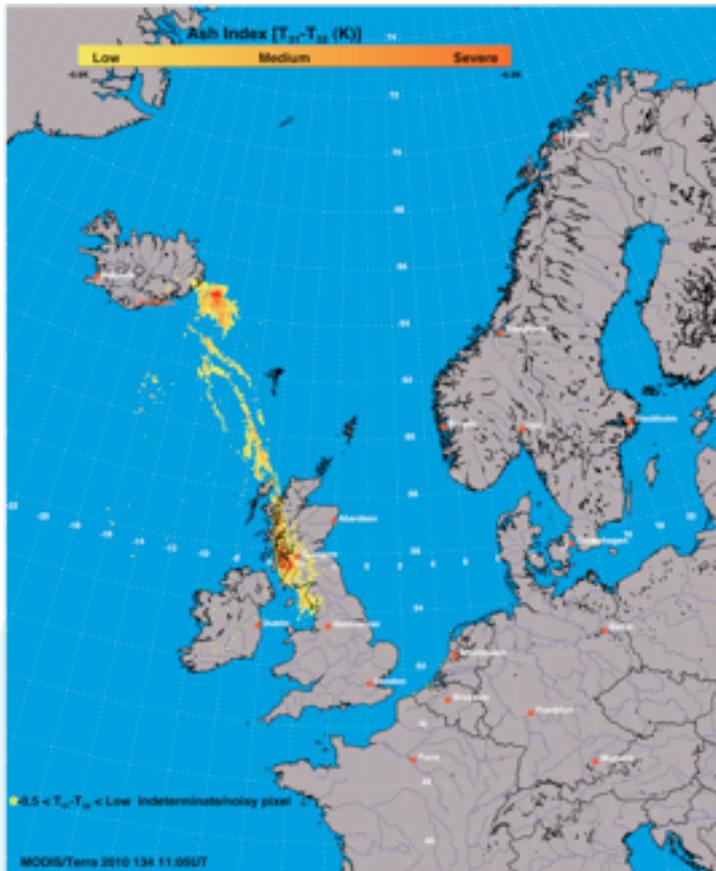
# Quantitative products



# Ash Index



# Ash Index for Eyjafjallajokull



# Summary of Ash Detection Methods

**Table 1** Summary of ash detection algorithms and techniques used with satellite infrared (IR) and visible channel data

Name	Principle	Reference
RA	2-band IR (11 and 12 $\mu\text{m}$ )	Prata (1989a, b)
Ratio	2-band IR (11 and 12 $\mu\text{m}$ )	Holasek and Rose (1991)
4-band	IR + visible	Mosher (2000)
TVAP	3-band IR (3.9, 11 and 12 $\mu\text{m}$ )	Ellrod et al. (2003)
PCI	Multi-band principal components	Hillger and Clark (2002a, b)
WVC	2-band IR + water vapour correction	Yu et al. (2002)
RAT	3-band IR (3.5, 11, 12 $\mu\text{m}$ )	Pergola et al. (2004)
3-band	3-band (IR and visible)	Pavolonis et al. (2006)

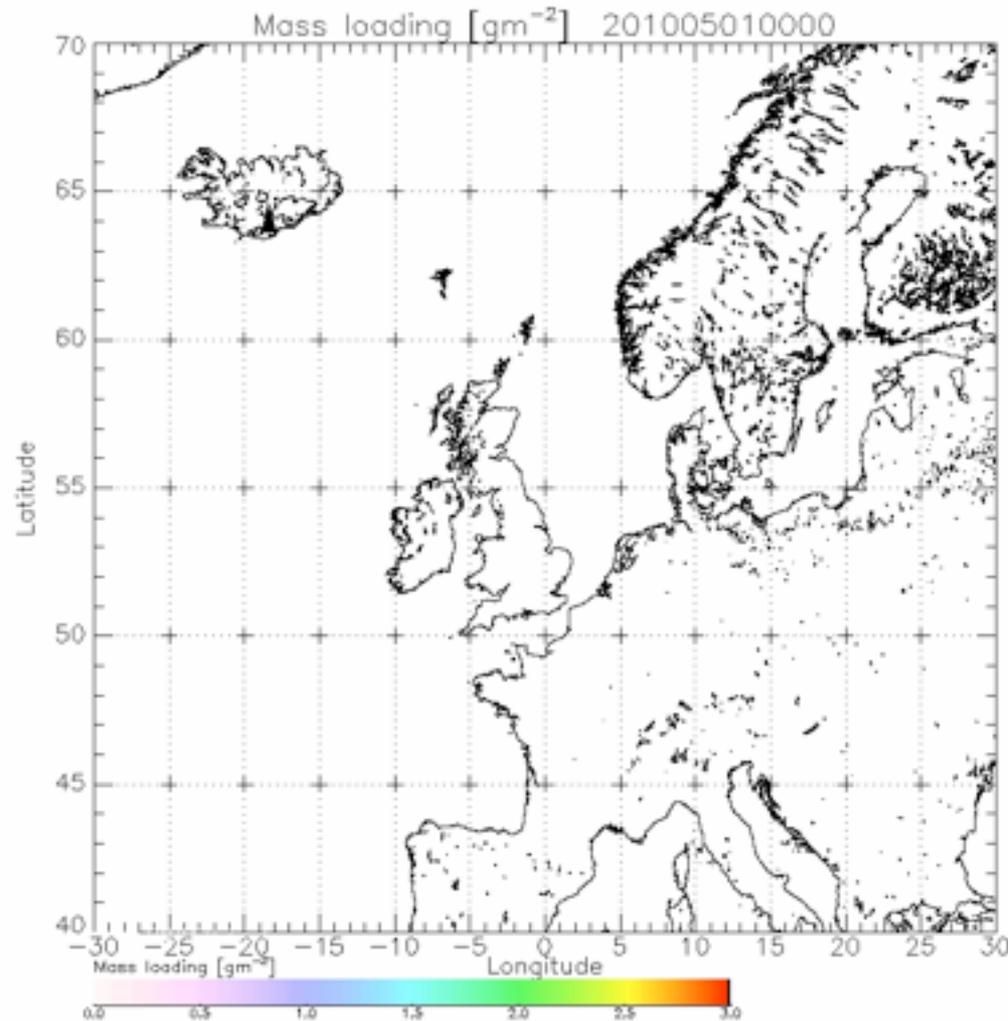
RA = Reverse absorption; TVAP = Three band volcanic ash product; PCI = Principle components; RAT = Ratio method; WVC = Water vapor correction method

Prata, A. J., 2009, Satellite detection of hazardous volcanic clouds and the risk to global air traffic, *Nat. Haz.*, **51**, 303–324.

VAAC	GEO Satellite(s)	Temporal Refresh	Spectral Capabilities	Next Generation GEO Satellite
Anchorage	GOES-11	30 minutes	Split-window	GOES-R (2015)
Buenos Aires	GOES-12 GOES-13 MSG	15 minutes 180 minutes 15 minutes	No split-window No split-window Advanced	GOES-R (2015) and <b>MTG (~2018)</b>
Darwin	MTSAT FY2D FY2E	60 minutes 60 minutes 60 minutes	Split-window Split-window Split-window	GOES-R like from JMA (2020?) and <b>FY4A from China (2014)</b>
London	MSG	15 minutes	Advanced	<b>MTG (~2018)</b>
Montreal	GOES-11 GOES-13	30 minutes 15 or 30 minutes	Split-window No split-window	GOES-R (2015)
Tokyo	MTSAT FY2D FY2E	30 minutes 60 minutes 60 minutes	Split-window Split-window Split-window	GOES-R like from JMA (2020?) and <b>FY4A from China (2014)</b>
Toulouse	MSG	5 or 15 minutes	Advanced	<b>MTG (~2018)</b>
Washington	GOES-11 GOES-12 GOES-13 MSG	30 minutes 15 minutes 15 or 30 minutes 15 minutes	Split-window No split-window No split-window Advanced	GOES-R (2015) and <b>MTG (~2018)</b>
Wellington	MTSAT GOES-11	60 minutes 180 minutes	Split-window Split-window	GOES-R like from JMA (2020?) and GOES-R (2015)

**Table 1: An overview of the geostationary satellite capabilities is shown as a function of Volcanic Ash Advisory Center (VAAC). The table summarizes the temporal and spectral capabilities (those relevant to volcanic ash remote sensing) of each instrument that covers each VAAC area of responsibility. In addition, future geostationary satellite capabilities are summarized. Next generation satellites that include a hyperspectral sounding capability are shown in bold.**

# A closer Look at Ash Detection using SEVIRI

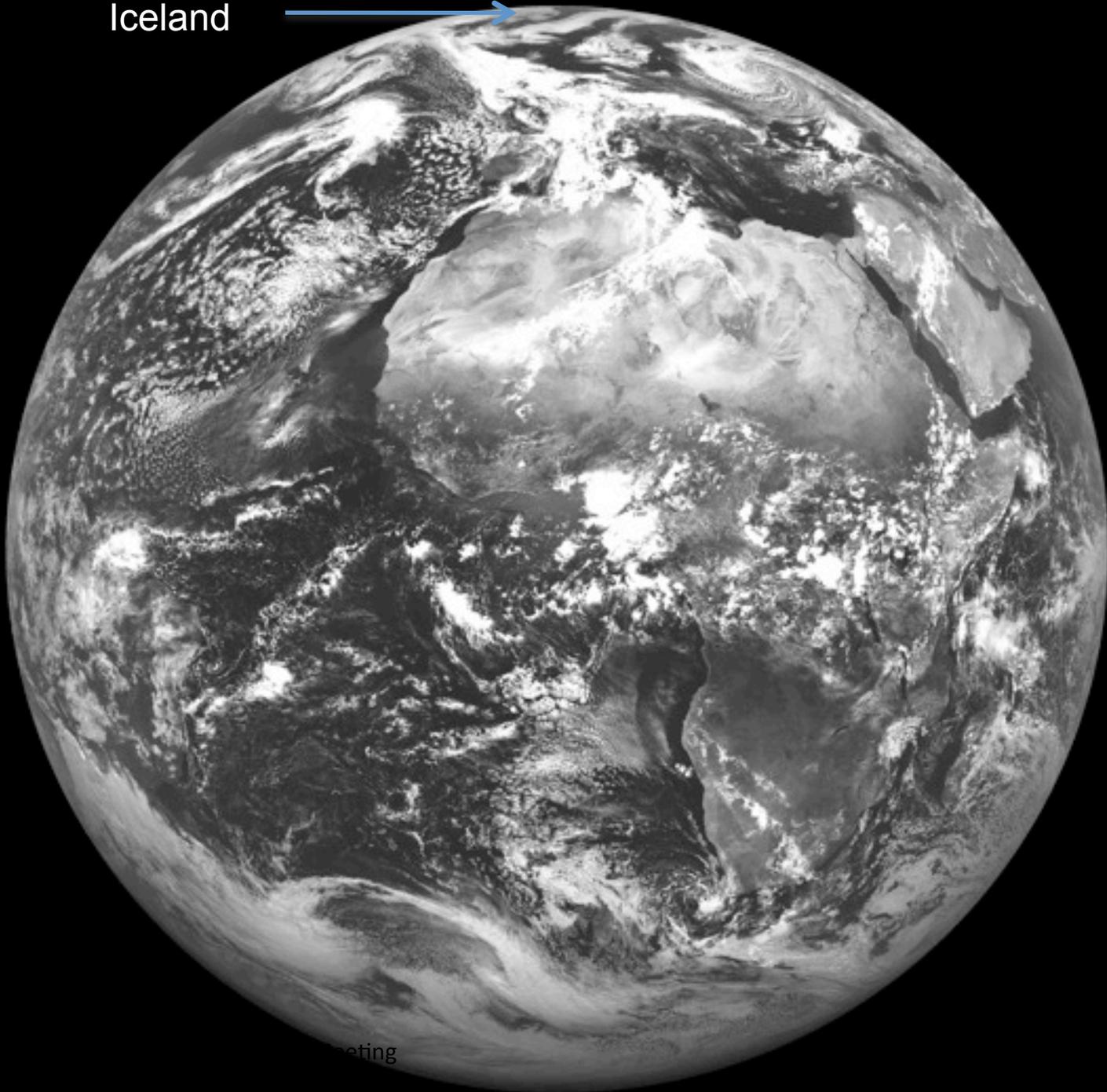


# SEVIRI

## IR channels

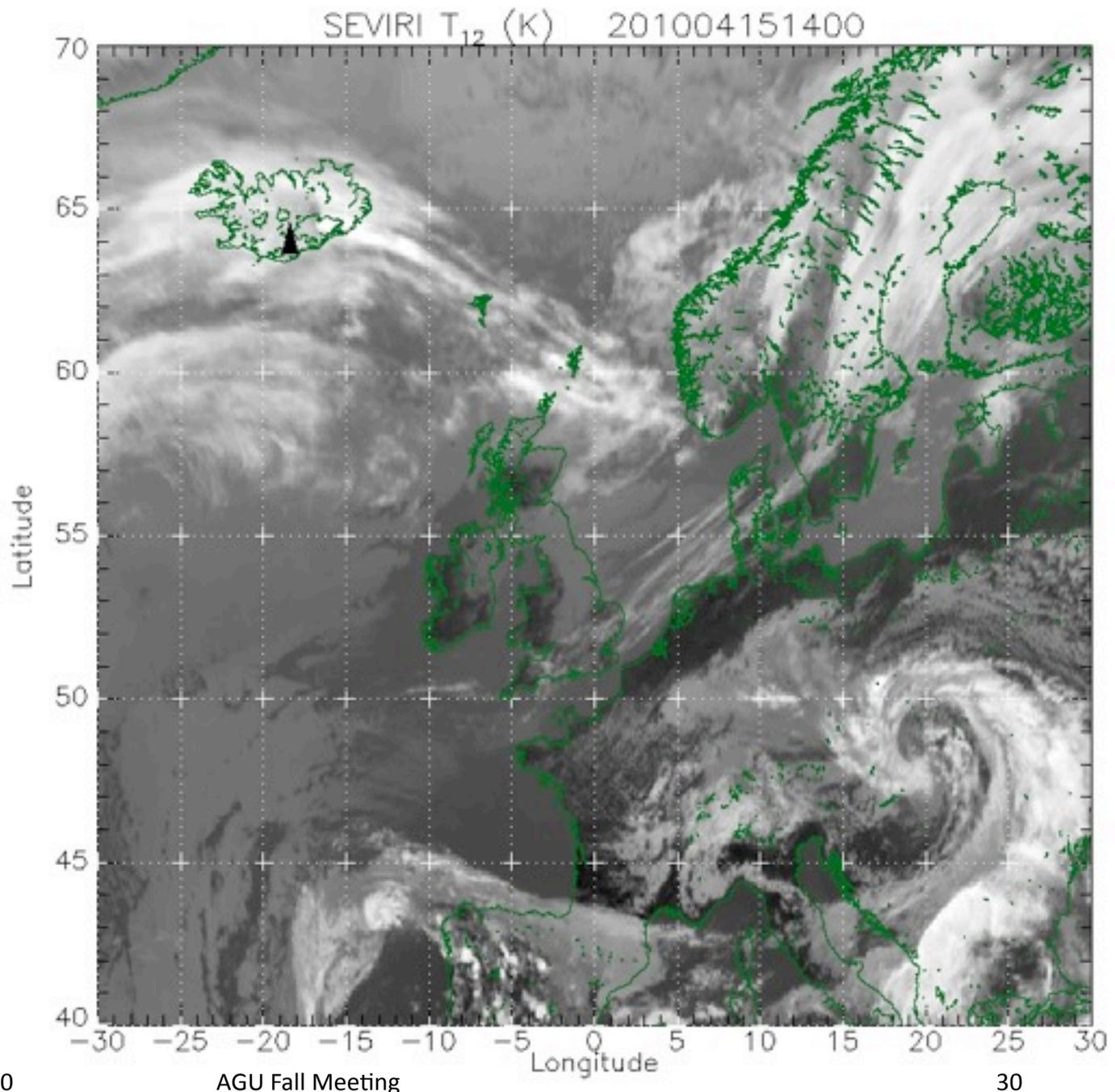
- 6.3  $\mu\text{m}$
- 7.4  $\mu\text{m}$
- 8.6  $\mu\text{m}$
- 9.7  $\mu\text{m}$
- 10.8  $\mu\text{m}$
- 12.0  $\mu\text{m}$
- 13.2  $\mu\text{m}$

Iceland



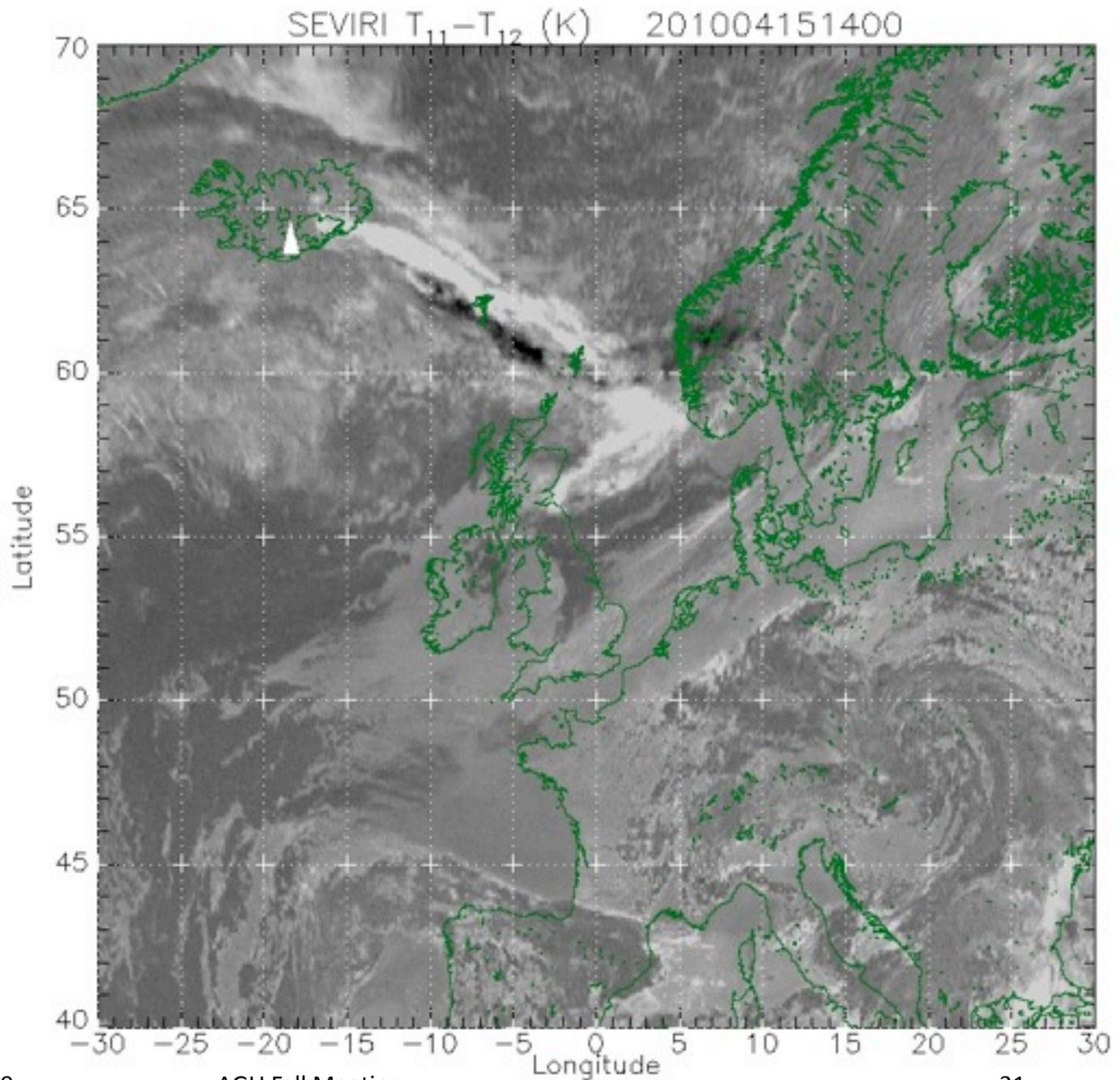
# SEVIRI

## $T_{12}$

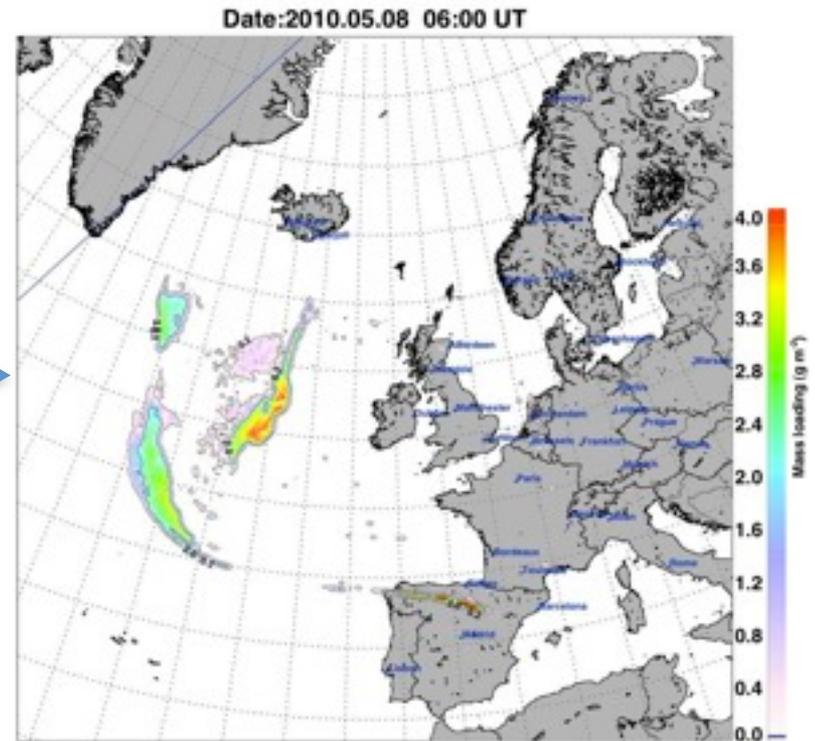
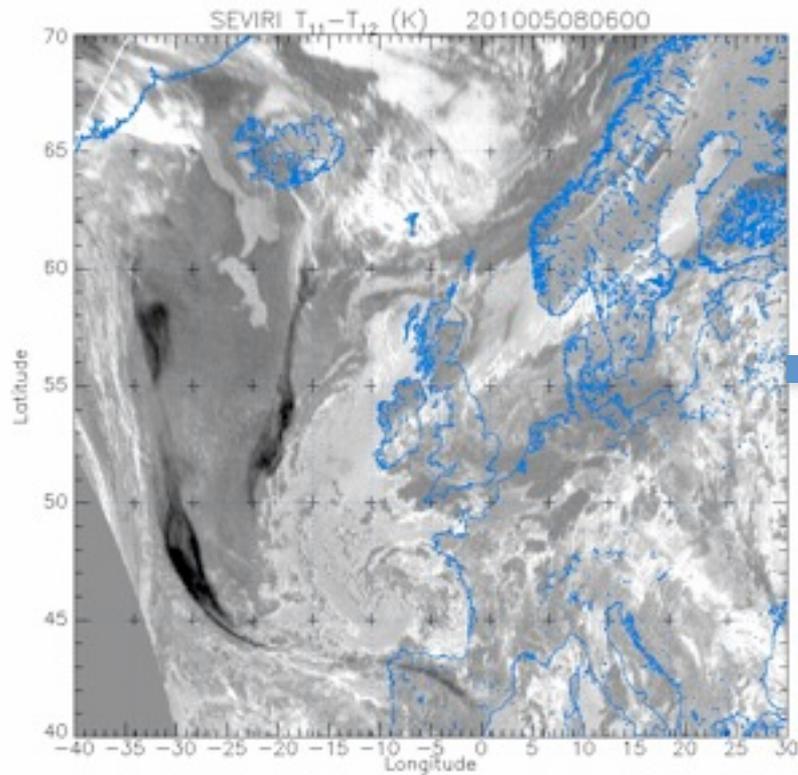


# SEVIRI

## $T_{11} - T_{12}$



# Computer Lab Example



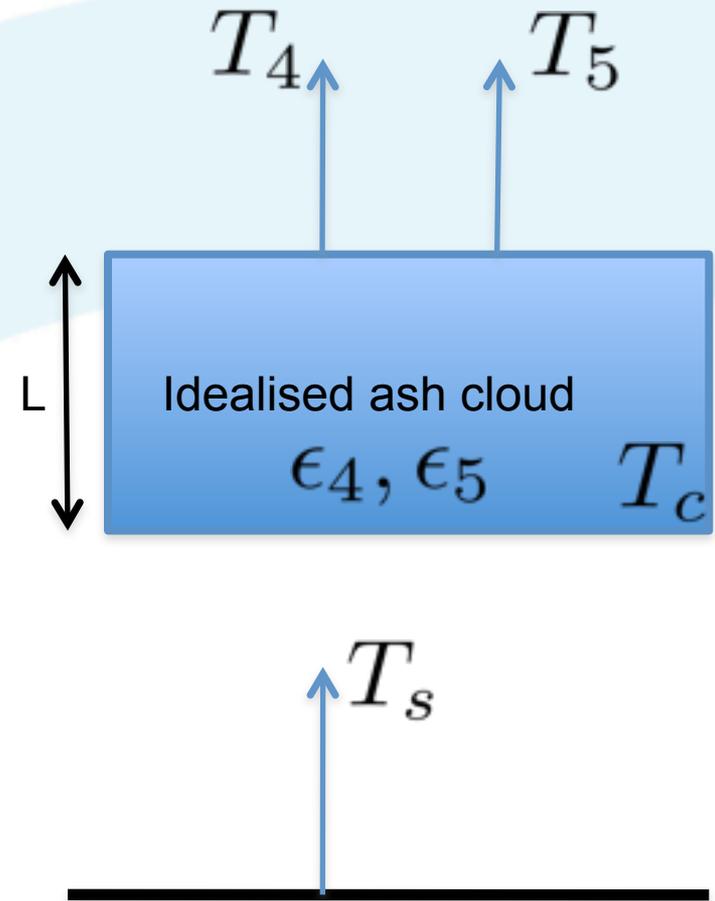
# Parametric equations

$$T_4 = \epsilon_4 T_c + (1 - \epsilon_4) T_s$$

$$T_5 = \epsilon_5 T_c + (1 - \epsilon_5) T_s$$

Solve these for:

$$\Delta T = T_4 - T_5$$



Denoting:

$$\Delta T_4 = T_s - T_4$$

$$\Delta T_{tcon} = T_s - T_c$$

$$\Psi = 1 - \frac{\Delta T_4}{\Delta T_{tcon}}$$

$$\epsilon_i = 1 - \exp(-k_i L)$$



$$\Psi_i = 1 - \epsilon_i$$

“like” cloud optical depth

$$\beta = \frac{k_j}{k_i}$$

$$\lambda_j > \lambda_i$$

Particle microphysics - radius

# Solution

$$\Delta T = \Delta T_{tcon}(\Psi - \Psi^\beta)$$

$\beta > 1$       *ice/water – cloud*

$\beta < 1$       *ash – cloud*

Note that:  $\Psi < 1$

# Solving for radius and optical depth

Given  $T_4$  and  $\Psi$  it is possible to construct “theoretical curves” that suggest the behaviour of “real” measurements on a plot of  $T_4$  vs  $\Delta T$

# Volcanic ash – Detection and discrimination

Prata, A. J., 1989, Infrared radiative transfer calculations for volcanic ash,  
 Geophys. Res. Lett., 16(11), 1293–1296

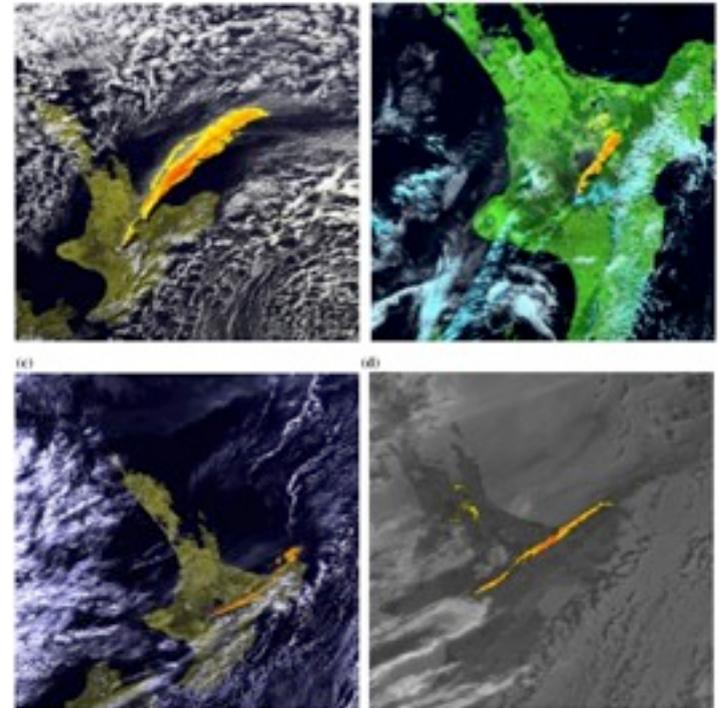
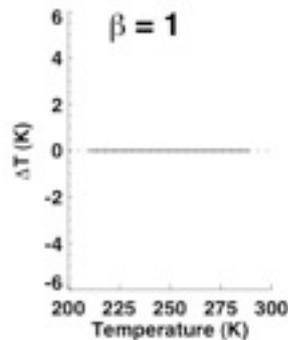
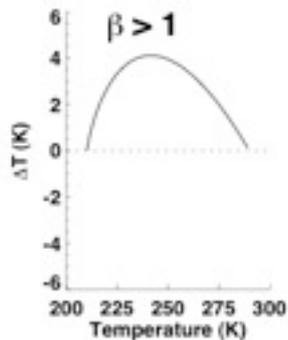
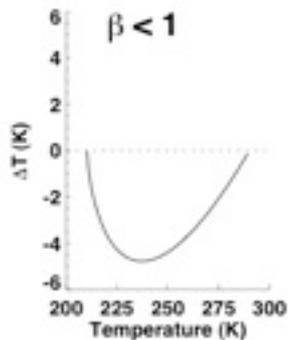
“reverse absorption” algorithm

$$T_{11} - T_{12} < 0$$

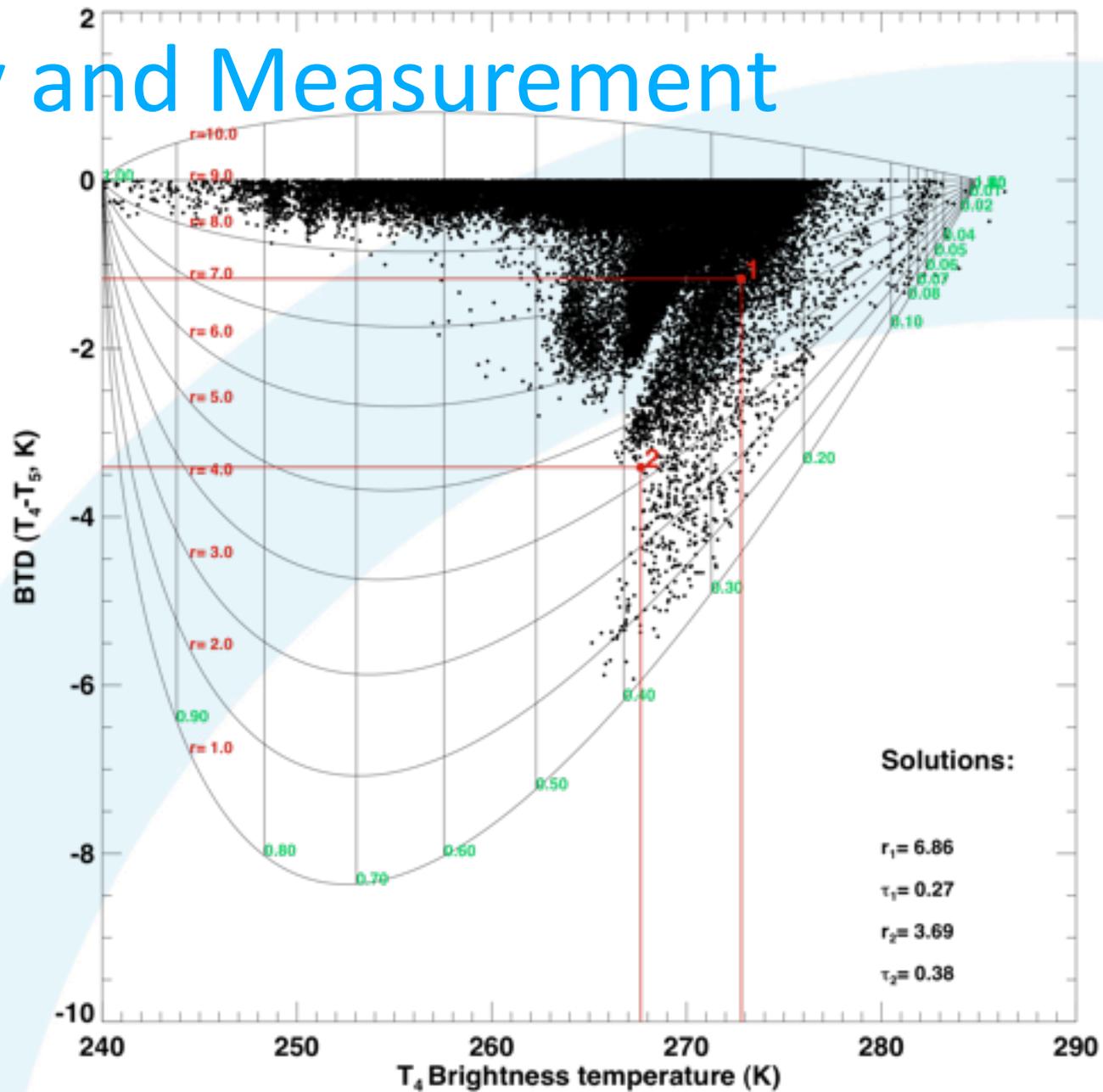
$$\Delta T = \Delta T_c [X - X^\beta],$$

$$X = 1 - \frac{\Delta T_4}{\Delta T_c},$$

$$\Delta T = T_4 - T_5, \quad \Delta T_c = T_s - T_c, \quad \Delta T_4 = T_s - T_4.$$



# Theory and Measurement



# Calculating the mass loading ( $\text{gm}^{-2}$ )

$$M_l = \frac{4}{3} \rho \frac{r\tau}{Q_{ext}}$$

$\rho$

Density

Units:  $\text{kg m}^{-3}$

$Q_{ext}$

Extinction efficiency

none

$\tau$

Optical depth

none

$r$

Effective radius

$\mu\text{m}$

# Eyjafjallajökull, April/May 2010



# Methodology

- Data source: SEVIRI 12-channel 15-minute data for 30W-30E, 40N-70N, from 14.04.2010 to 24.05.2010. Only use IR channels.
- Automatically detect pixels affected by ash
- Apply water vapour correction
- Apply parallax correction
- Retrieve effective particle radius, IR optical depth and mass loading ( $\text{g m}^{-2}$ )
- Validate
- Determine concentrations using coincident CALIOP space-based lidar measurements

# Some problems with ash detection

Water vapour makes  $T_{11}-T_{12} > 0$

*Correction applied based on Yu et al. (2003)*

Clear land at night gives  $T_{11}-T_{12} < 0$

*No good correction algorithm available*

Mixed pixels

*Difficult to correct for without sub-pixel data*

Improvement possible for some pixels by using spatial information

# Retrieval scheme

Solve the RT equation with scattering for a plane-parallel cloud:

$$\mu \frac{\partial I}{\partial \tau}(\tau, \mu) = I(\tau, \mu) - (1 - \varpi_o) B(T) \frac{\varpi_o}{2} \int_{-1}^1 P(\mu; \mu') I(\tau, \mu') d\mu'$$

Boundary conditions:

$$I(0, -\mu) = 0$$

$$I(\tau_1, +\mu) = B(T_s)$$

Solution:

$$I(\tau, \mu_i) = \sum_j L_j W_j(\mu_i) \exp(-k_j \tau) + B(T)$$

# Microphysics

Refractive index as a function of wavelength.  
Spherical particles.  
Mie scattering equation solver

Modified-gamma and log-normal size distributions

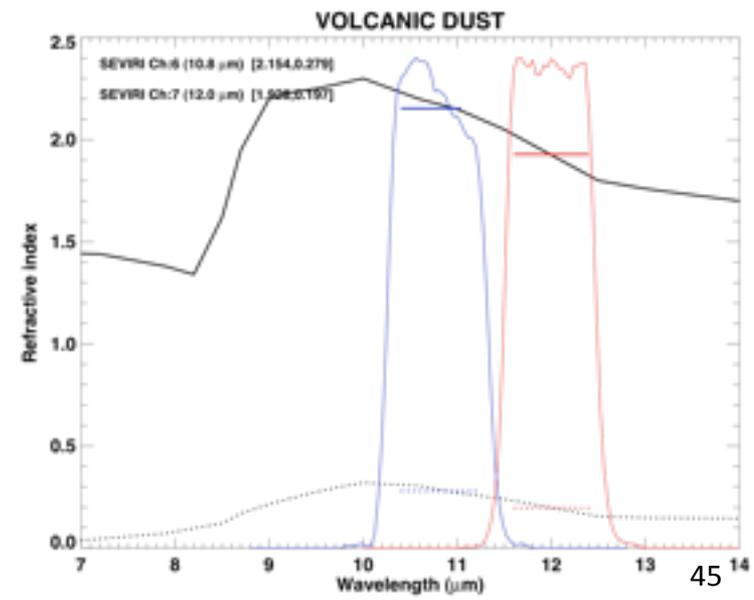
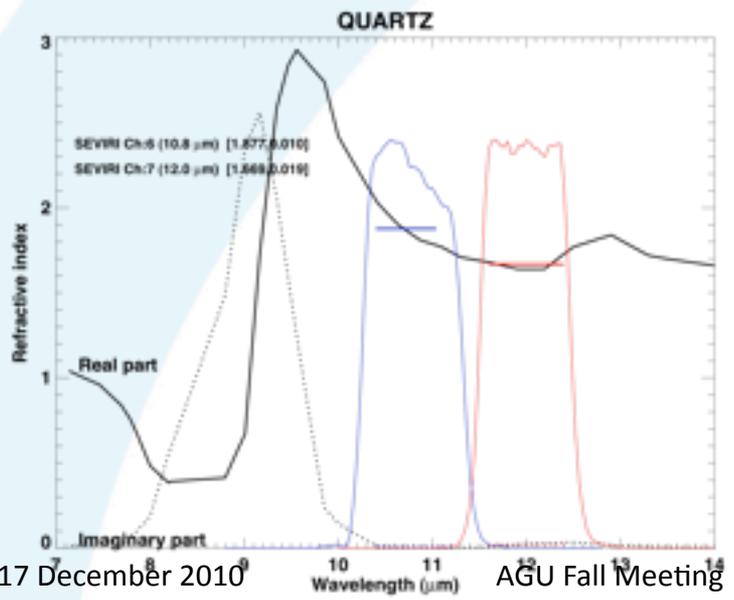
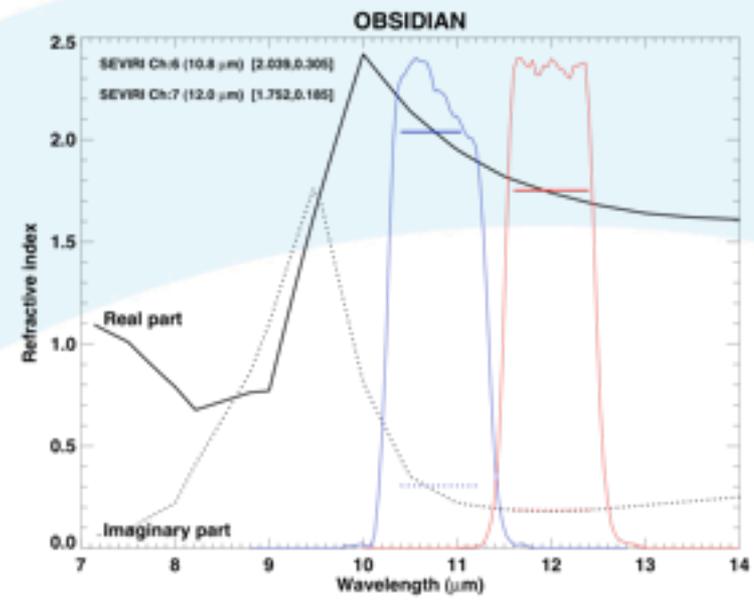
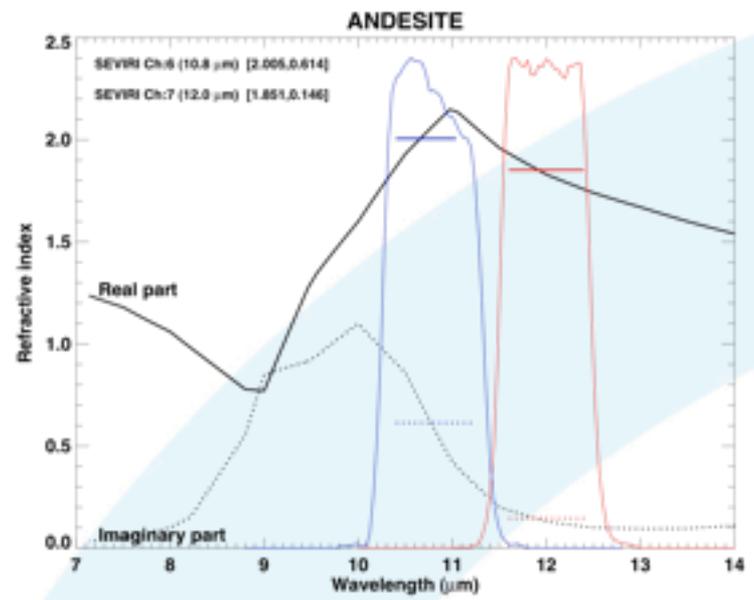
$$\frac{dn(r)}{dr} = \frac{Nb^7}{6!} r^6 \exp(-br)$$

$$\frac{dn(r)}{dr} = \frac{N}{r\sigma\sqrt{2\pi}} \exp\left\{-\frac{(\ln[r] - \ln[r_o])^2}{2\sigma^2}\right\}$$

Efficiency factors for polydispersions:

$$\hat{Q}_f = \frac{\int_0^\infty \pi r^2 Q_f\left(\frac{2\pi r}{\lambda}, m\right) \frac{dn(r)}{dr} dr}{\int_0^\infty \pi r^2 \frac{dn(r)}{dr} dr}$$

# Refractive indices



# Procedure

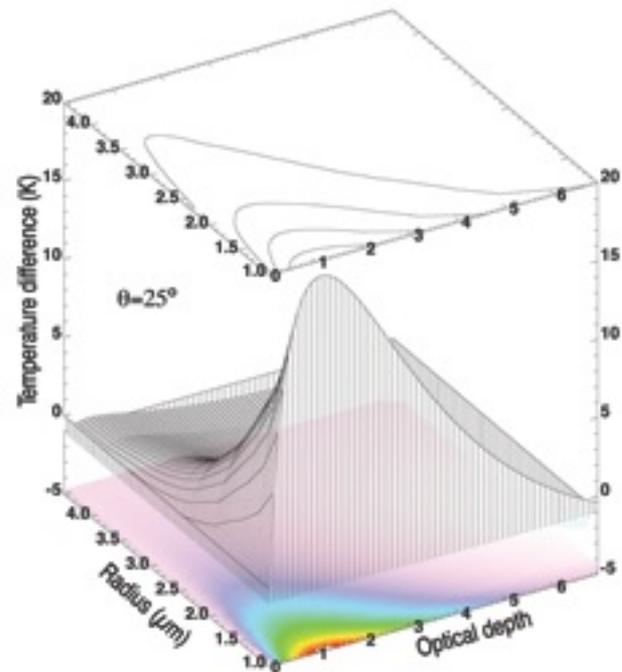
Run Mie code for particle sizes from 0.2 to 32  $\mu\text{m}$  in steps of 0.2  $\mu\text{m}$

Run RT code for all particle sizes, optical depths from 0.02 to 8, in steps of 0.02, 16 streams, and cloud-top and surface temperatures from  $T_c \pm 10$  K,  $T_s \pm 10$  K

Generate very large look-up table containing simulated top-of-atmosphere (no water vapour) brightness temperatures at 10.8, 12.0  $\mu\text{m}$

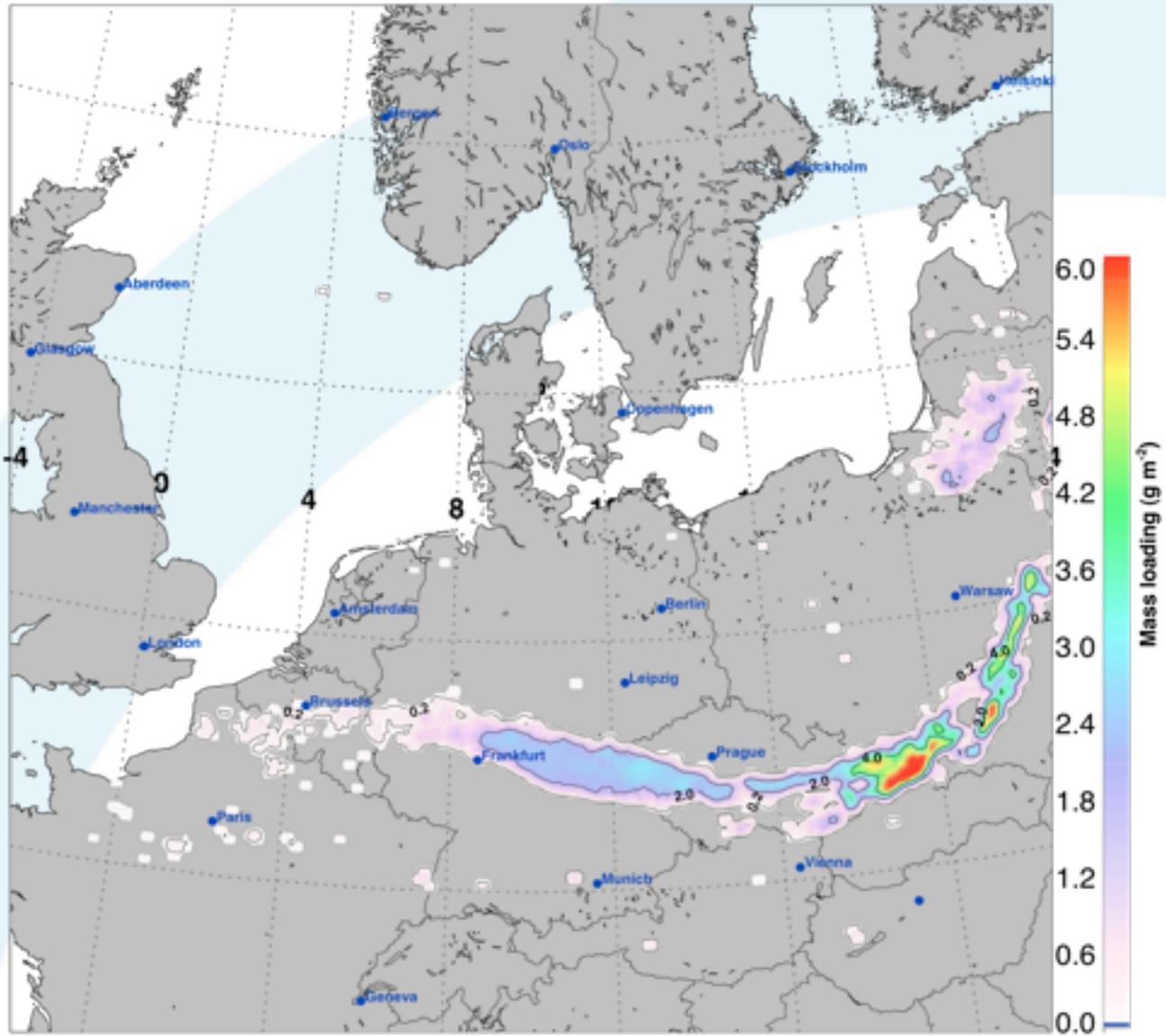
Use ash-identified pixels with measured BTs at 10.8 and 12.0  $\mu\text{m}$  to find “best fit” to simulations.

(Sometimes several values found).

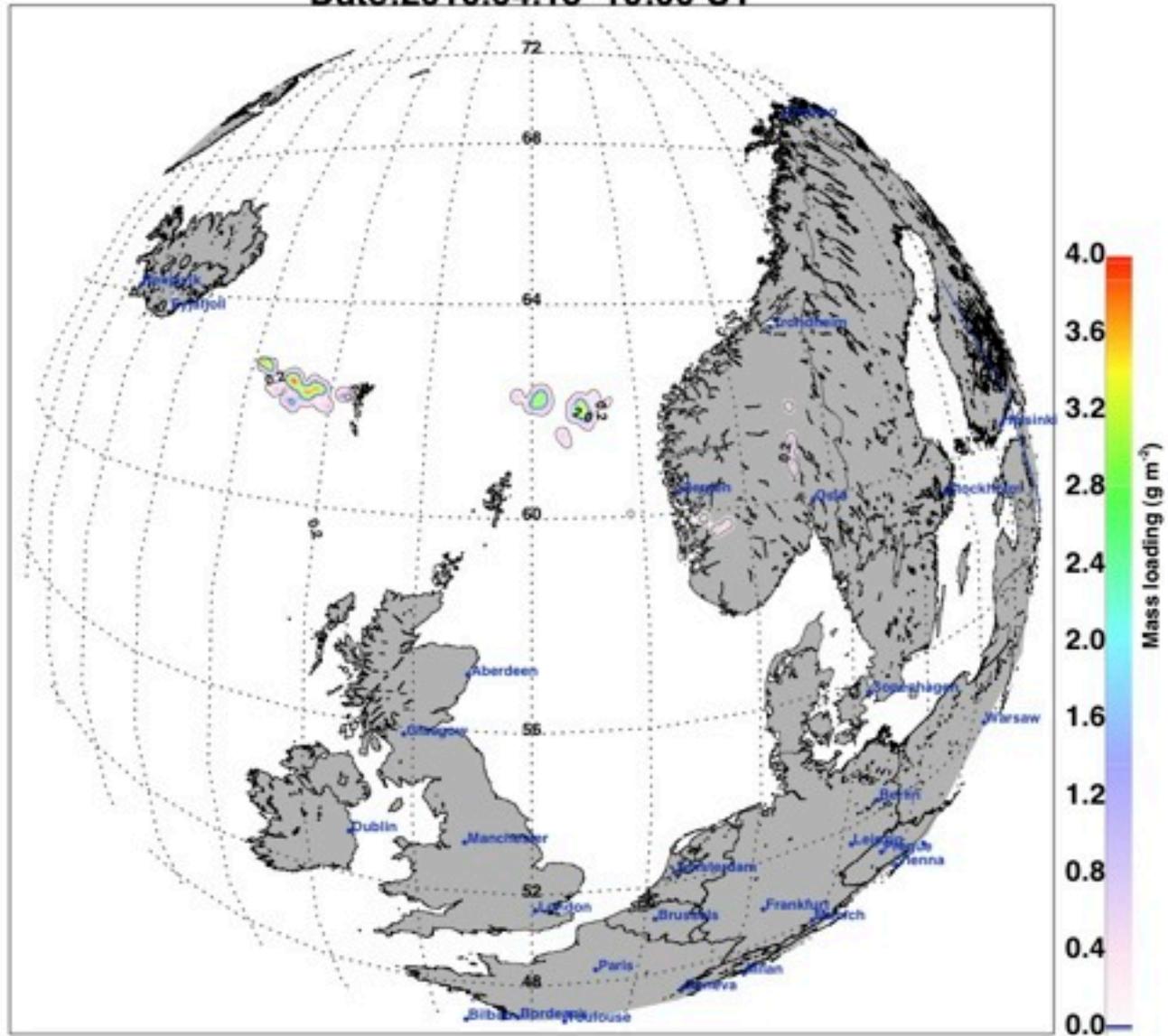


# 16 April 20:00 UT

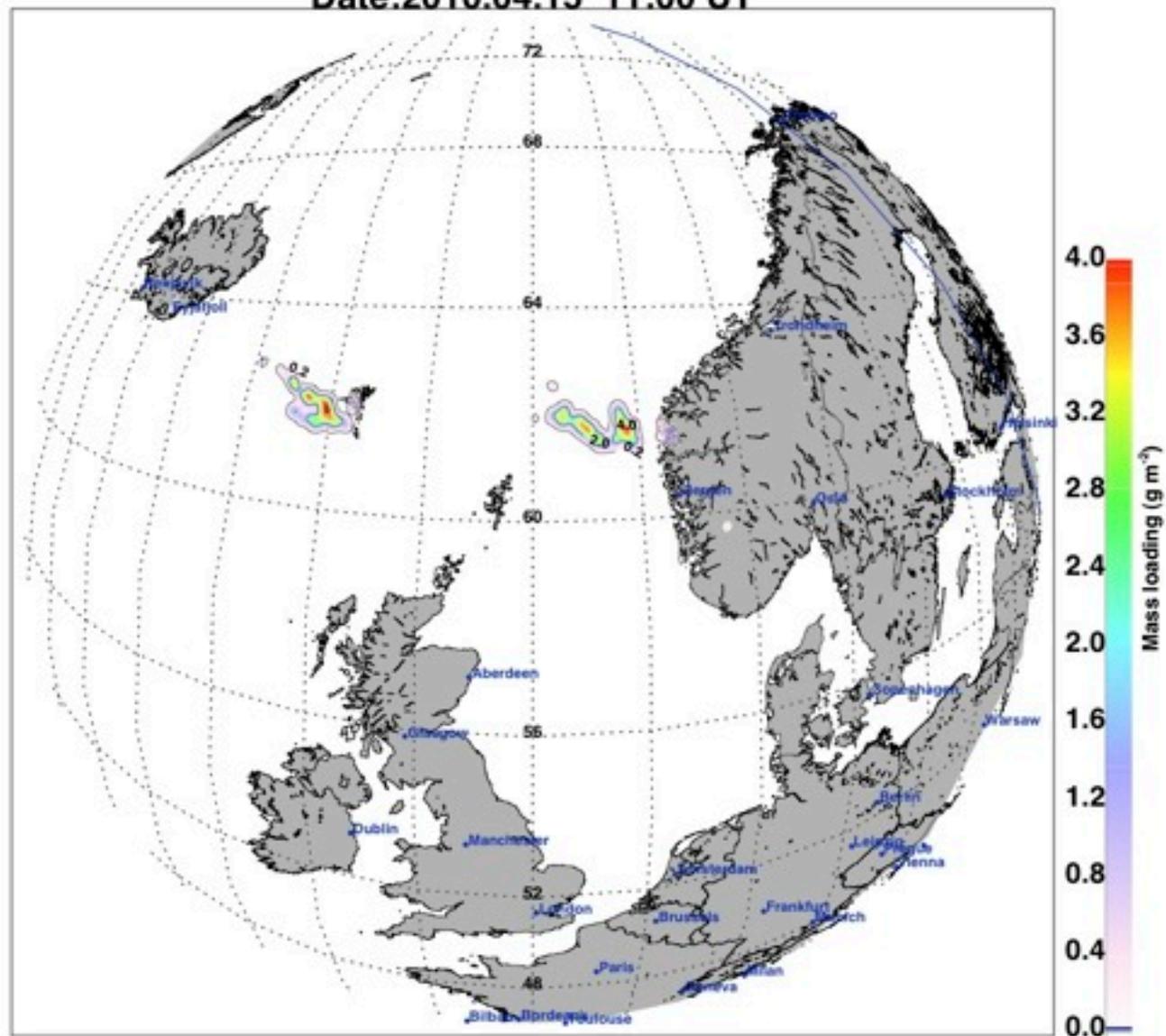
Date:2010.04.16 20:00 UT



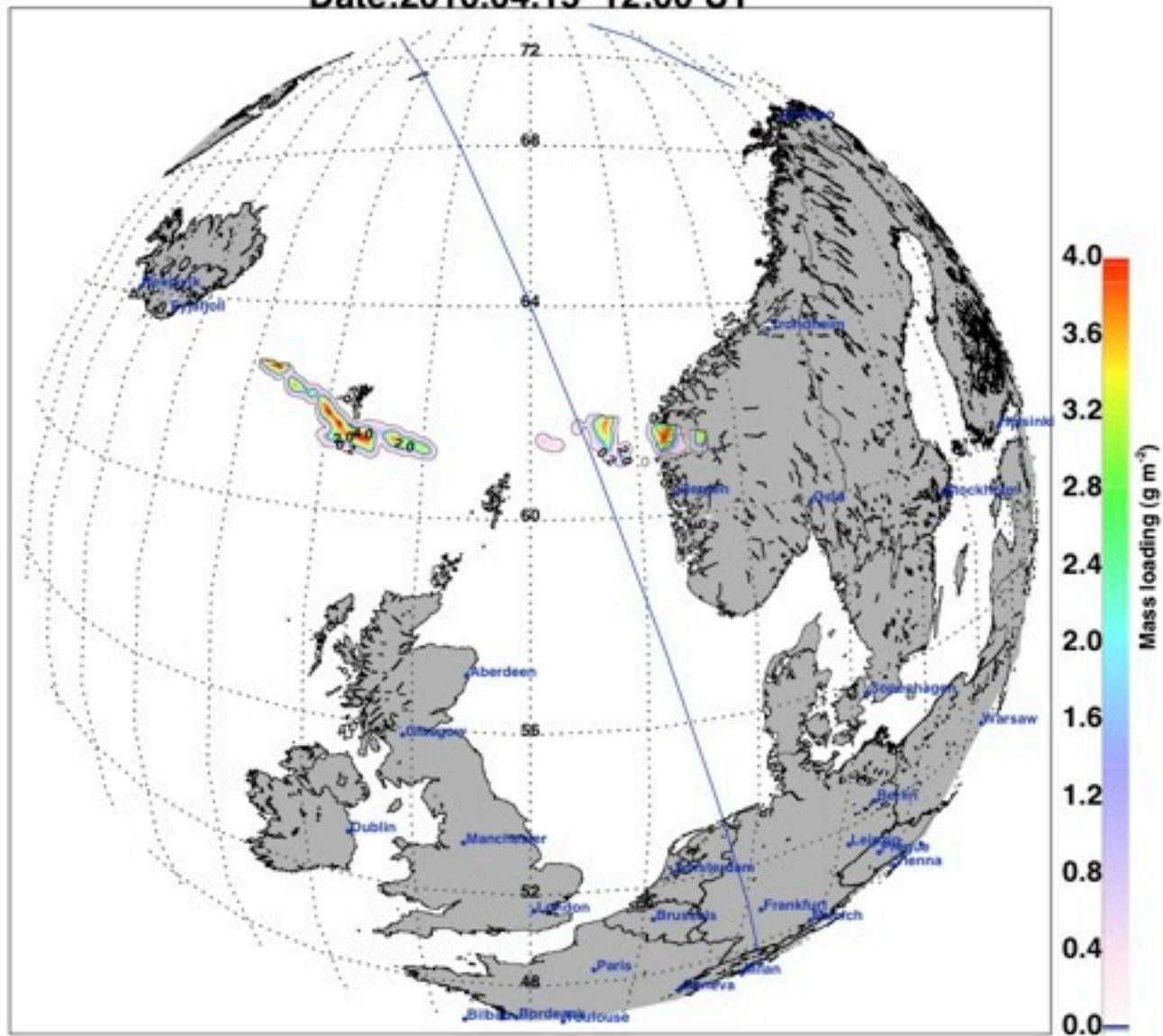
Date:2010.04.15 10:00 UT



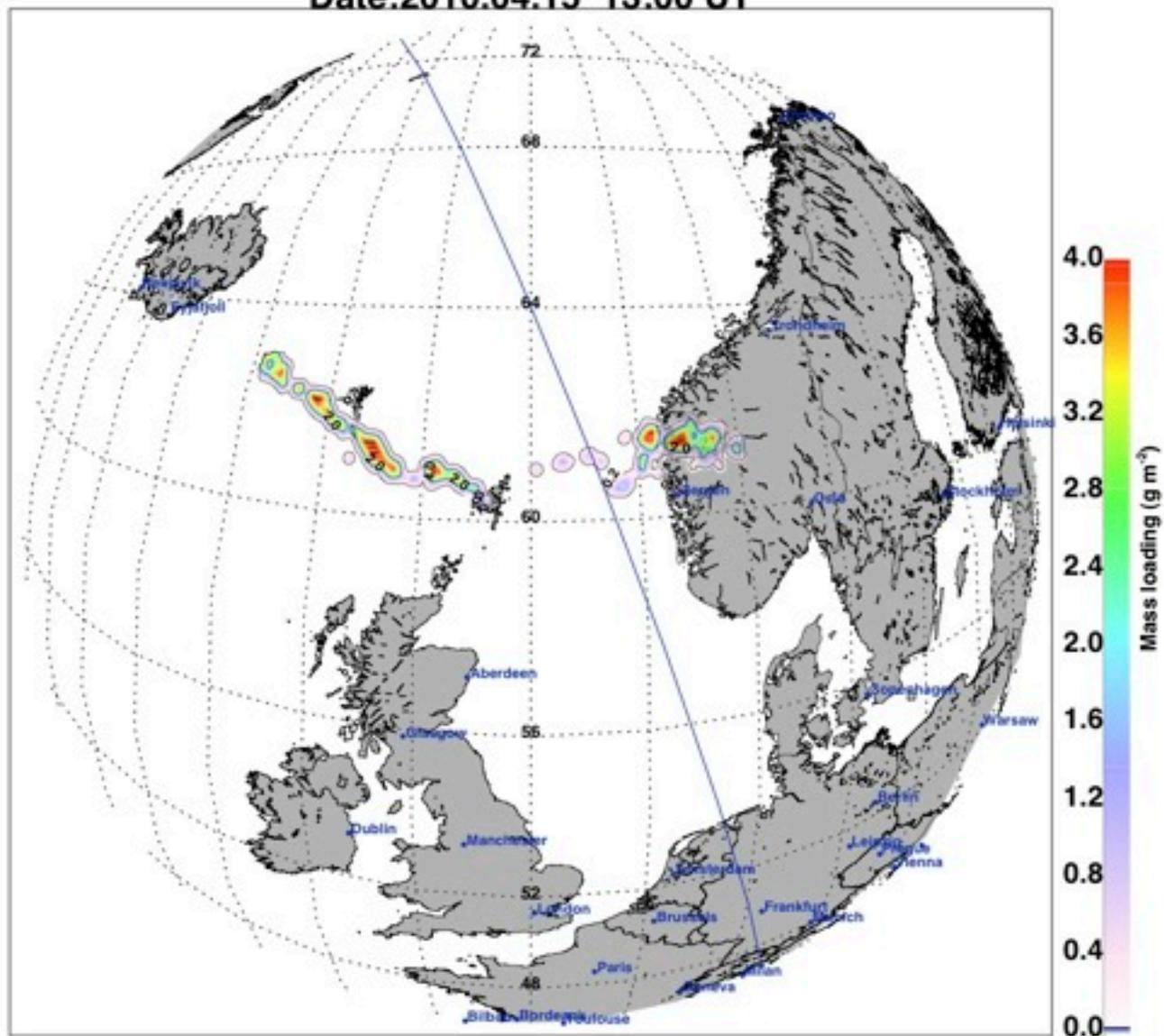
Date:2010.04.15 11:00 UT



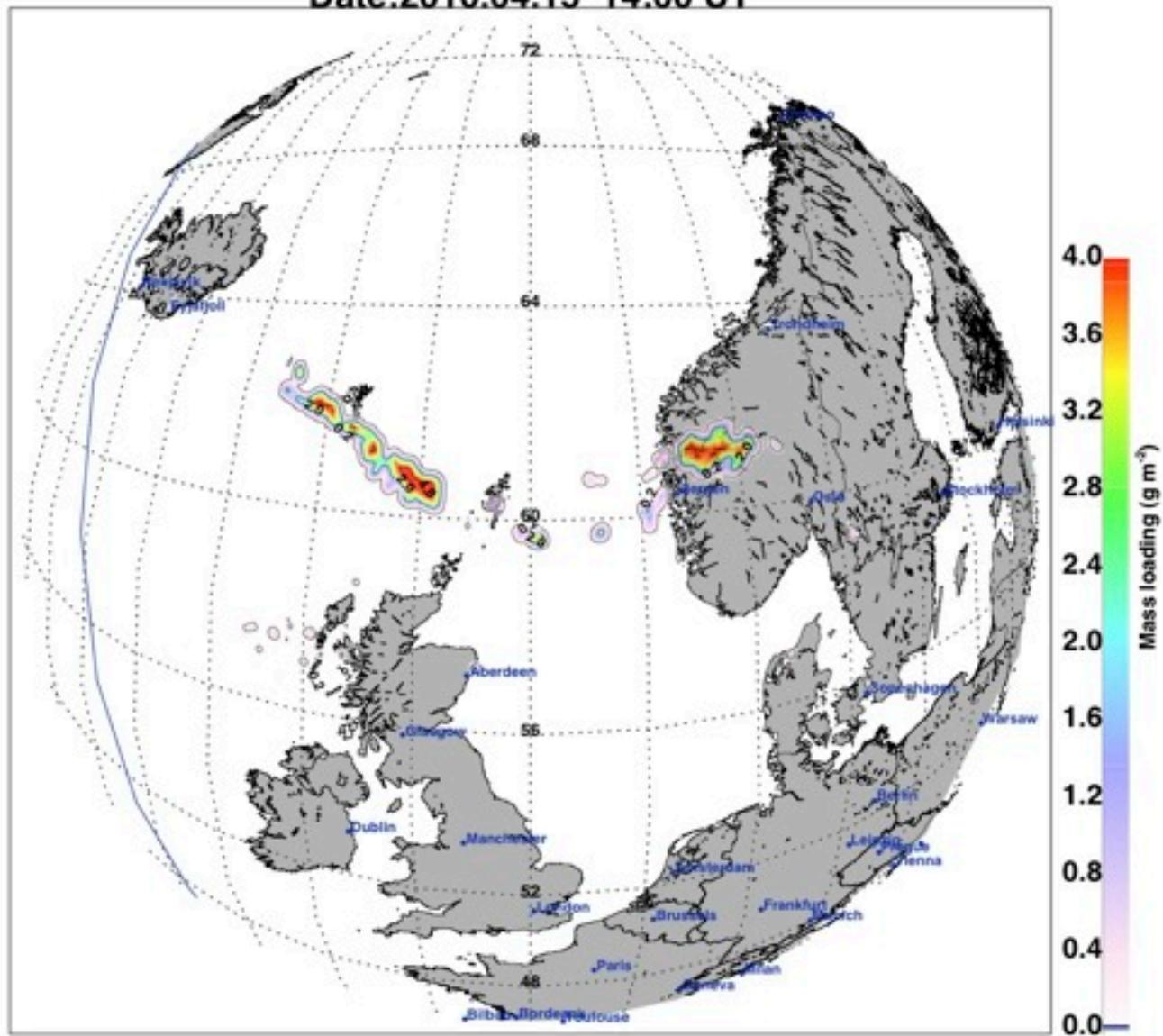
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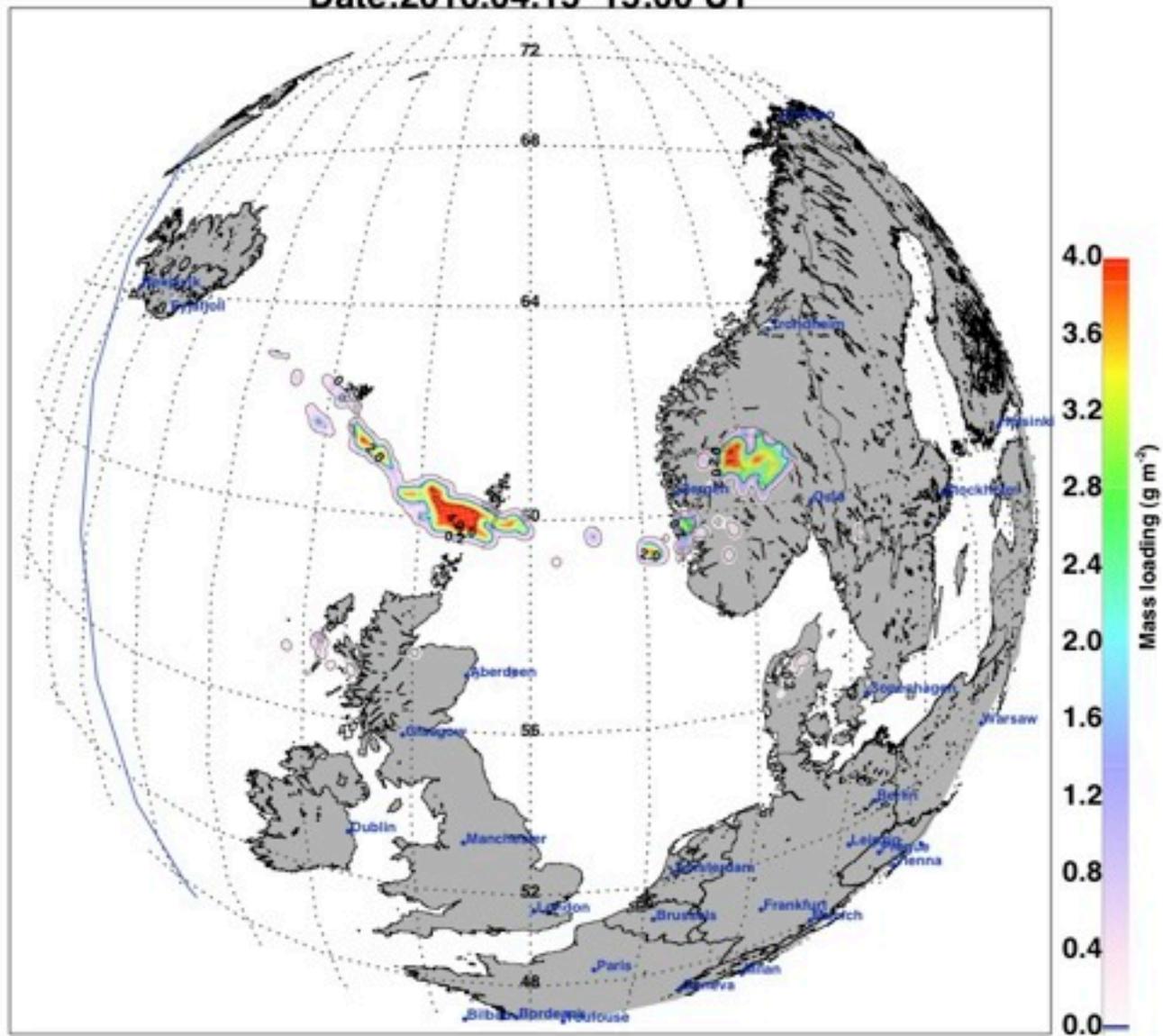
Date:2010.04.15 13:00 UT



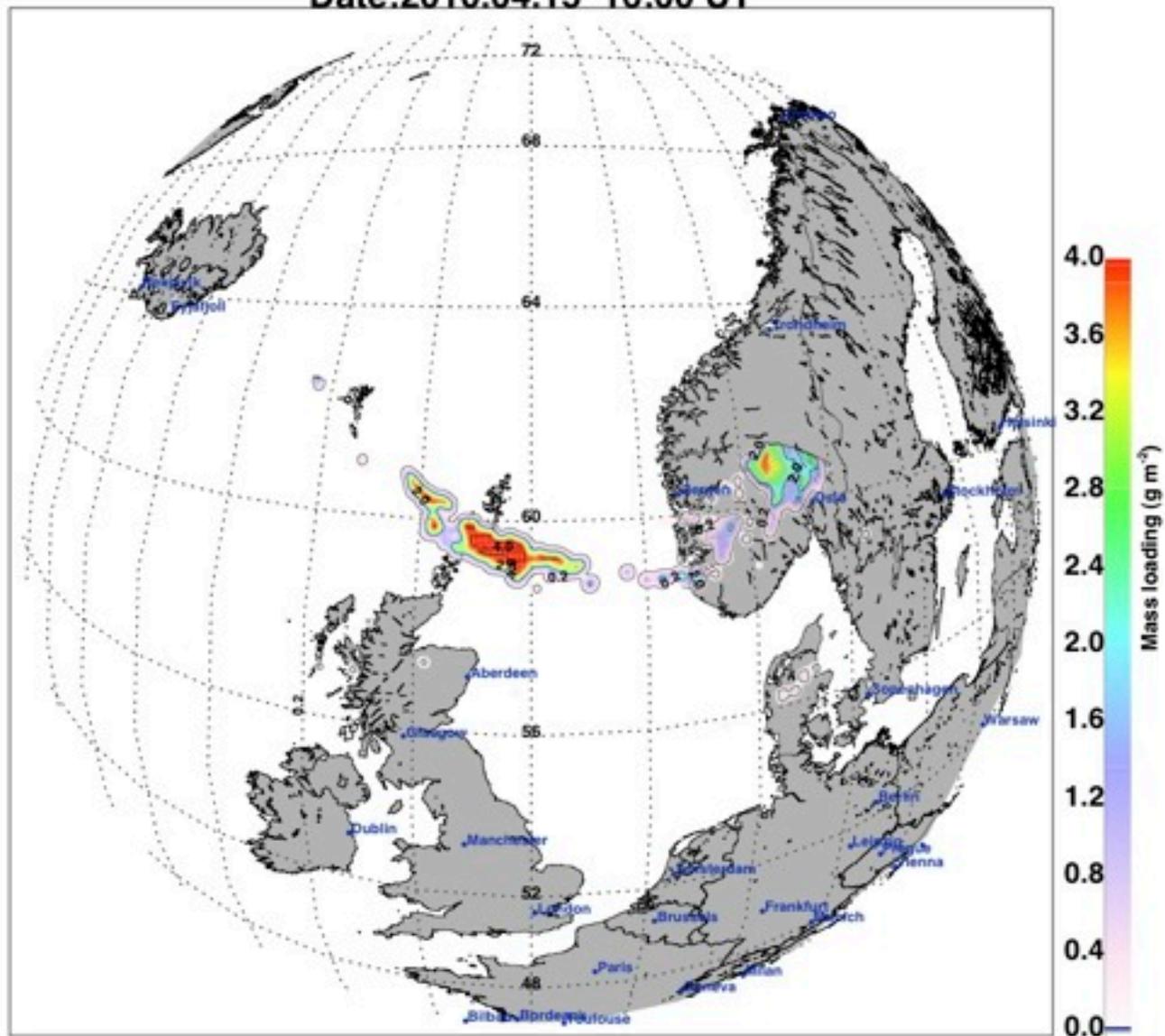
Date:2010.04.15 14:00 UT



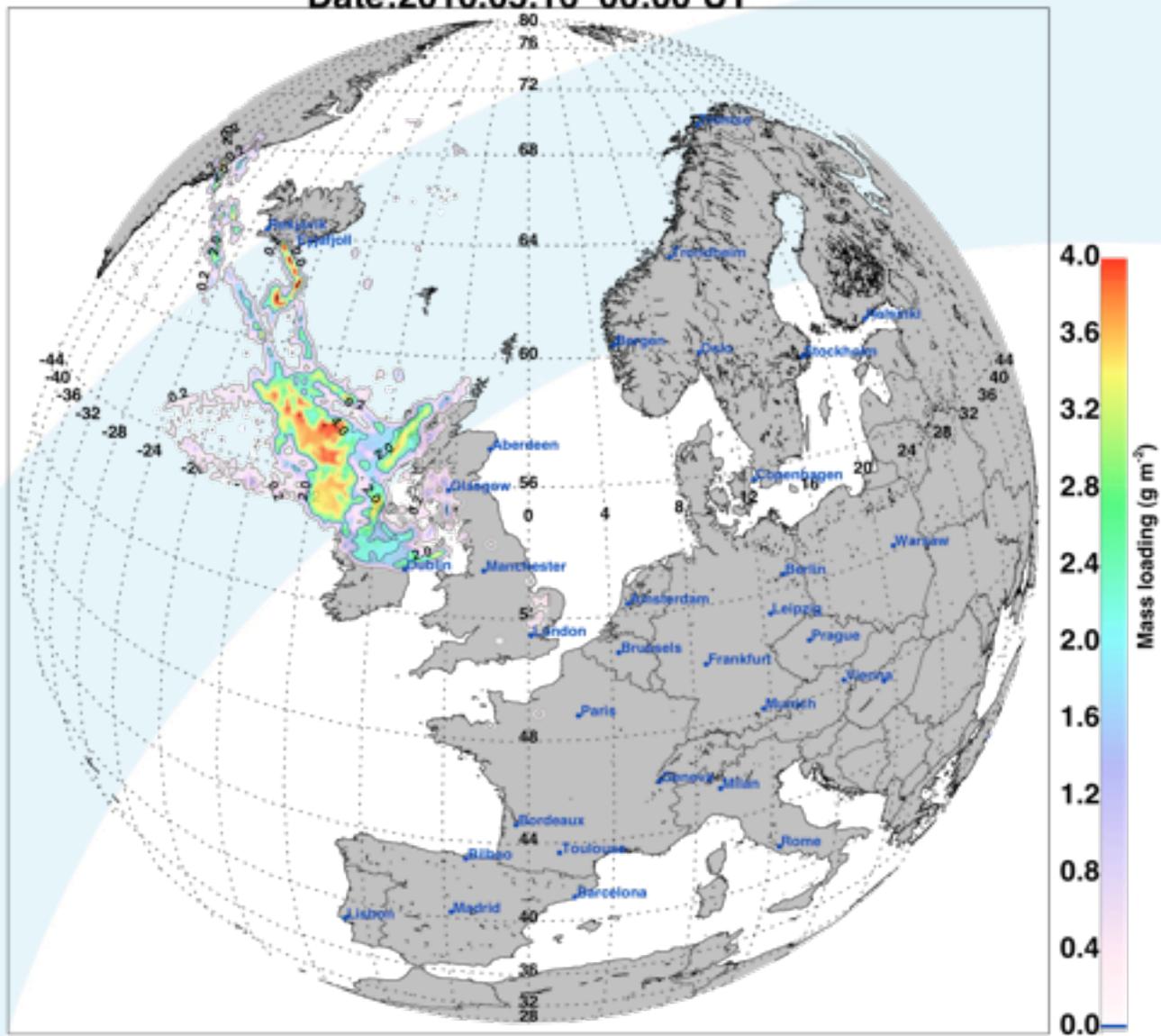
Date:2010.04.15 15:00 UT



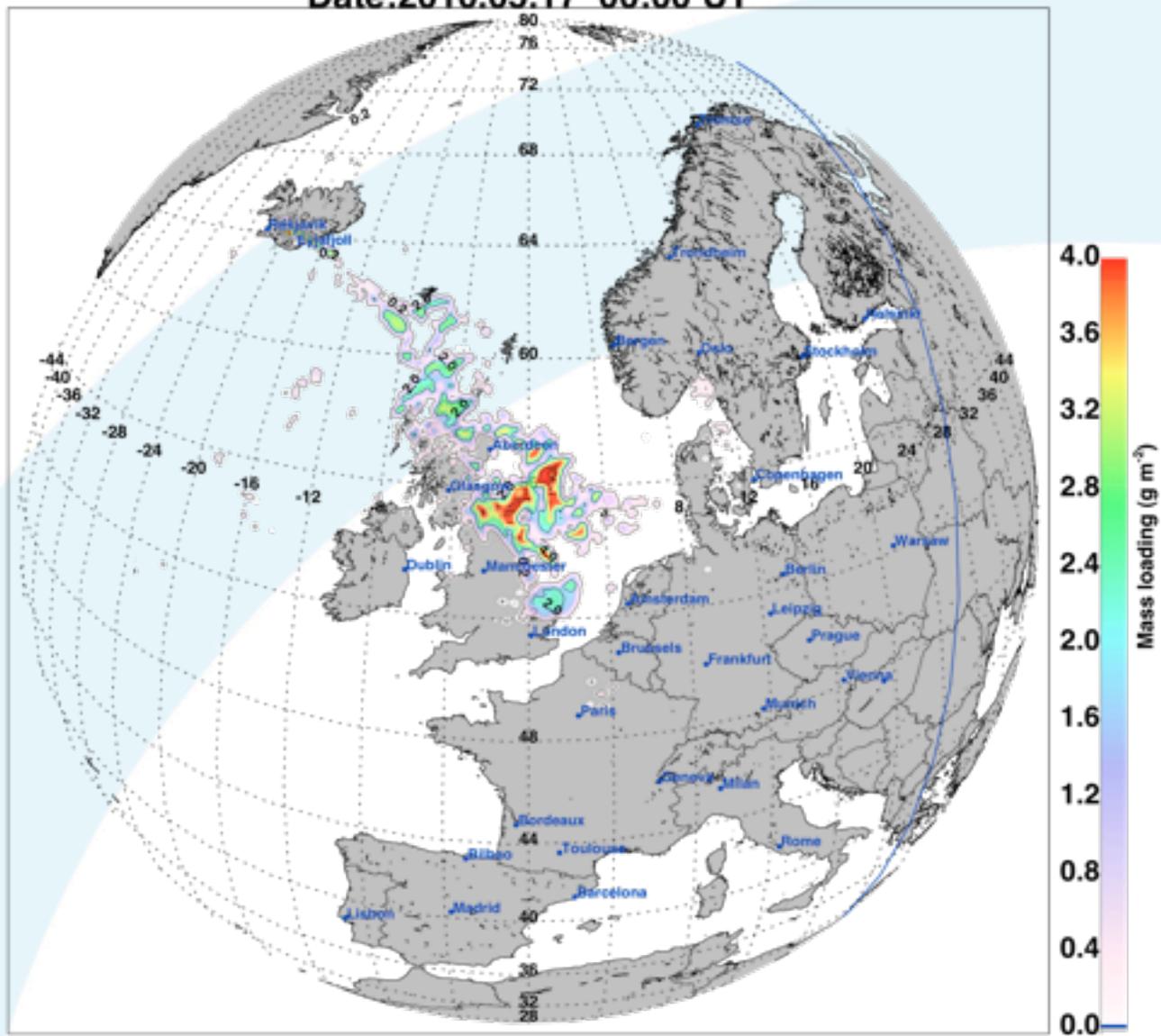
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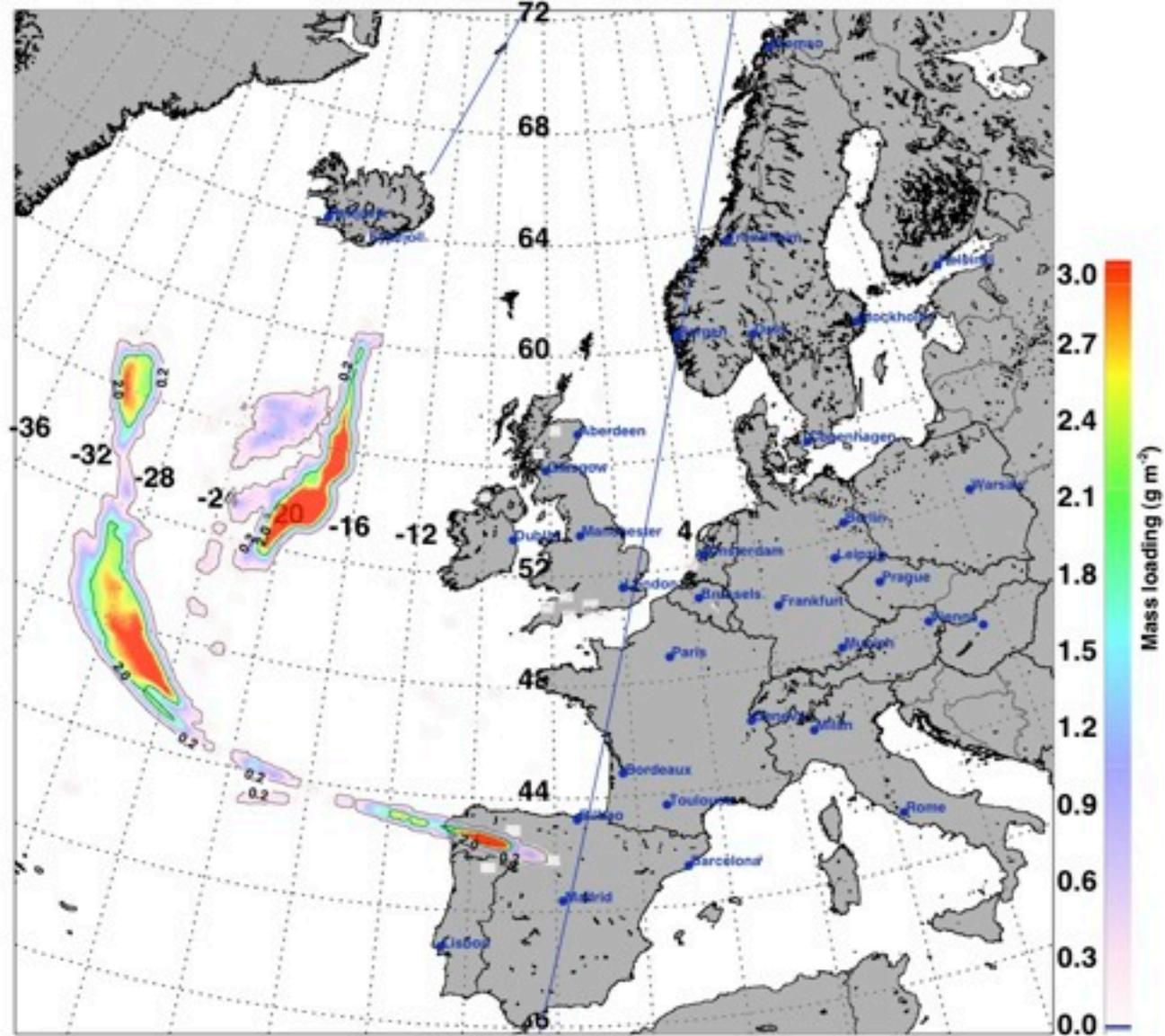
Date:2010.05.16 00:00 UT



Date:2010.05.17 00:00 UT

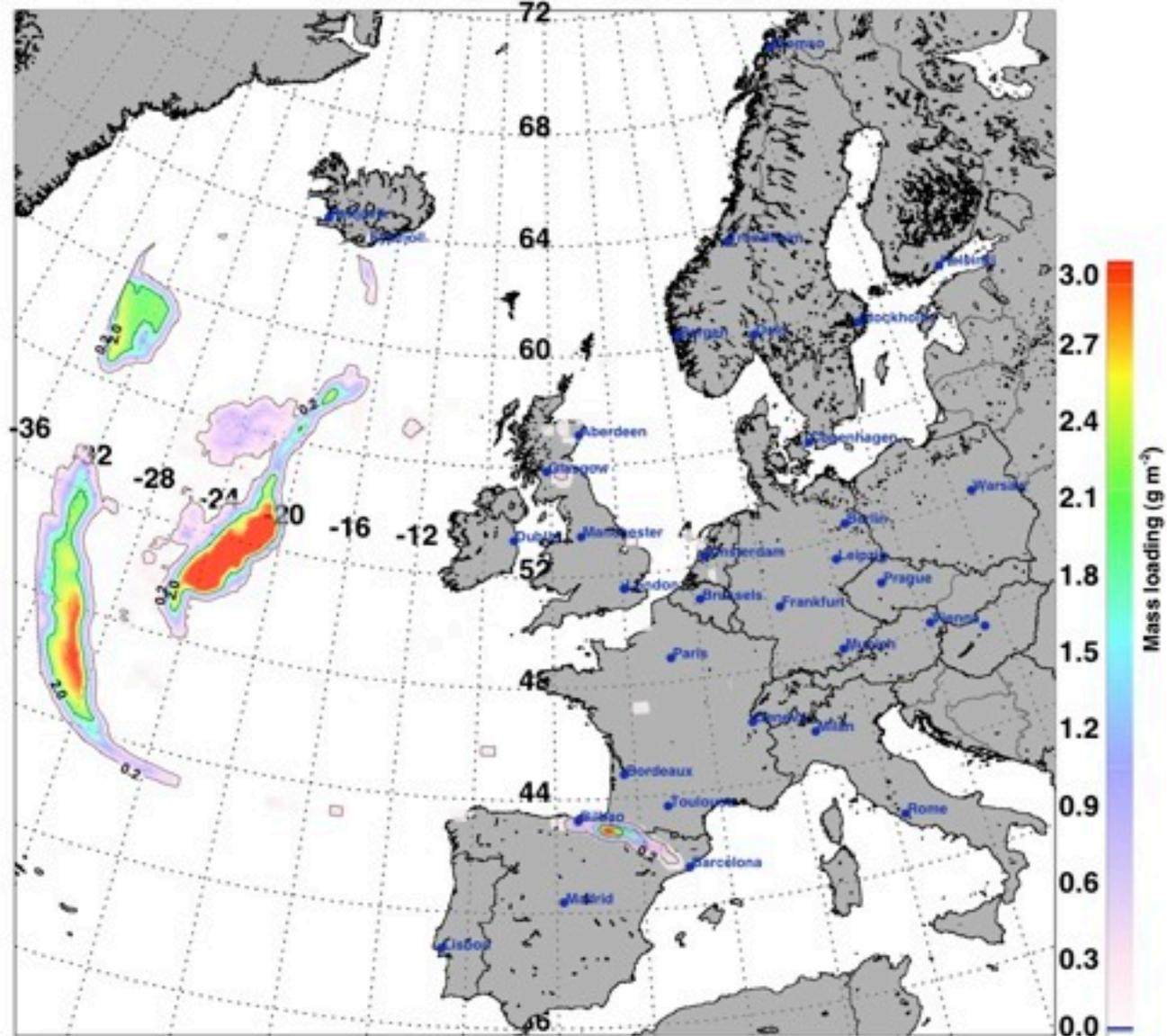


Date:2010.05.08 03:00 UT

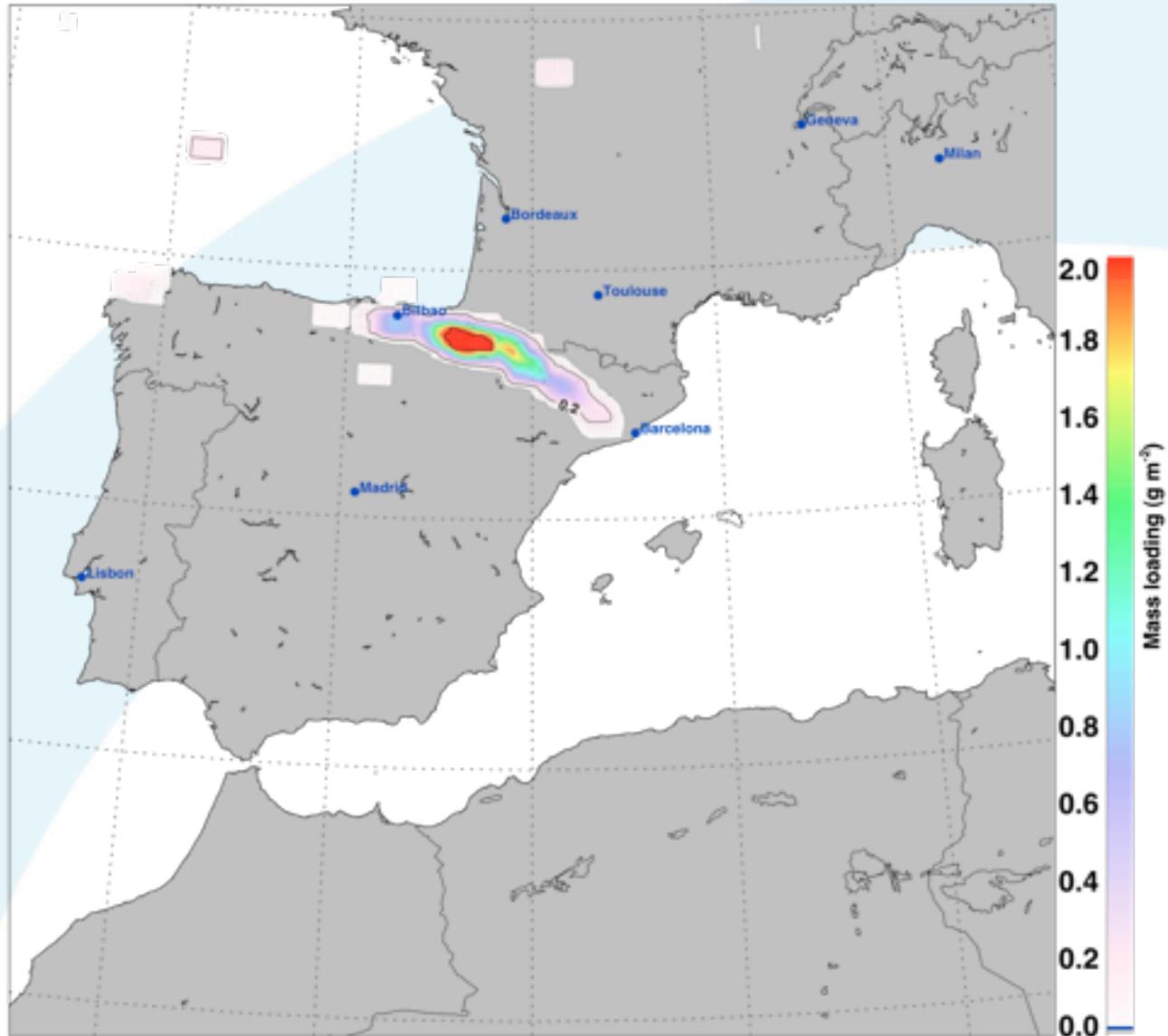




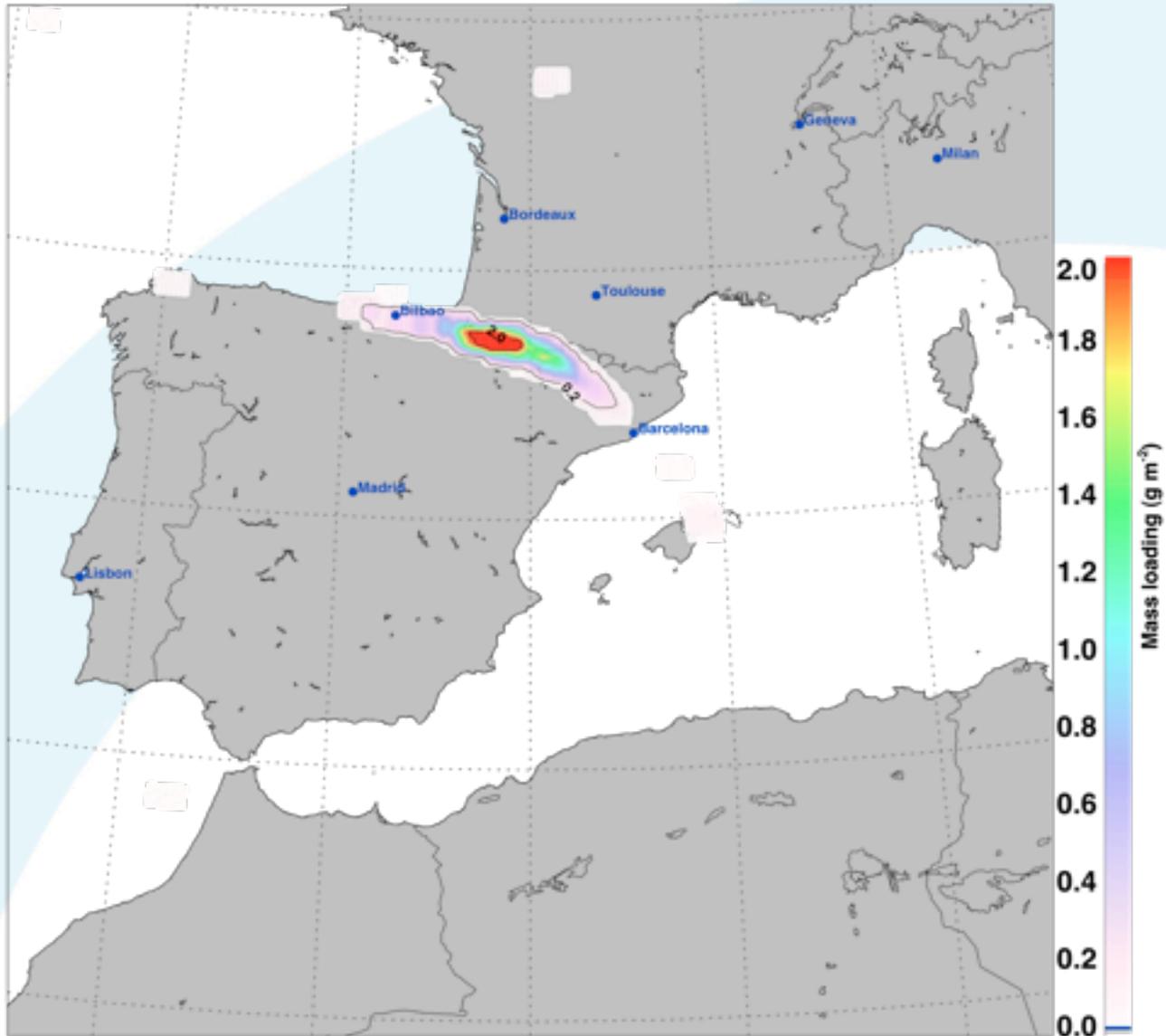
Date: 2010.05.08 09:00 UT



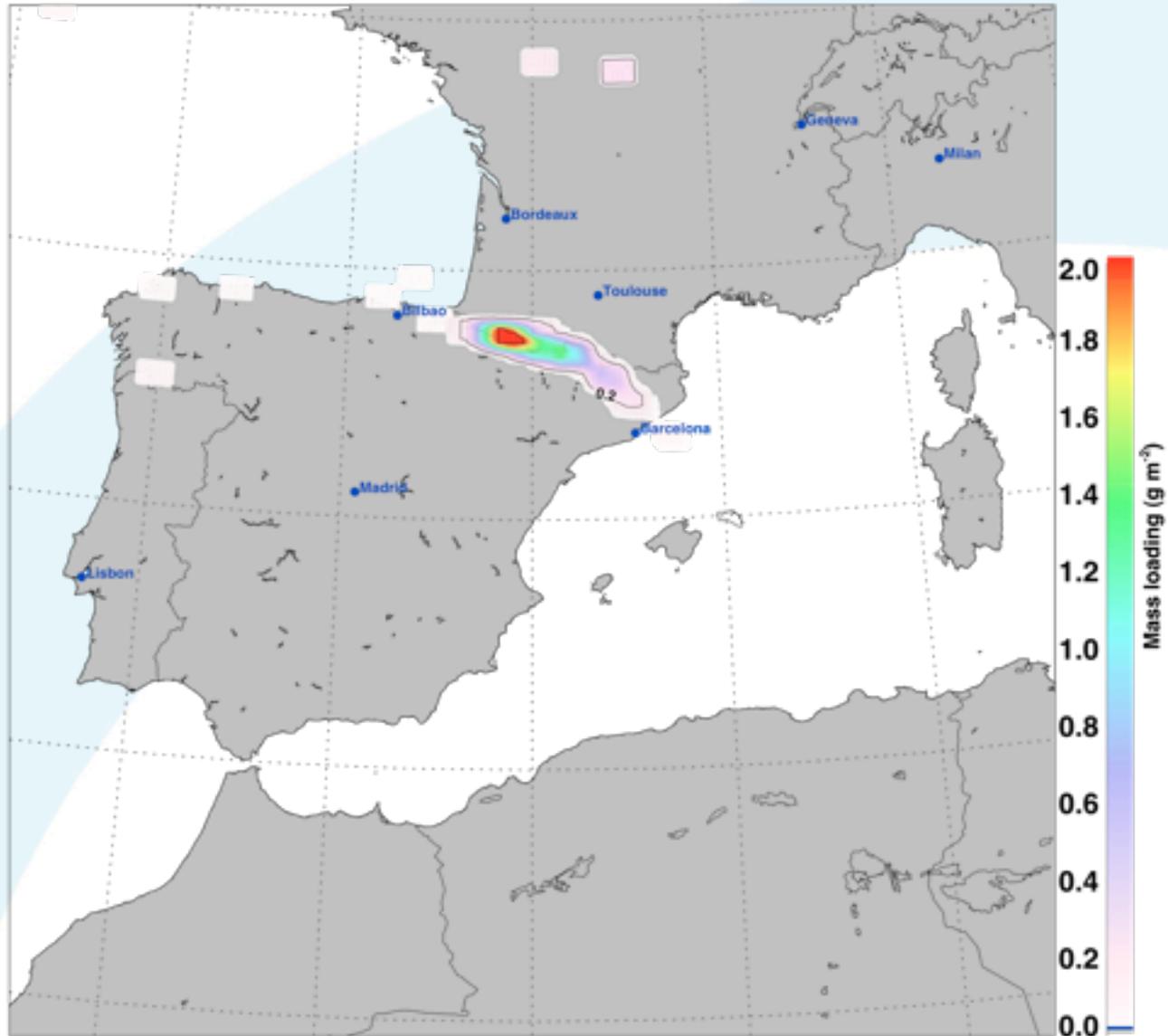
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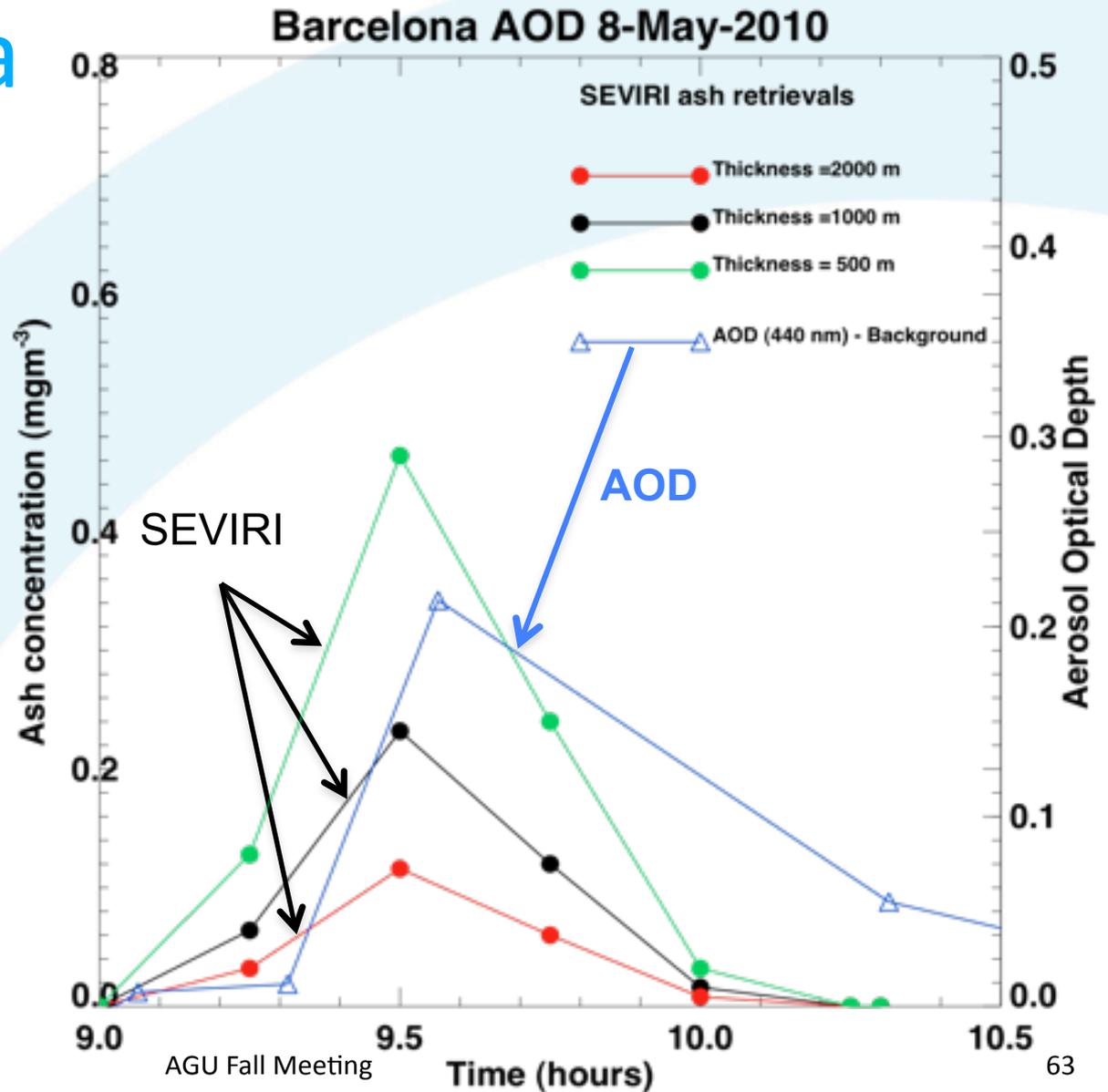
Date:2010.05.08 09:30 UT



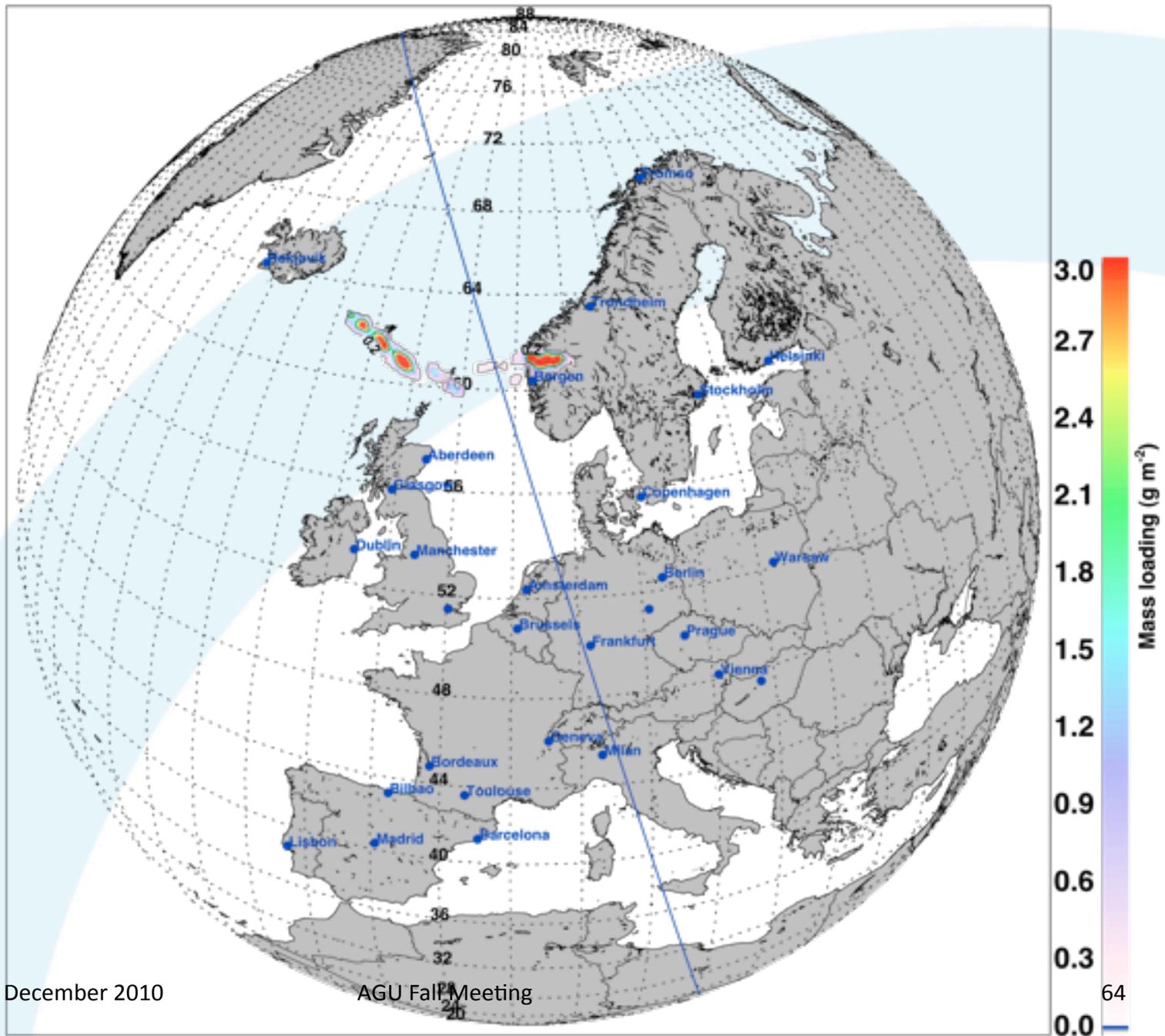
Date:2010.05.08 10:00 UT



# Sun photometer at Barcelona



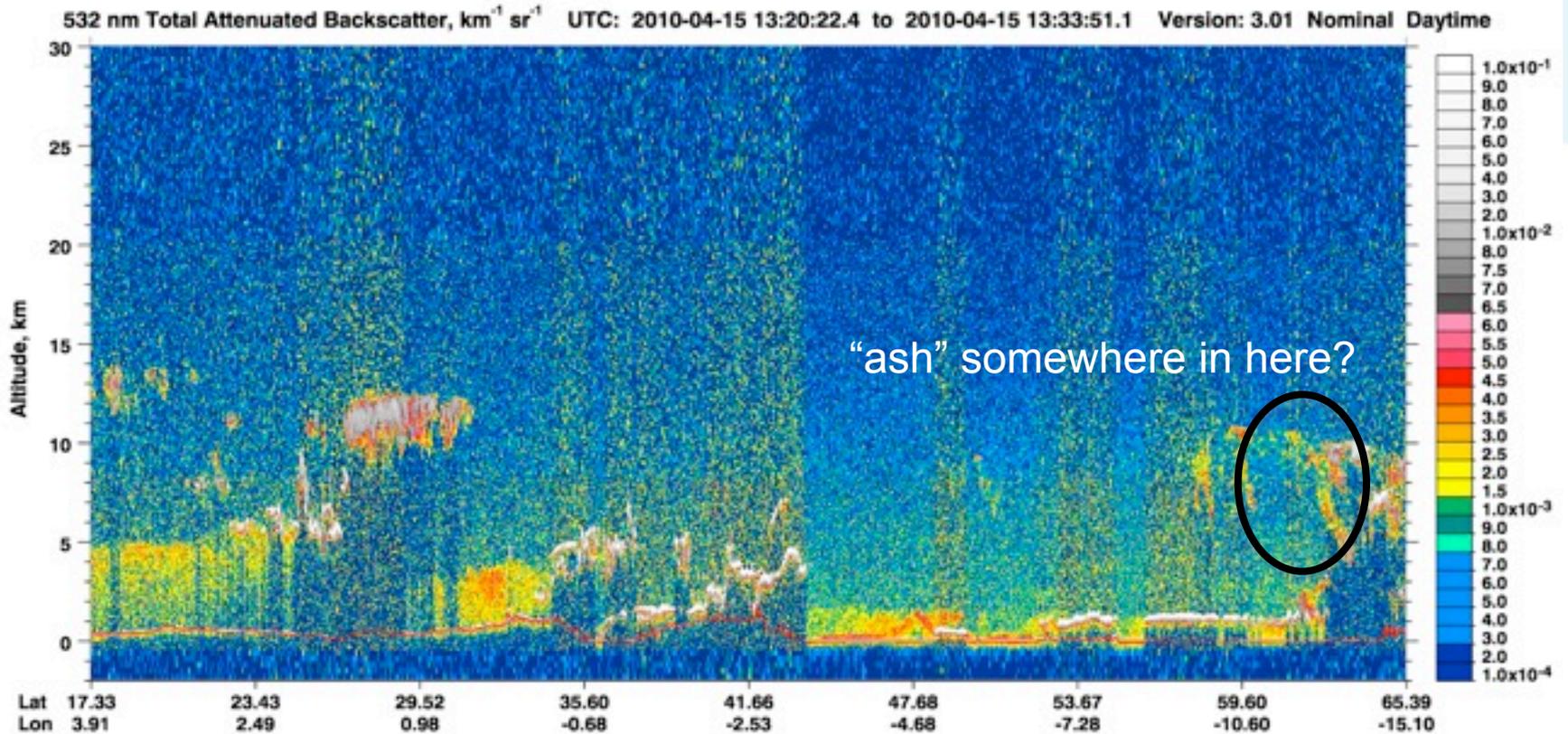
Date:2010.04.15 13:30 UT



17 December 2010

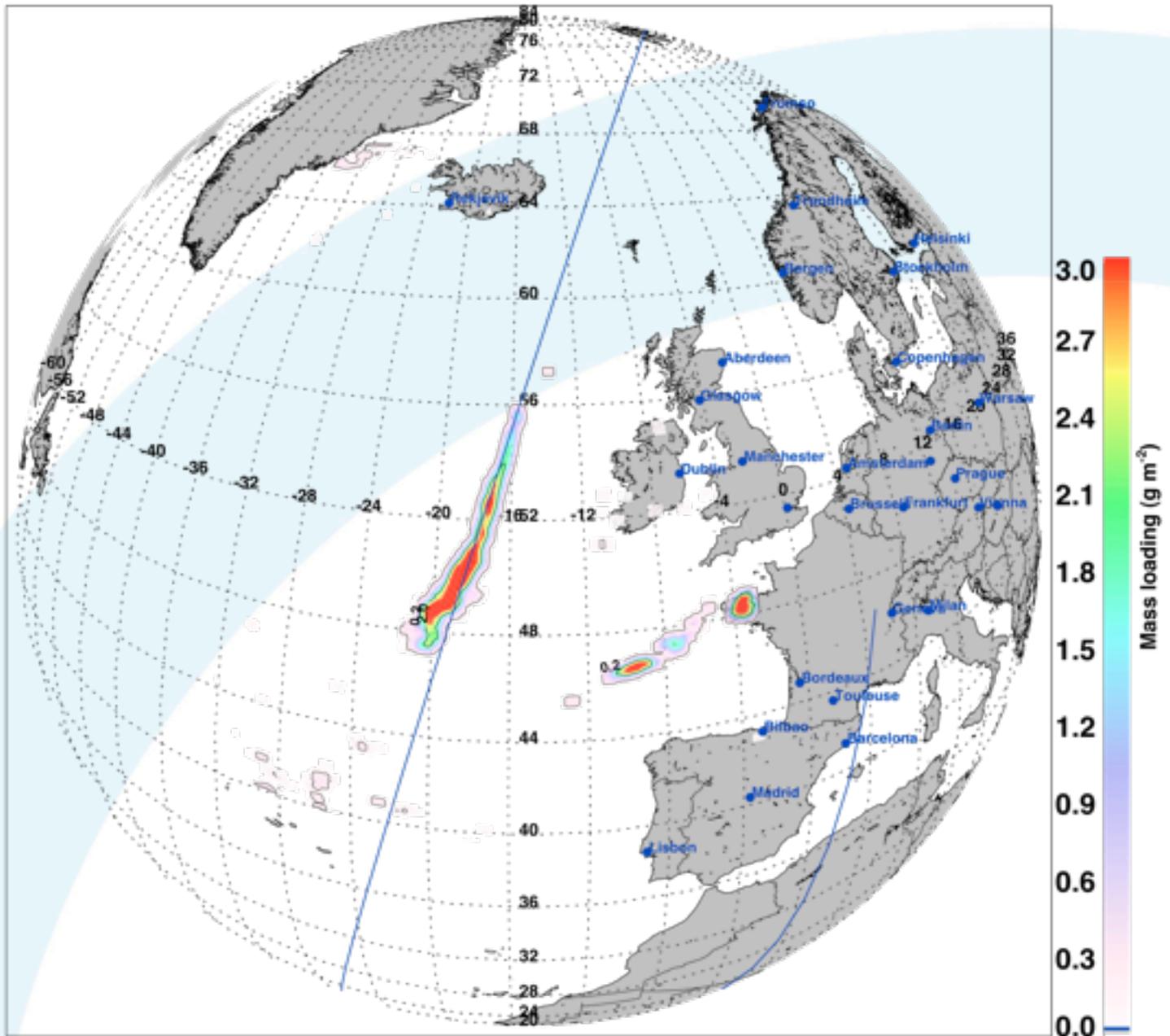
AGU Fall Meeting

# CALIPSO



CALIOP “curtain” 532 nm Total Backscatter

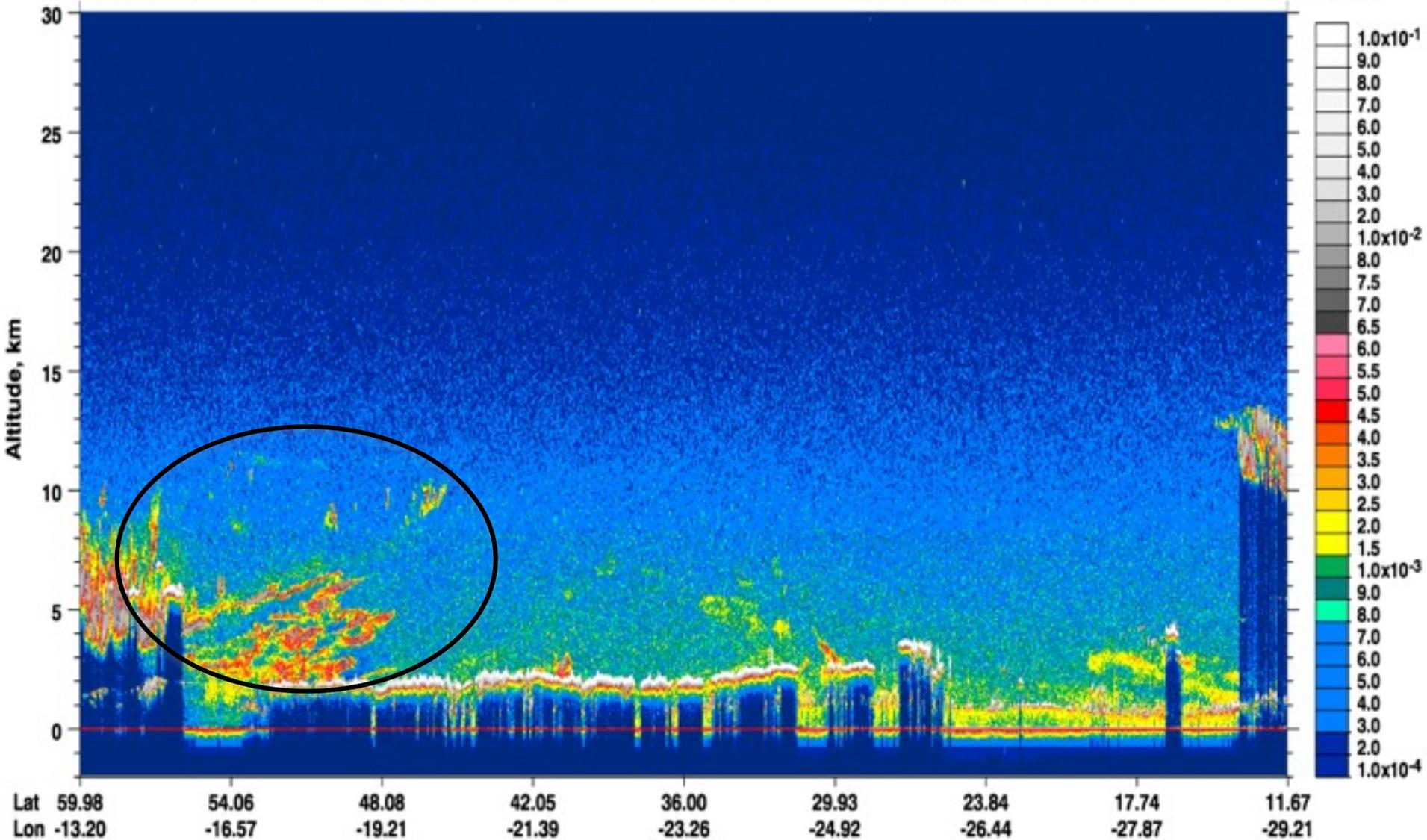
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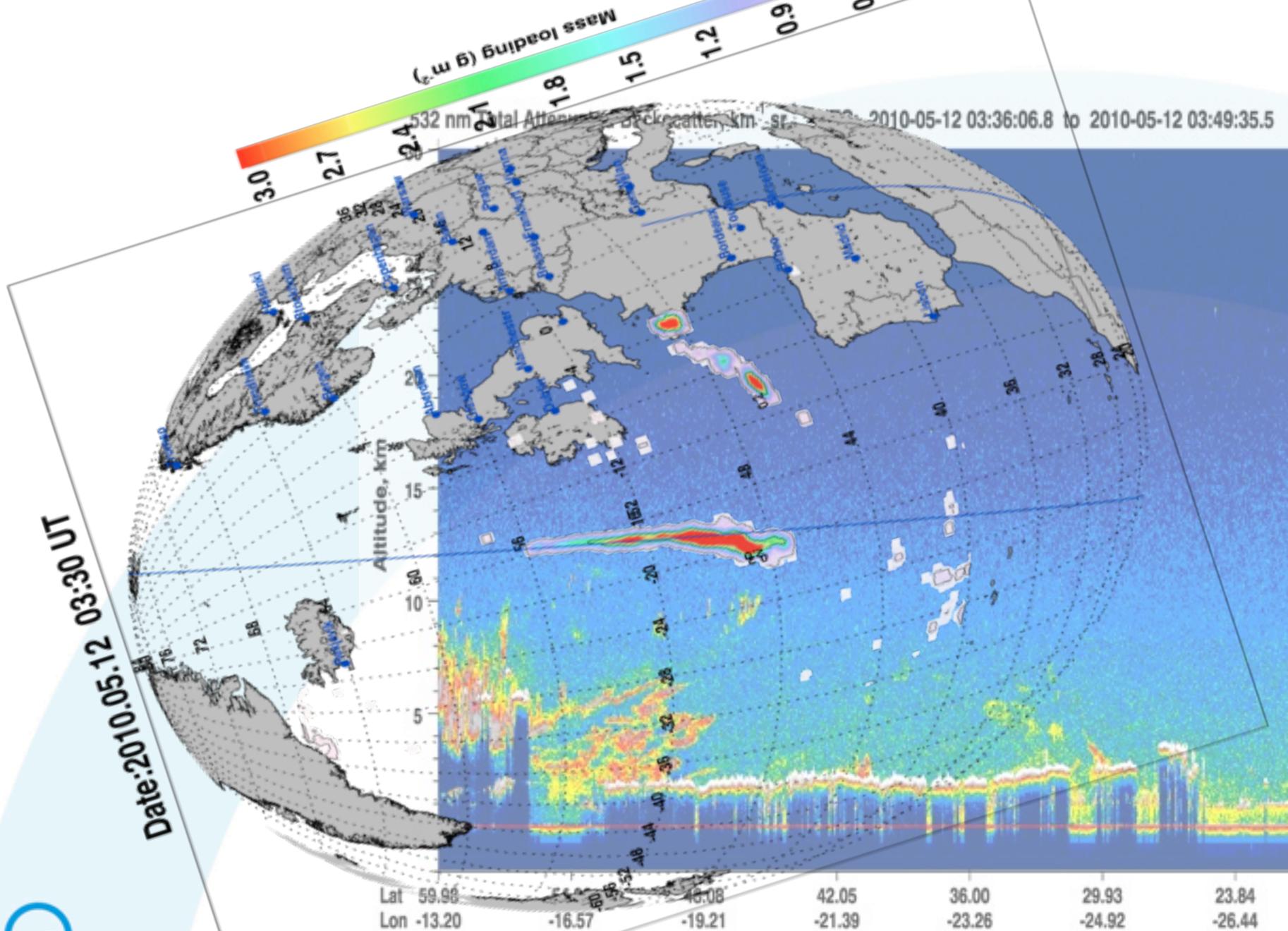


17 December 2010

AGU Fall Meeting

66



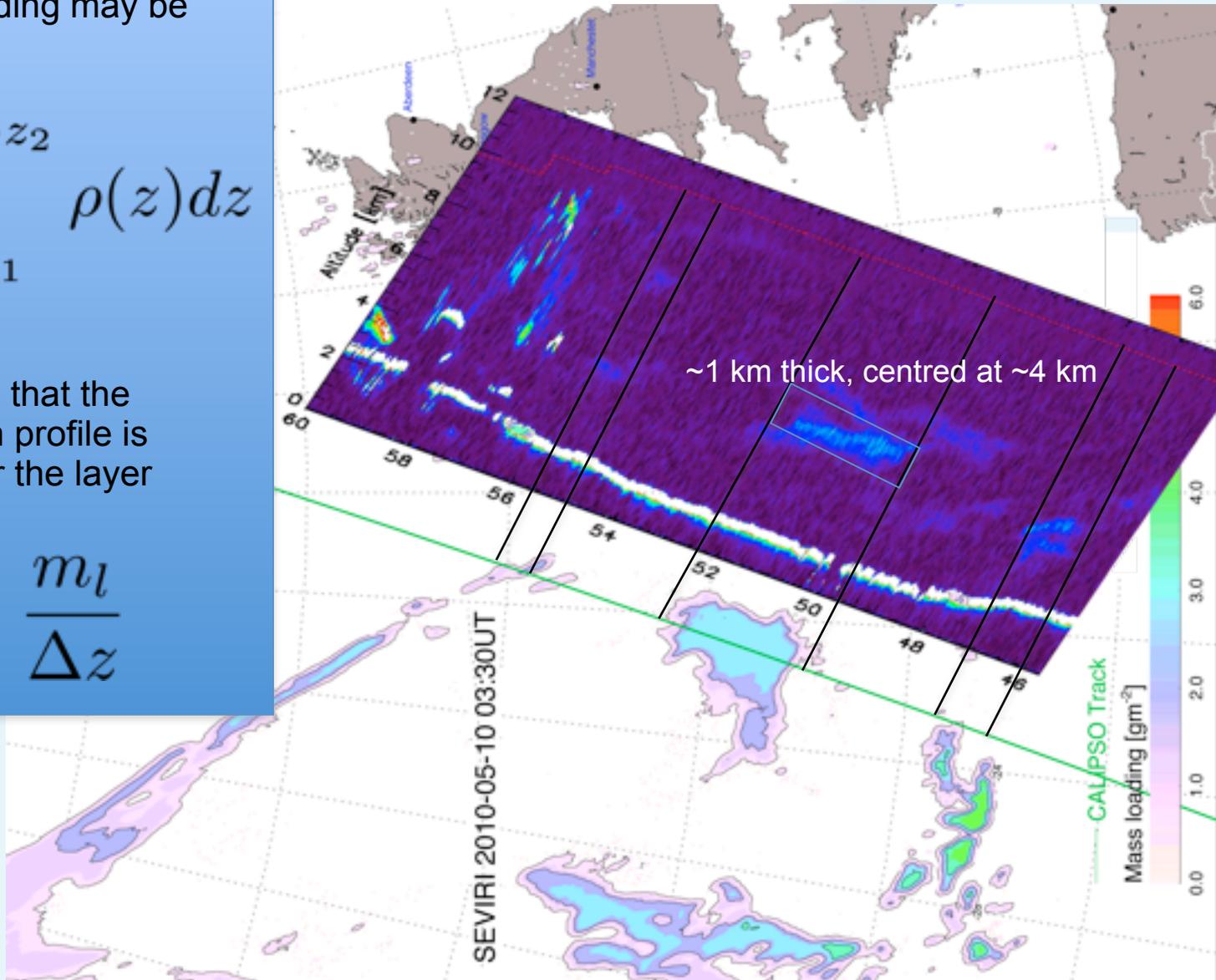


For a single aerosol type the mass loading may be written:

$$m_l = \int_{z_1}^{z_2} \rho(z) dz$$

If we assume that the concentration profile is constant over the layer then:

$$\rho = \frac{m_l}{\Delta z}$$



$$m_l = \int_{z_1}^{z_2} \rho(z) dz$$

$$m_l \sim 0.4 \text{ g m}^{-2}$$

Date: 2010.04.16 14:30 UT

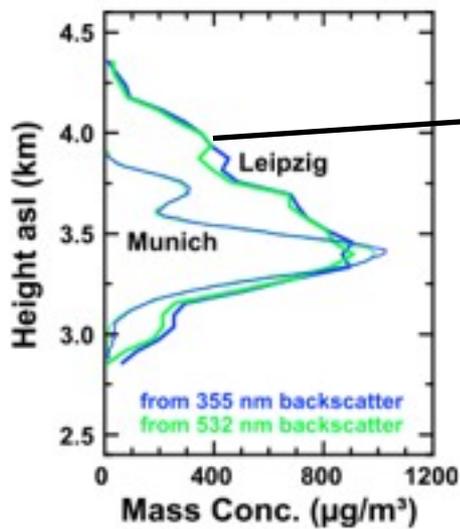
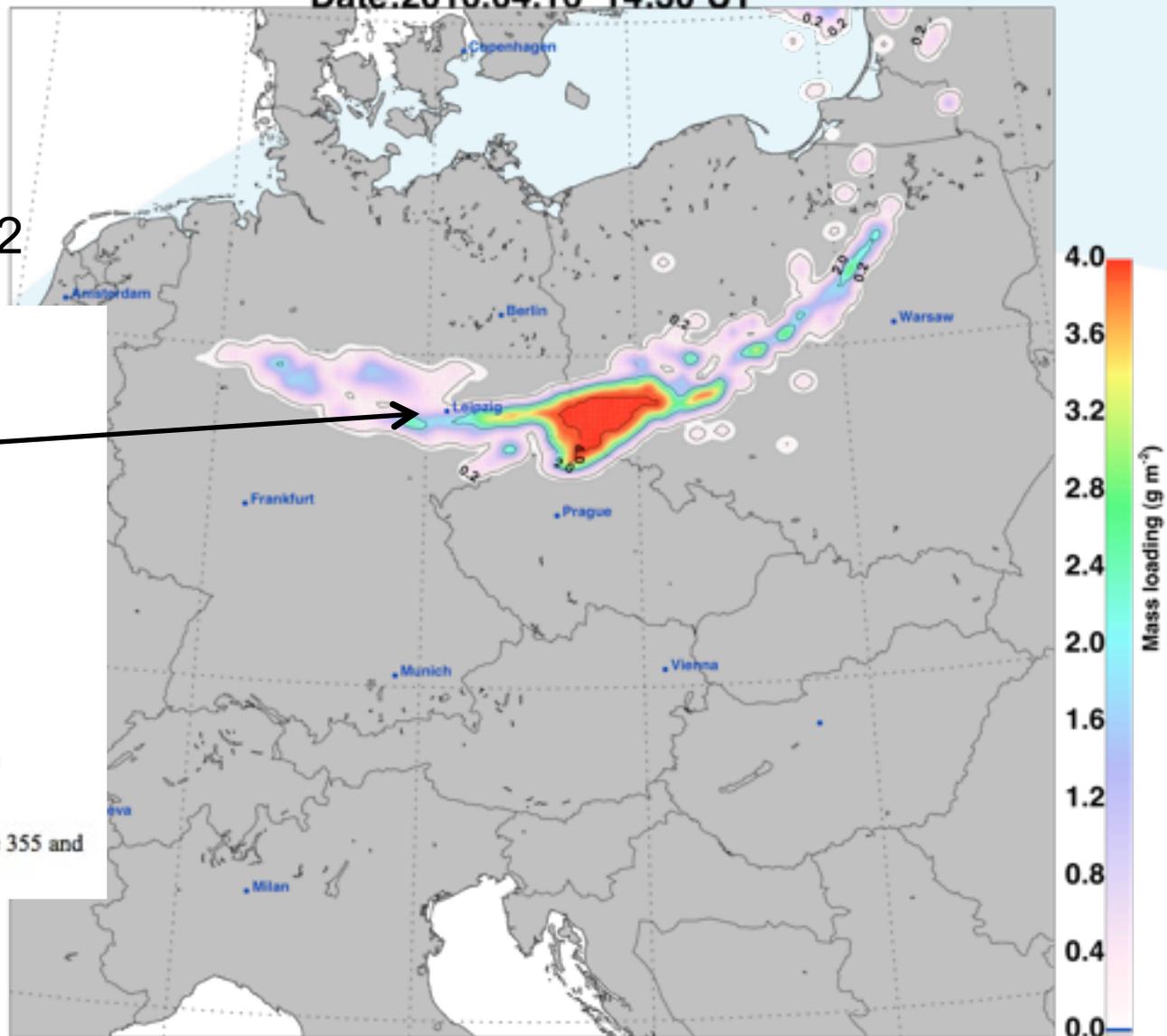
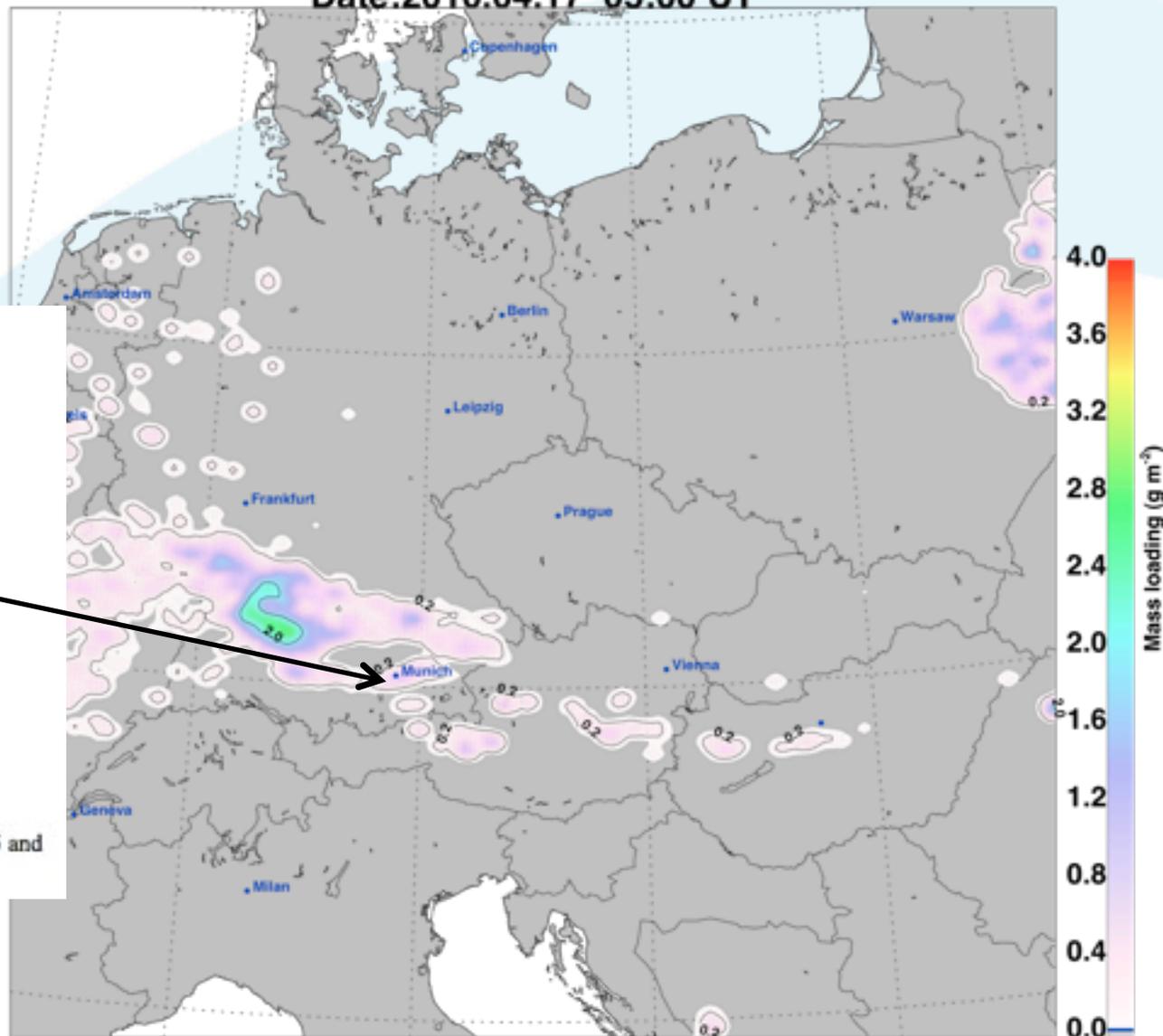


Figure 4. Mass concentrations estimated from the 355 and 532 nm backscatter profiles in Figure 3.



Date:2010.04.17 05:00 UT



$$m_l \sim 0.2 \text{ g m}^{-2}$$

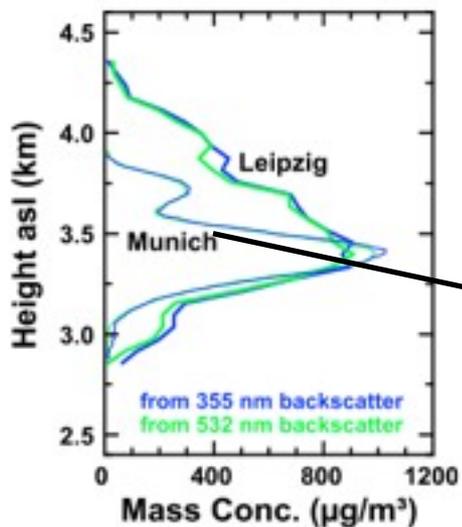


Figure 4. Mass concentrations estimated from the 355 and 532 nm backscatter profiles in Figure 3.

# Visualisation



Courtesy: Vince Realmuto, JPL

Eyjafjallajokull, Iceland

Time: 02:30Z  
Date: 15/05/2010  
Location: 51.50, -2.491  
Ash Loading: 4.75 g/m<sup>2</sup>

London Luton

Courtesy: Andrew Probst, UCLA/JPL

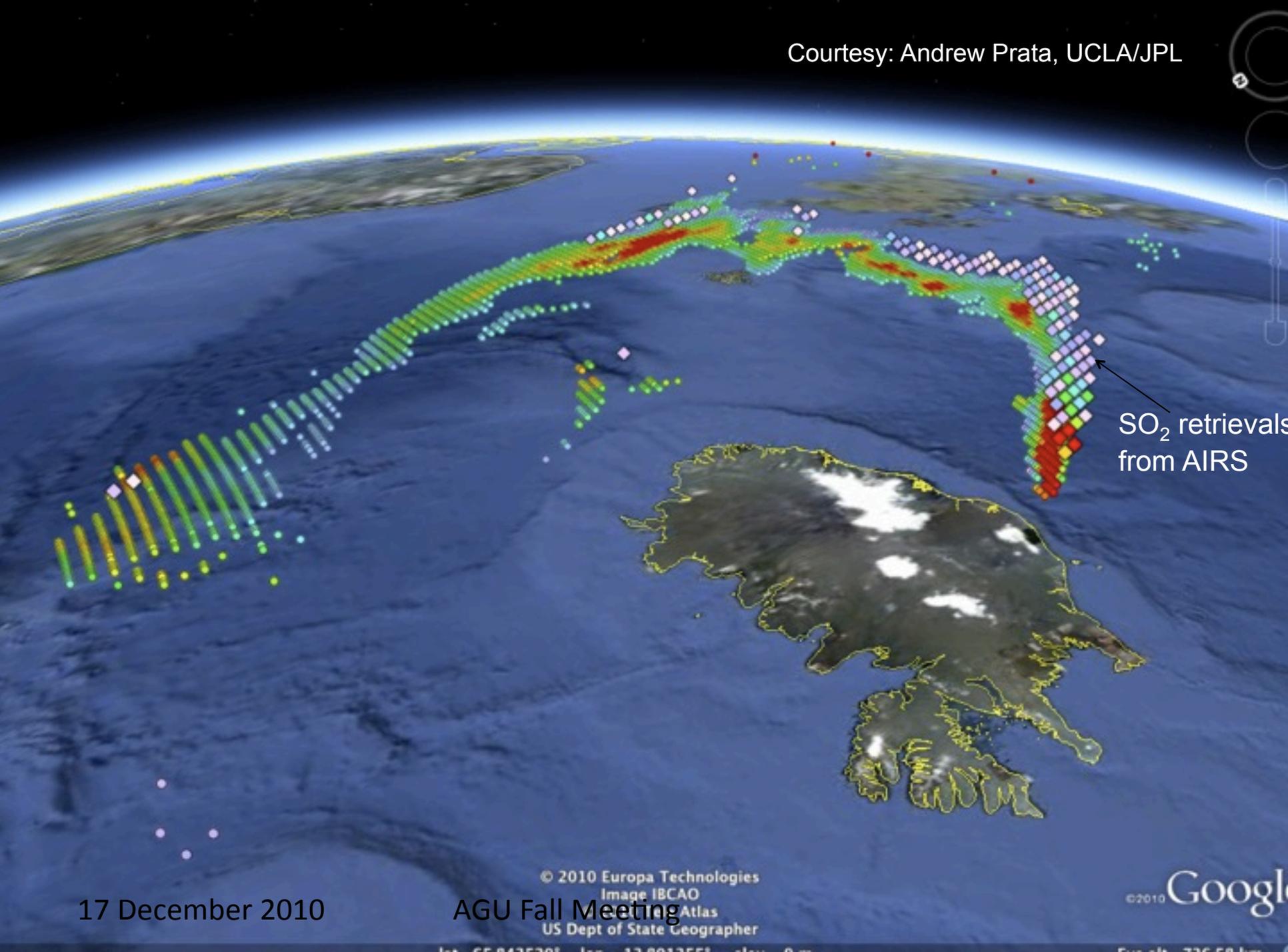
17 December 2010

AGU Fall Meeting

Data SIO, NOAA, U.S. Navy, NGA, GEBCO  
© 2010 InTerra Ltd & Bluesky  
Image © 2010 TerraMetrics  
Image © 2010 IGN-France

©2010 Google

Courtesy: Andrew Prata, UCLA/JPL



SO<sub>2</sub> retrievals  
from AIRS

17 December 2010

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© 2010 Europa Technologies  
Image IBCAO  
US Dept of State Geographer

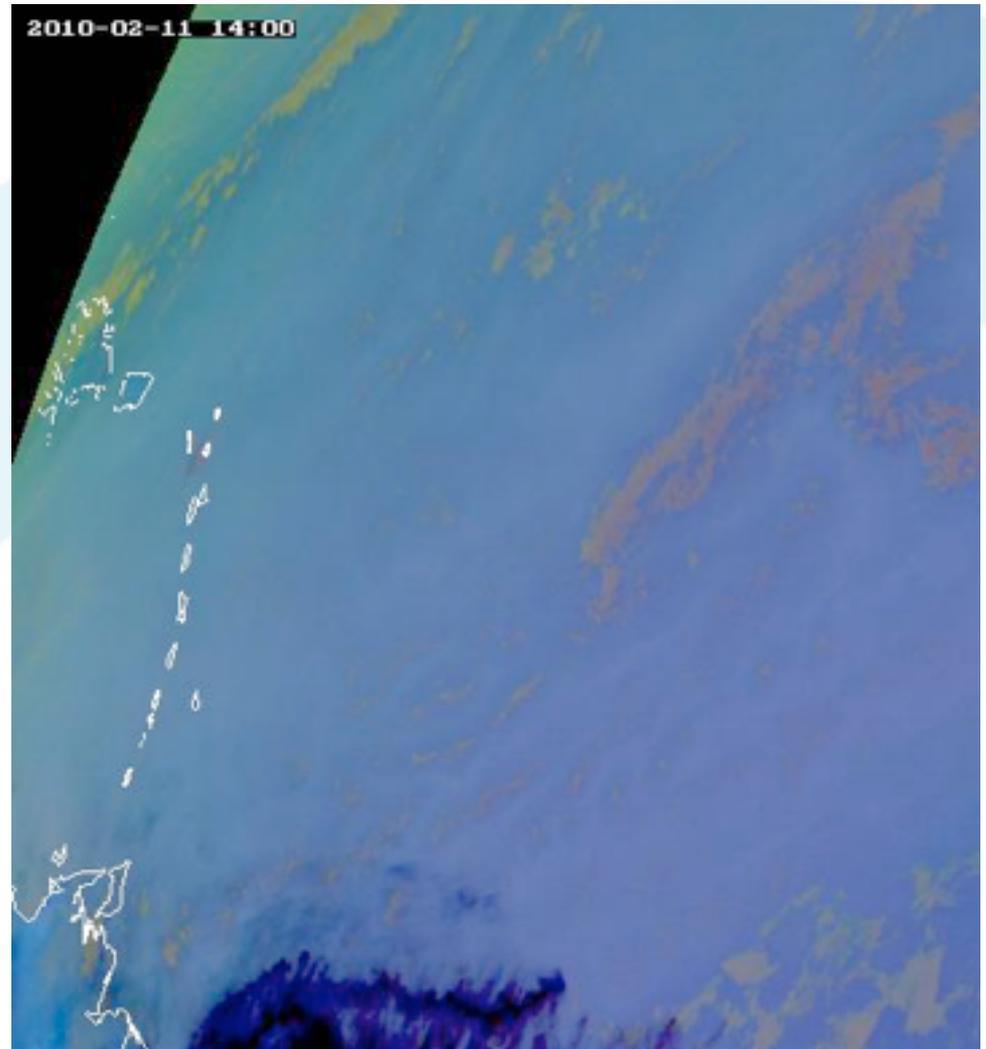
©2010 Google



# SEVIRI

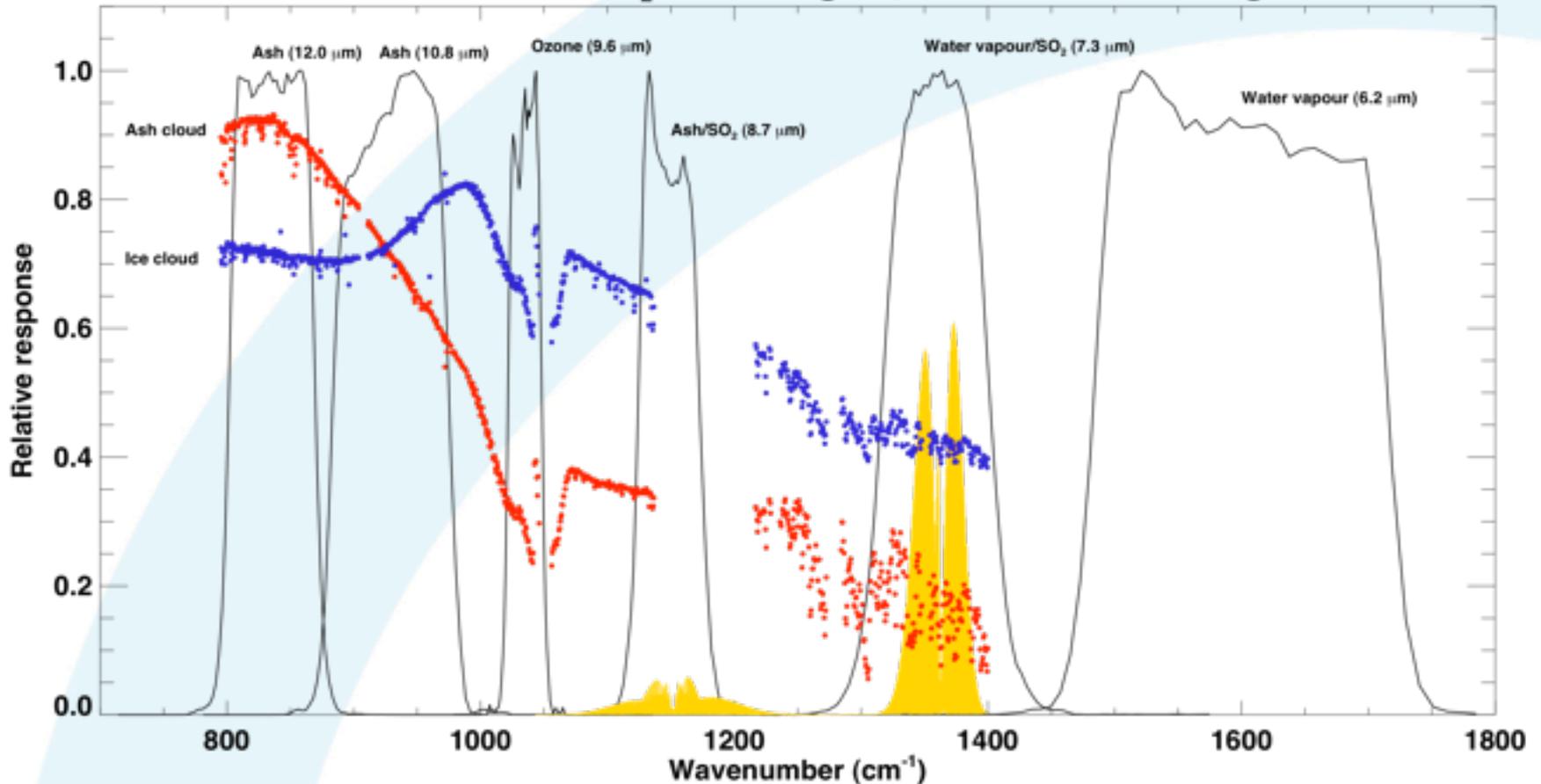
Soufriere Hills  
11-12 February 2010

36 hrs at 15 min resolution

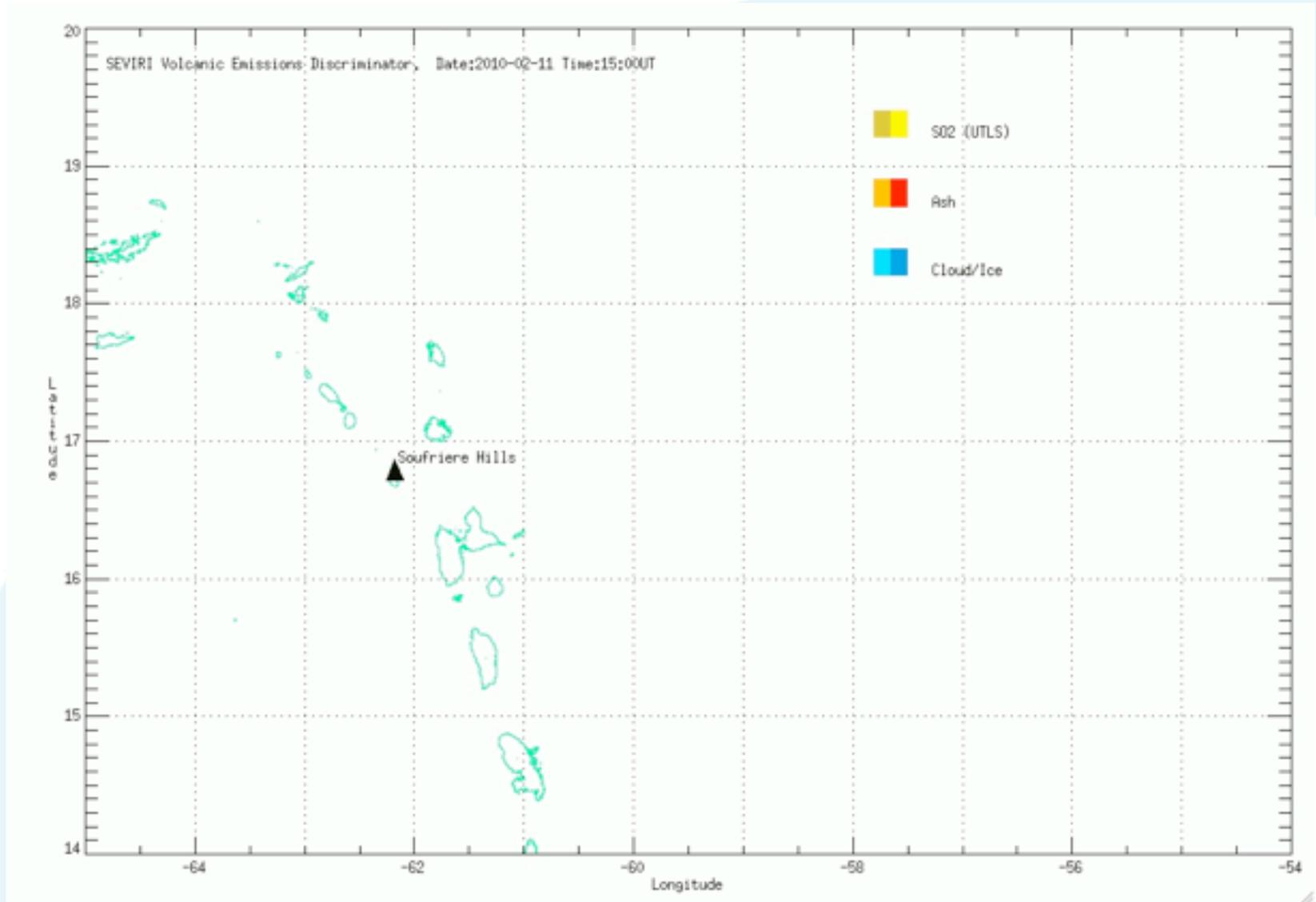


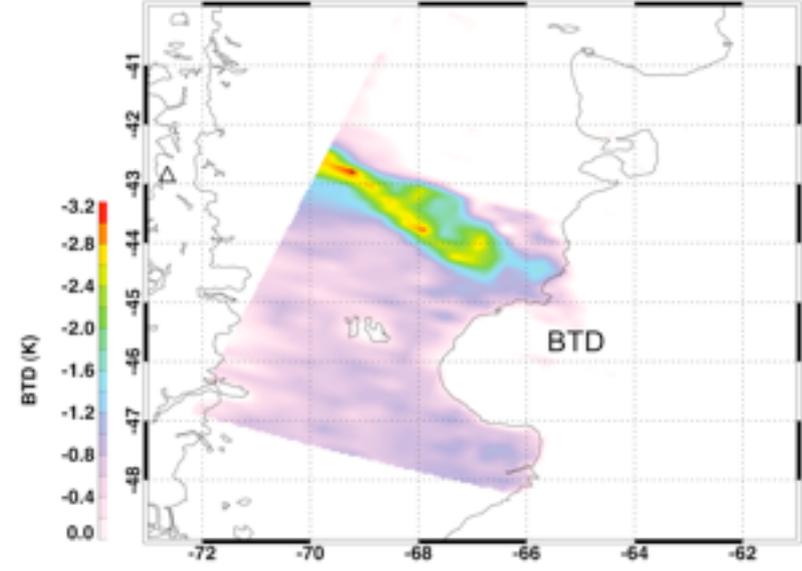
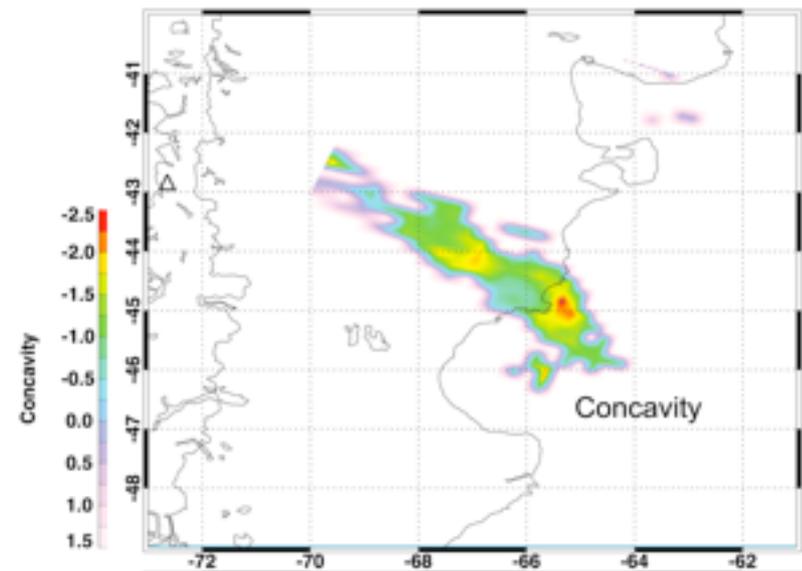
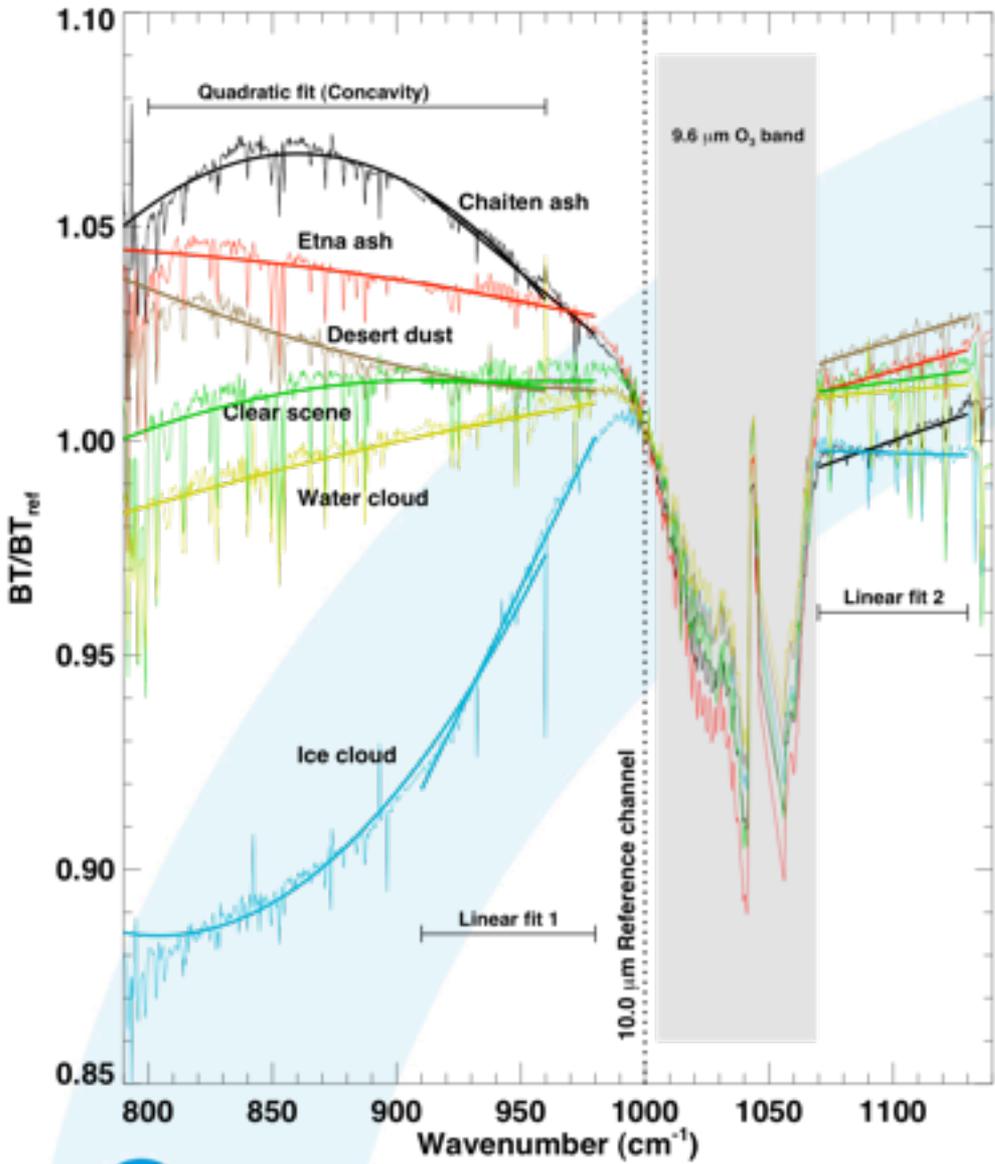
# SEVIRI – triple detection

SEVIRI filter functions, SO<sub>2</sub> line strengths, ash and ice cloud signatures



# SEVIRI – triple detection





# Linear 1 – SO<sub>2</sub>



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## The infrared spectral signature of volcanic ash determined from high-spectral resolution satellite measurements

G. Gangale<sup>a</sup>, A.J. Prata<sup>b,\*</sup>, L. Clarisse<sup>c</sup>

<sup>a</sup> Dipartimento di Ingegneria dei Materiali e dell'Aerospazio, Università di Modena e Reggio Emilia, Italy

<sup>b</sup> Climate and Atmospheric Department, Norwegian Institute for Air Research, PO Box 100, 2007 Kjeller, Norway

<sup>c</sup> Spectroscopie de l'Atmosphère, Service de Chimie Quantique Physicochimie, Université Libre de Bruxelles, Brussels, Belgium

### ARTICLE INFO

#### Article history:

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#### Keywords:

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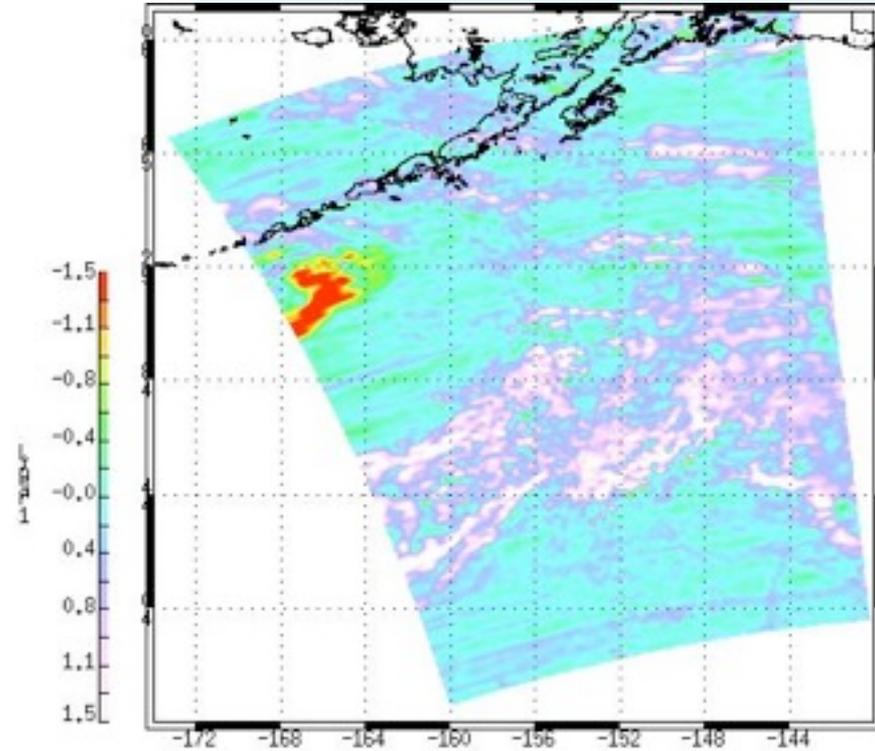
Radiative transfer

Infrared remote sensing

### ABSTRACT

High-spectral resolution infrared spectra of the earth's atmosphere and surface are routinely available from satellite sensors, such as the Atmospheric Infrared Sounder (AIRS) and the Infrared Atmospheric Sounding Interferometer (IASI). We exploit the spectral content of AIRS data to demonstrate that airborne volcanic ash has a unique signature in the infrared (8–12  $\mu\text{m}$ ) that can be used to infer particle size, infrared opacity and composition. The spectral signature is interpreted with the aid of a radiative transfer model utilizing the optical properties of andesite, rhyolite and quartz. Based on the infrared spectral signature, a new volcanic ash detection algorithm is proposed that can discriminate volcanic ash from other airborne substances and we show that the algorithm depends on particle size, optical depth and composition. The new algorithm has an improved sensitivity to optically thin ash clouds, and hence can detect them for longer (~4 days) and at greater distances from the source (~5000 km).

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# Linear 2 – SO<sub>2</sub>



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<sup>b</sup> Climate and Atmospheric Department, Norwegian Institute for Air Research, PO Box 100, 2027 Kjeller, Norway

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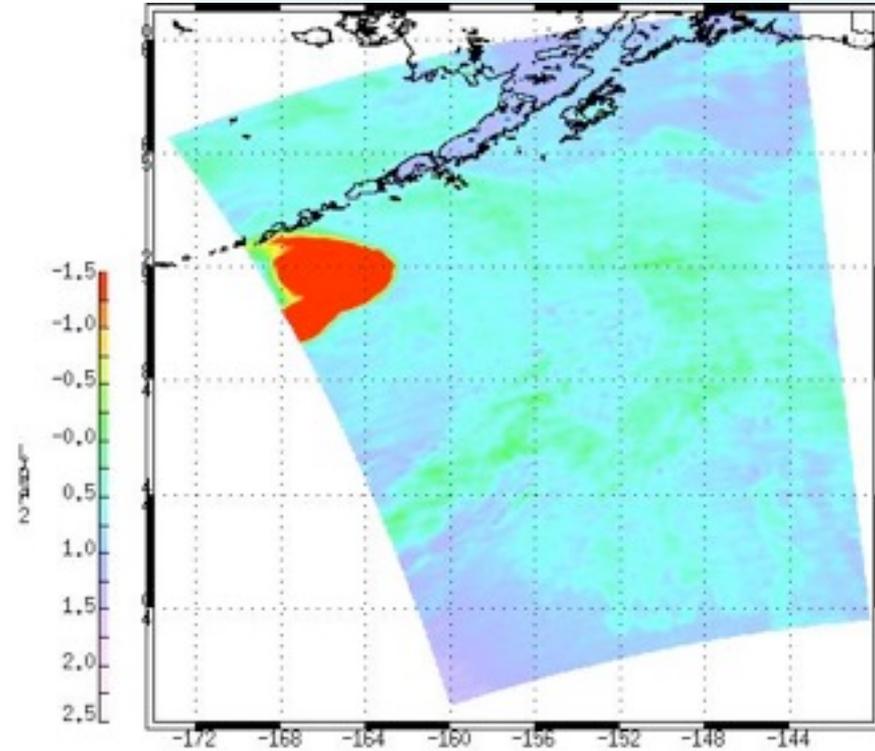
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# Concavity



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## The infrared spectral signature of volcanic ash determined from high-spectral resolution satellite measurements

G. Gangale<sup>a</sup>, A.J. Prata<sup>b,\*</sup>, L. Clarisse<sup>c</sup>

<sup>a</sup> Dipartimento di Ingegneria dei Materiali e dell'Ambiente, Università di Modena e Reggio Emilia, Italy

<sup>b</sup> Climate and Atmospheric Department, Norwegian Institute for Air Research, PO Box 100, 2027 Kjeller, Norway

<sup>c</sup> Spectroscopie de l'Atmosphère, Service de Chimie Quantique Physicochimie, Université Libre de Bruxelles, Brussels, Belgium

### ARTICLE INFO

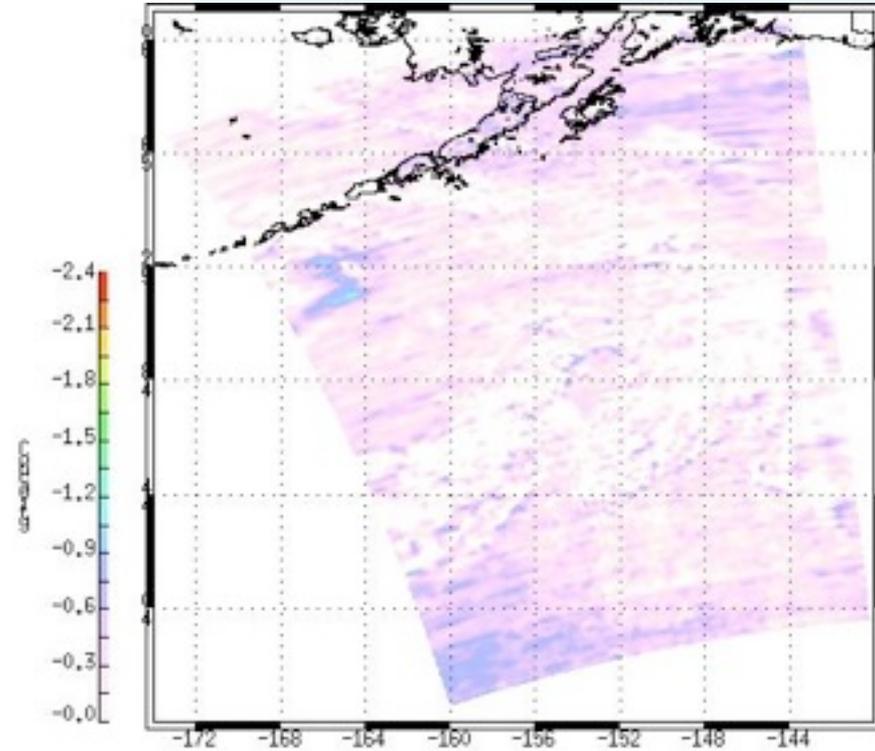
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Received 22 June 2009  
Received in revised form 21 September 2009  
Accepted 23 September 2009  
Available online xxxx

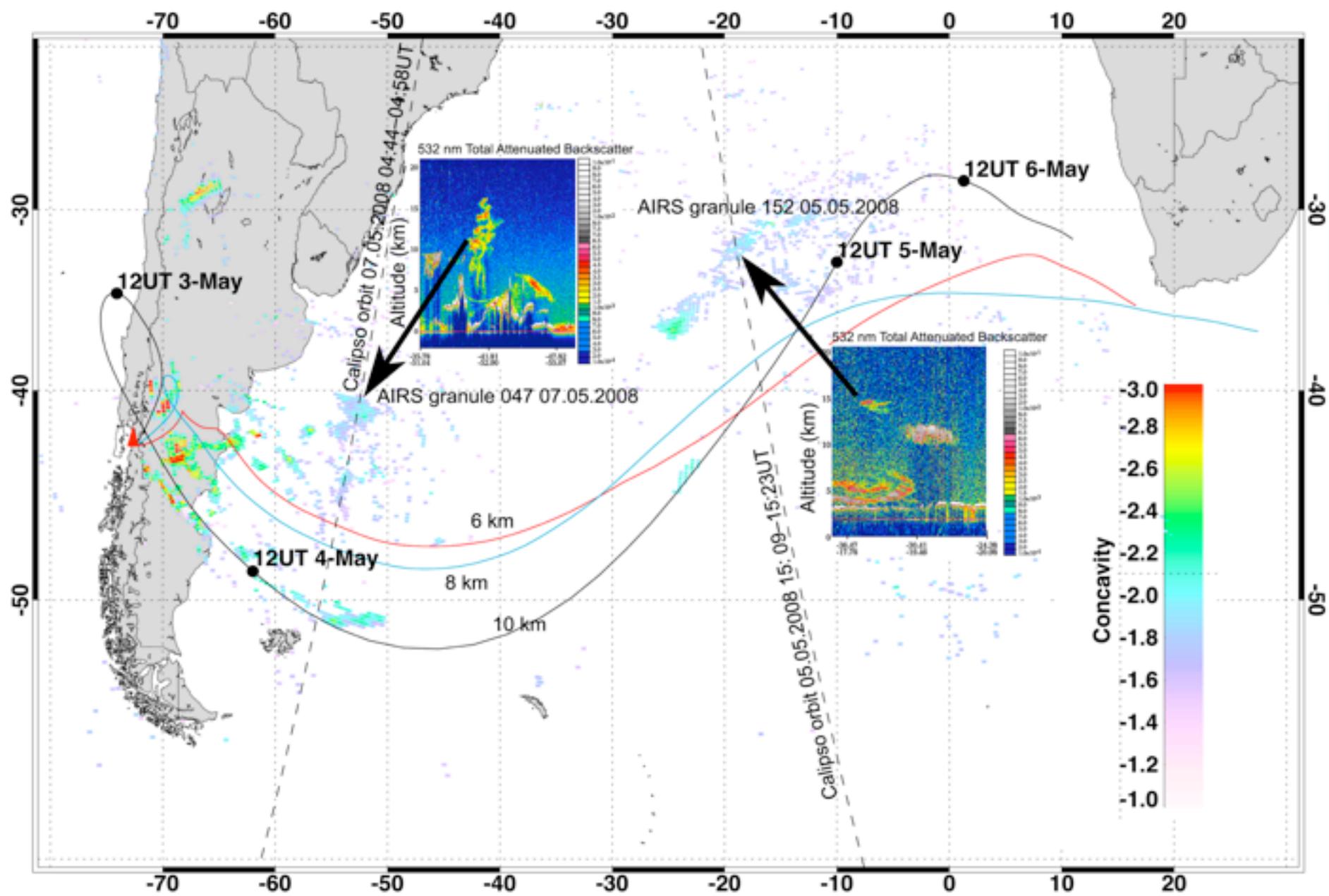
Keywords:  
Volcanic ash  
Radiative transfer  
Infrared remote sensing

### ABSTRACT

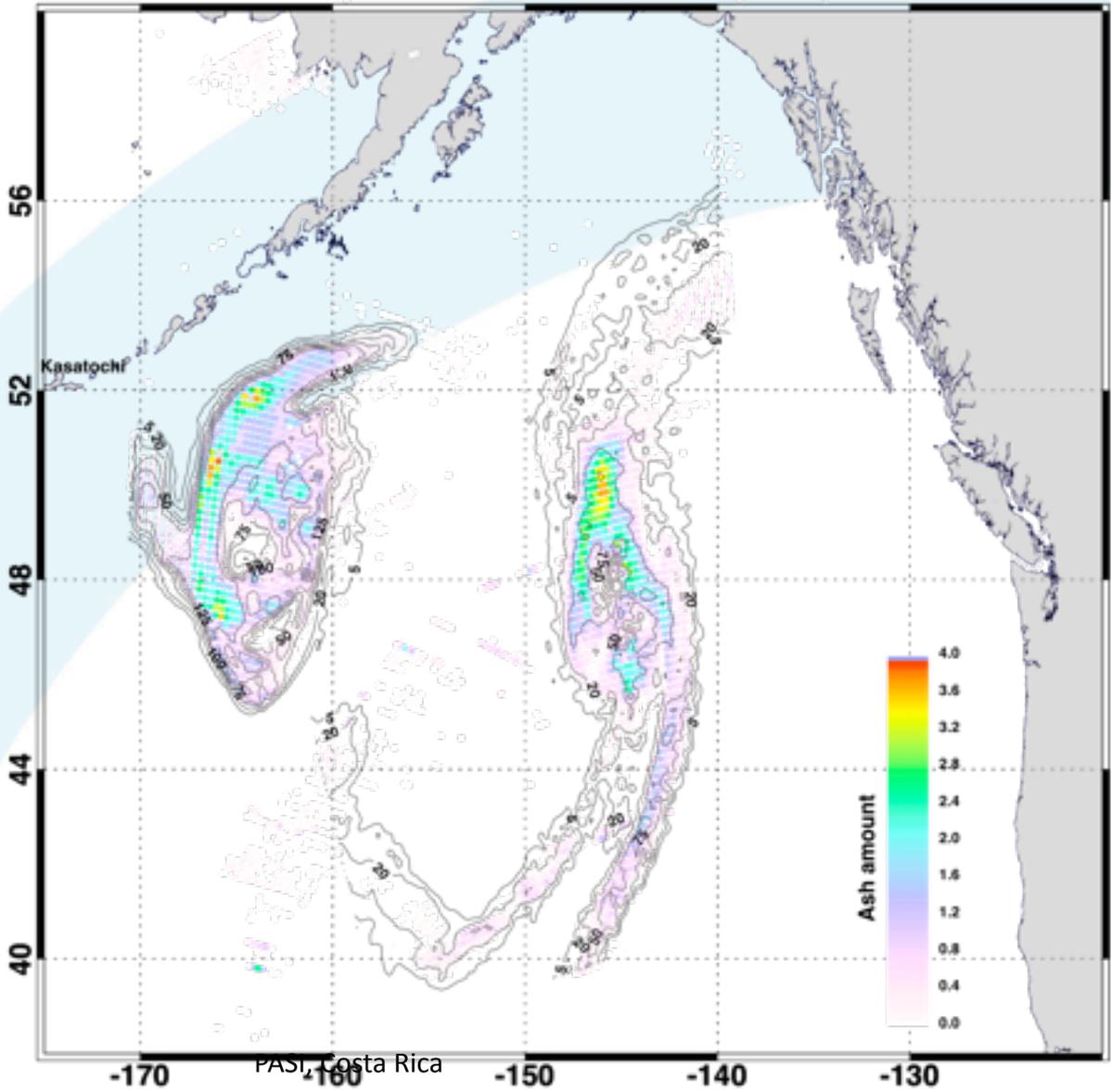
High-spectral resolution infrared spectra of the earth's atmosphere and surface are routinely available from satellite sensors, such as the Atmospheric Infrared Sounder (AIRS) and the Infrared Atmospheric Sounding Interferometer (IASI). We exploit the spectral content of AIRS data to demonstrate that airborne volcanic ash has a unique signature in the infrared (8–12  $\mu\text{m}$ ) that can be used to infer particle size, infrared opacity and composition. The spectral signature is interpreted with the aid of a radiative transfer model utilizing the optical properties of andesite, rhyolite and quartz. Based on the infrared spectral signature, a new volcanic ash detection algorithm is proposed that can discriminate volcanic ash from other airborne substances and we show that the algorithm depends on particle size, optical depth and composition. The new algorithm has an improved sensitivity to optically thin ash clouds, and hence can detect them for longer (~4 days) and at greater distances from the source (~5000 km).

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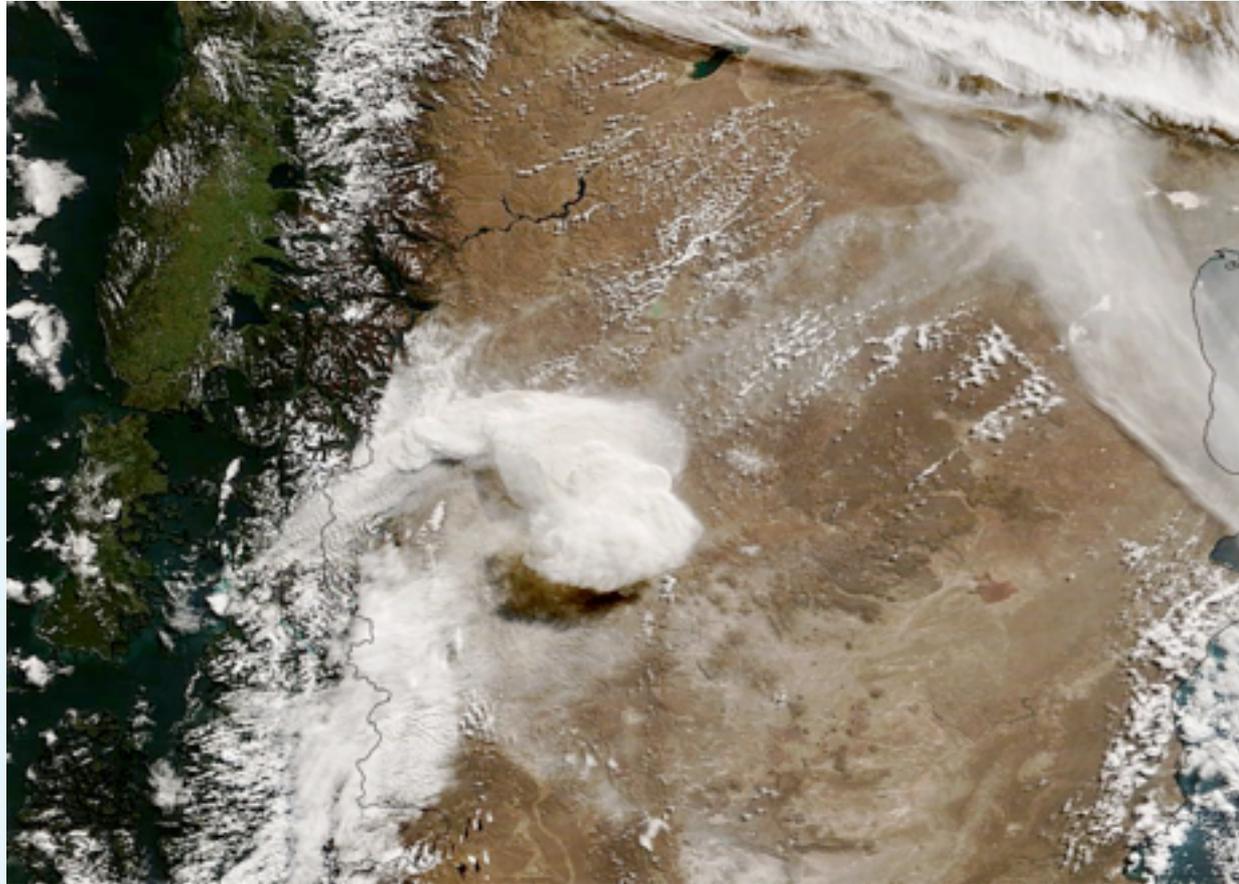
# Collocation of ash and SO<sub>2</sub>



10 January 2011

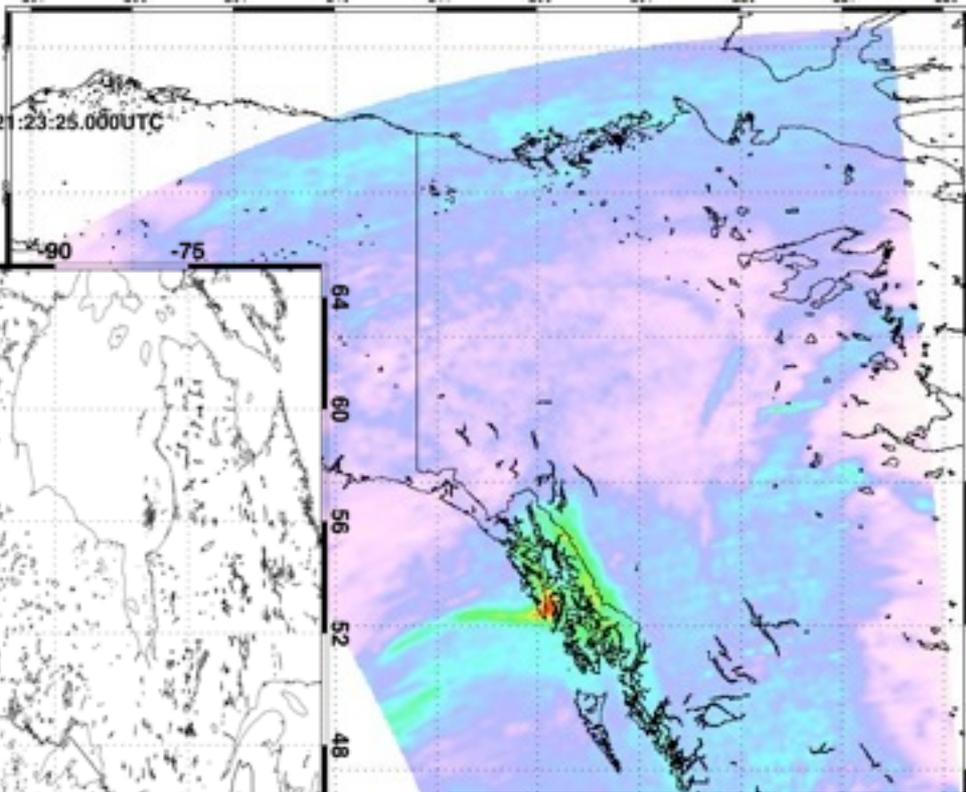
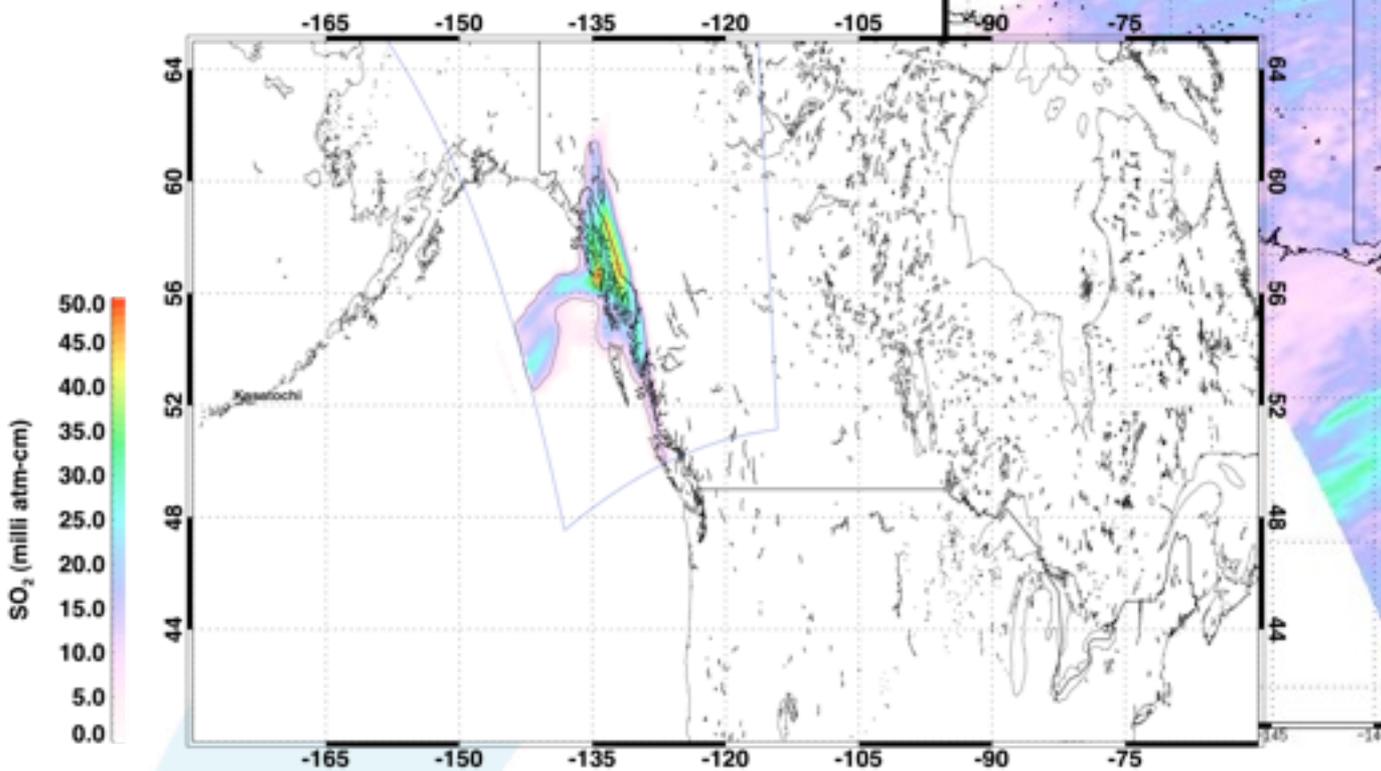
# Chaitén, Chile

First rhyolitic ash eruption in the satellite era (~30 years)

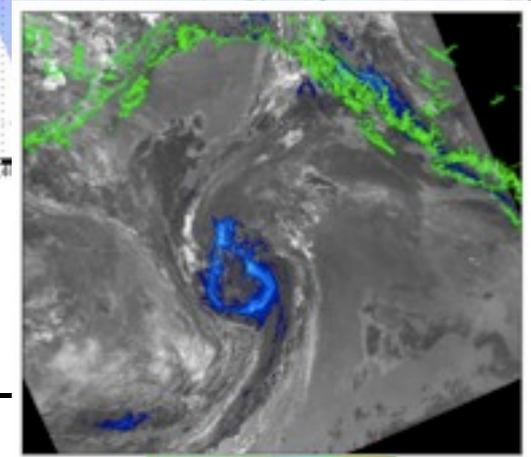


S.A. Carn, J.S. Pallister, L. Lara, J. Ewert, G. Villarosa, M. Fromm, S. Watt, A.J. Prata, D.M. Pyle, T.A. Mather, N. Matthews, R.S. Martin, A. Pavez, R. Aguilera, R. Thomas, W. Rison, P. Krehbiel, J. Johnson, A. Folch, D. Basualto, T.J. Casadevall, M. Guffanti, C. Benitez, J.G. Viramonte, 2008, The awakening of Chaitén volcano, Chile, *EOS Trans.* (to appear).

Mass= 0.1619 Tg Area= 345762 km<sup>2</sup> Max SO<sub>2</sub>= 49.9DU Date: 2008.08.10 21:23:25.000UTC

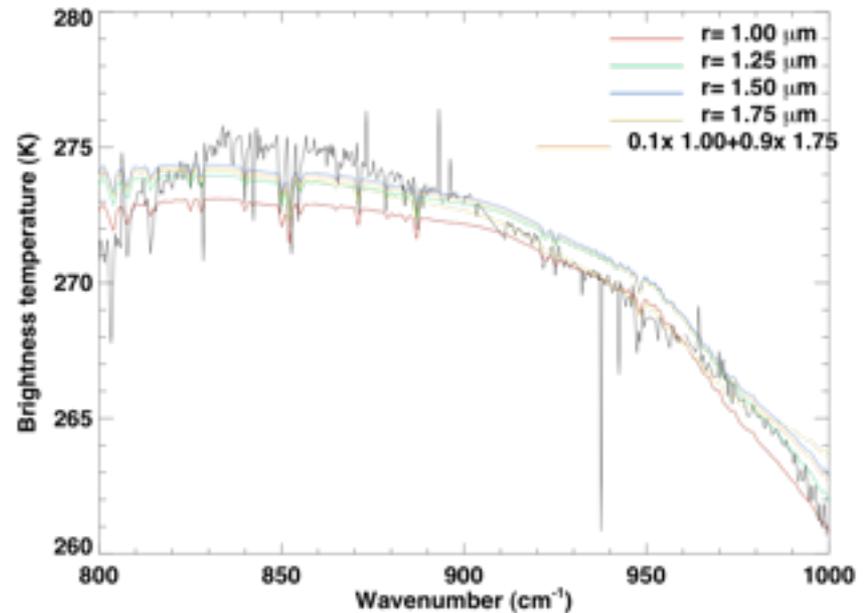
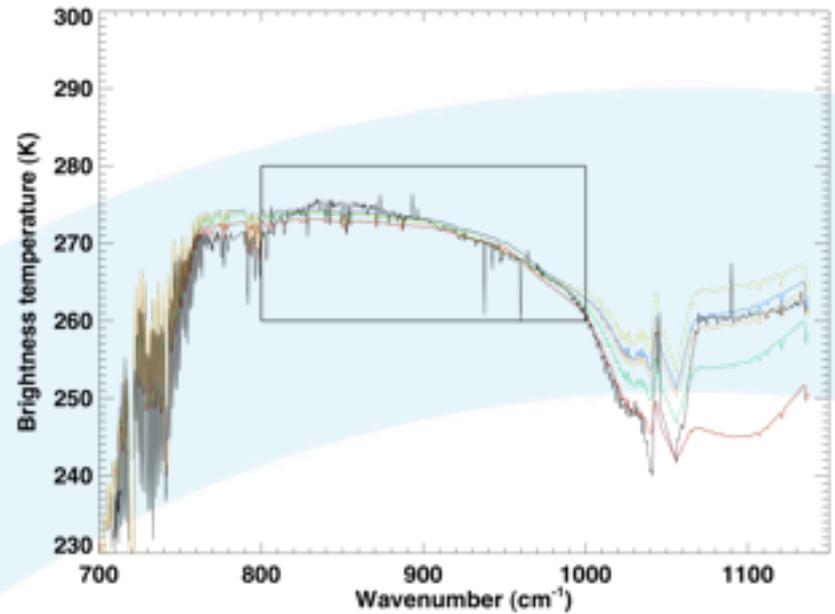


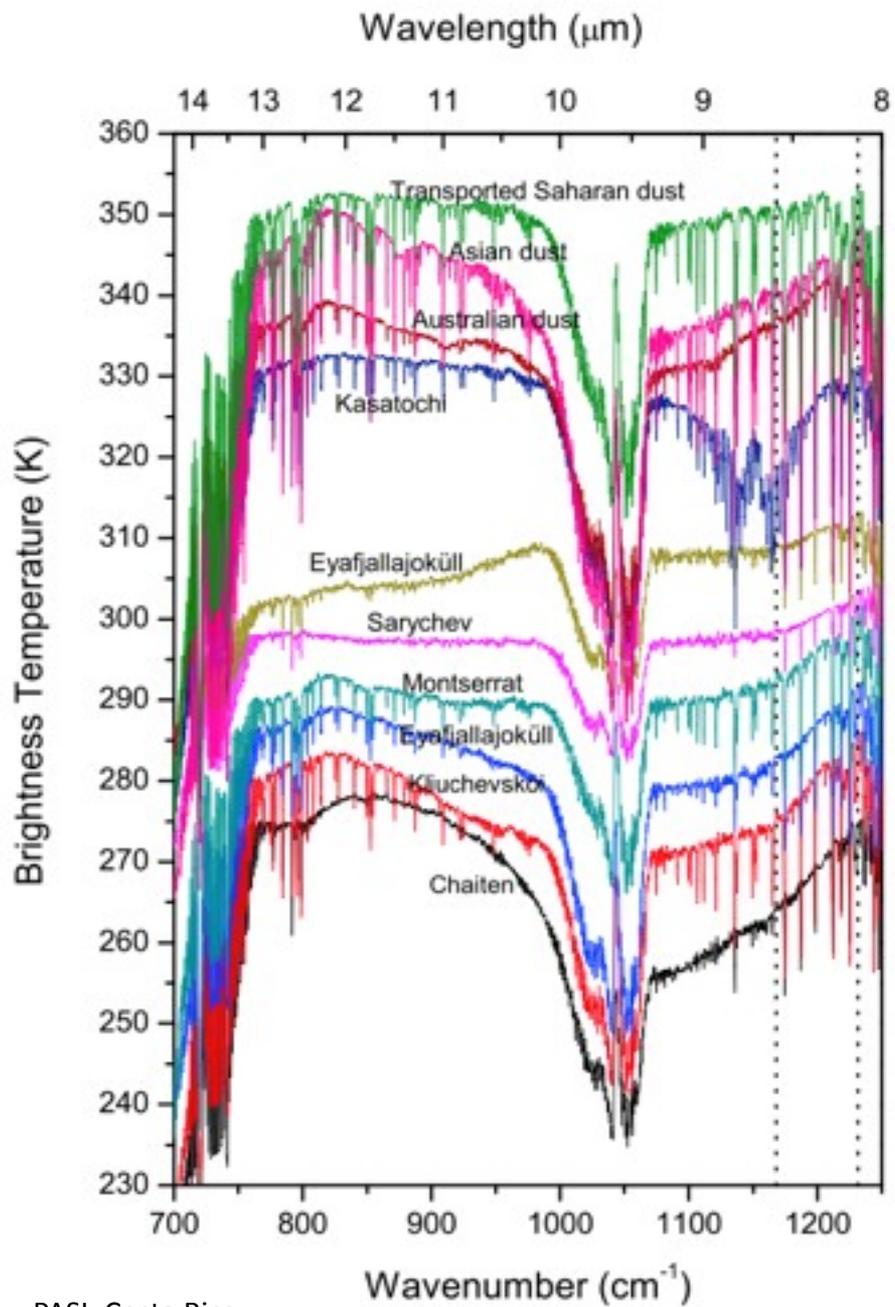
2008.08.10.214 10 August



AVHRR 4m5 - NOAA-18 2008/08/10 22:31 UTC

# Sensitivity to particle size and





## A correlation method for volcanic ash detection using hyperspectral infrared measurements

Lieven Clarisse,<sup>1</sup> Fred Prata,<sup>2</sup> Jean-Lionel Lacour,<sup>1</sup> Daniel Hurtmans,<sup>1</sup> Cathy Clerbaux,<sup>1,3</sup> and Pierre-François Coheur<sup>1</sup>

Received 23 July 2009; revised 26 August 2010; accepted 27 August 2010; published 12 October 2010.

[1] Remote satellite detection of airborne volcanic ash is important for mitigating hazards to aviation and for calculating plume altitudes. Infrared sounders are essential for detecting ash, as they can distinguish aerosol type and can be used day and night. While broadband sensors are mainly used for this purpose, they have inherent limitations. Typically, water and ice can mask volcanic ash, while wind blown dust can yield false detection. High spectral resolution sounders should be able to overcome some of these limitations. However, existing detection methods are not easily applicable to hyperspectral sounders and there is therefore a pressing need for novel techniques. In response, we propose a sensitive and robust volcanic ash detection method for hyperspectral sounders based on correlation coefficients and demonstrate it on IASI observations. We show that the method differentiates ash from clouds and dust. Easy to implement, it could contribute to operational volcanic hazard mitigation. **Citation:** Clarisse, L., F. Prata, J.-L. Lacour, D. Hurtmans, C. Clerbaux, and P.-F. Coheur (2010), A correlation method for volcanic ash detection using hyperspectral infrared measurements, *Geophys. Res. Lett.*, 37, L19806, doi:10.1029/2010GL044828.

infrared imagers (e.g., MODIS, AVHRR, GOES, SEVIRI). Most of the available algorithms find their origin in the reverse absorption method [Prata, 1989; Zipper *et al.*, 2004]. It exploits the differential absorption between channels in the region 10 to 12  $\mu\text{m}$ . While water and ice clouds absorb preferentially towards 12  $\mu\text{m}$ , ash clouds absorb preferentially towards 10  $\mu\text{m}$ . There are a number of limitations to the reverse absorption technique [Pavolonis *et al.*, 2006], such as the fact that water and ice clouds can have a masking effect on the observation of volcanic ash, while wind blown mineral dust can yield false detection (see Figure 1). Several multichannel extensions and new algorithms have been proposed, enhancing both the sensitivity and robustness [see, e.g., Elrod *et al.*, 2003; Pergola *et al.*, 2008; Zipper *et al.*, 2004; Pavolonis and Sieglaff, 2009].

[1] With the advent of the infrared sounders AIRS and IASI, and in the perspective of the planned GEMS Sentinel missions for atmospheric composition, hyperspectral sounders are becoming increasingly important. While it is likely that these can overcome some of the limitations of broadband sensors (Figure 1 is a good example of how a spectrometer can discriminate more than an imager), there has been little progress in developing a universal volcanic ash detection

## Retrieving radius, concentration, optical depth, and mass of different types of aerosols from high-resolution infrared nadir spectra

Lieven Clarisse,<sup>1,\*</sup> Daniel Hurtmans,<sup>1</sup> Alfred J. Prata,<sup>2</sup> Federico Karagulian,<sup>1</sup> Cathy Clerbaux,<sup>3,1</sup> Martine De Mazière,<sup>4</sup> and Pierre-François Coheur<sup>1</sup>

<sup>1</sup>Spectroscopie de l'Atmosphère, Service de Chimie Quantique et Photophysique, Université Libre de Bruxelles (ULB), Brussels, Belgium

<sup>2</sup>Climate and Atmosphere Department, Norwegian Institute for Air Research, P.O. Box 100, 2027 Kjeller, Norway

<sup>3</sup>UPMC University Paris 06, Université Versailles Saint Quentin; CNRS/INSU, LATMOS-IPSL, Paris, France

<sup>4</sup>Belgian Institute for Space Aeronomy, 3 Avenue Circulaire, B-1180 Brussels, Belgium

\*Corresponding author: lclariss@ulb.ac.be

Received 15 March 2010; revised 12 May 2010; accepted 18 May 2010; posted 2 June 2010 (Doc. ID 125477); published 23 June 2010

We present a sophisticated radiative transfer code for modeling outgoing IR radiation from planetary atmospheres and, conversely, for retrieving atmospheric properties from high-resolution nadir-observed spectra. The forward model is built around a doubling-adding routine and calculates, in a spherical refractive geometry, the outgoing radiation emitted by the Earth and the atmosphere containing one layer of aerosol. The inverse model uses an optimal estimation approach and can simultaneously retrieve atmospheric trace gases, aerosol effective radius, and concentration. It is different from existing codes, as most forward codes dealing with multiple scattering assume a plane-parallel atmosphere, and as for the retrieval, it does not rely on precalculated spectra, the use of microwindows, or two-step retrievals. The simultaneous retrieval on a broad spectral range exploits the full potential of current state-of-the-art hyperspectral IR sounders, such as AIRS and IASI, and should be particularly useful in studying major pollution events. We present five example retrievals of IASI spectra observed in the range from 800 to 1200  $\text{cm}^{-1}$  above dust, volcanic ash, sulfuric acid, ice particles, and biomass burning aerosols. © 2010 Optical Society of America

OCIS codes: 010.1100, 010.0280, 010.5620.

# Three year SO<sub>2</sub> measurements from IASI

Lieven Clarisse, Fred Prata, Andreas Richter, Simon Carn, Daniel Hurtmans, Pierre Coheur, Juliette Hadji-Lazaro, Cathy Clerbaux



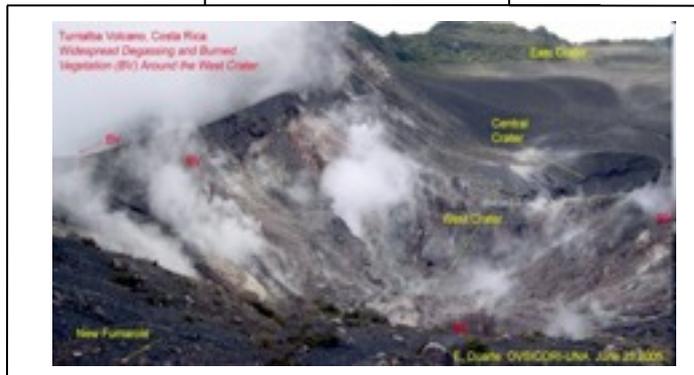
# The SO<sub>2</sub> Pyramid



Upper Troposphere  
Lower Stratosphere

-0.5 W/m<sup>2</sup> (± 2)

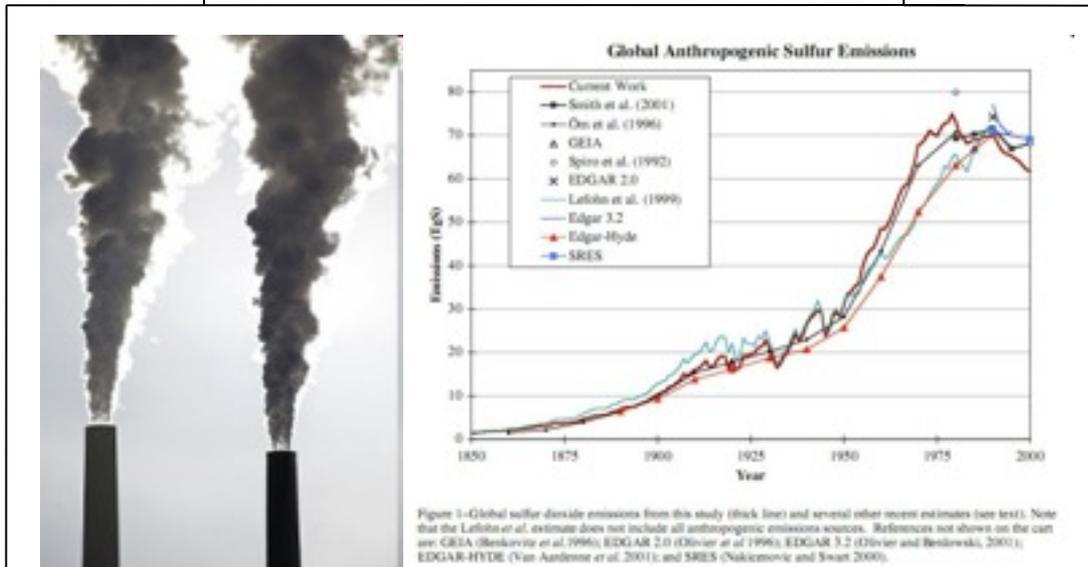
1Tg ± a lot



Free troposphere:

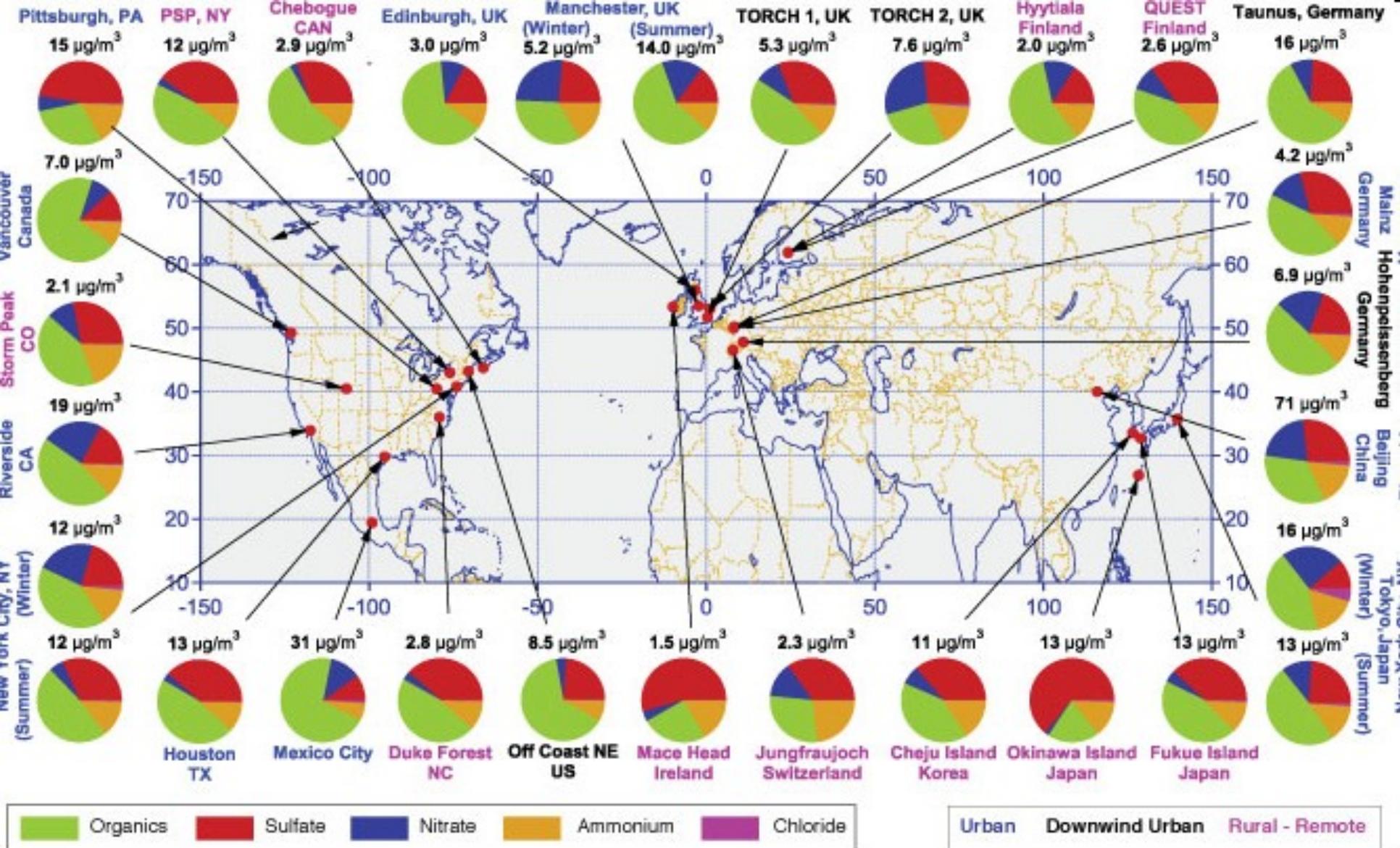
-0.3 W/m<sup>2</sup>

10 ± 5 Tg

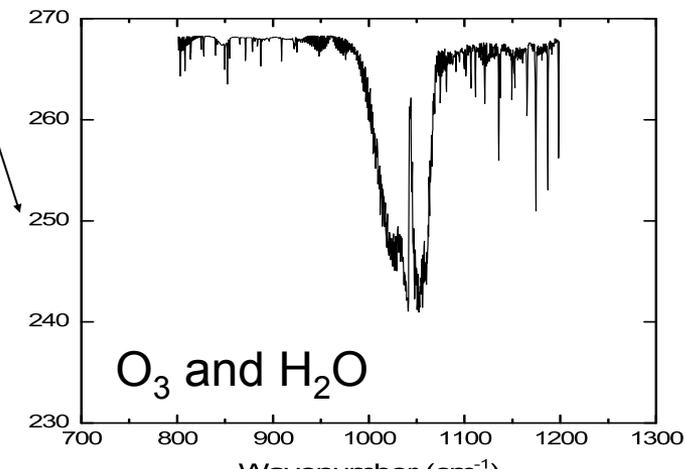
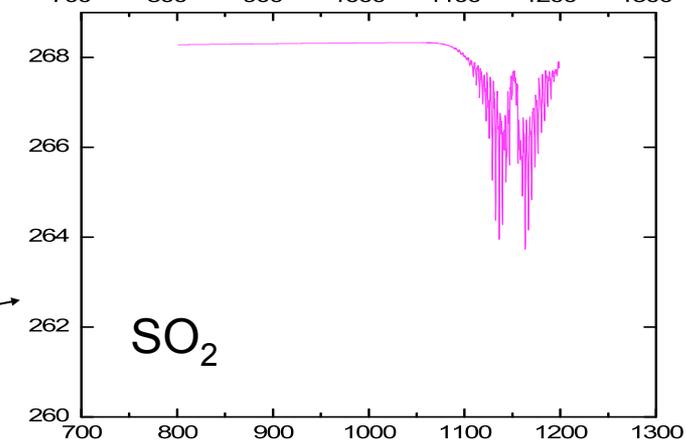
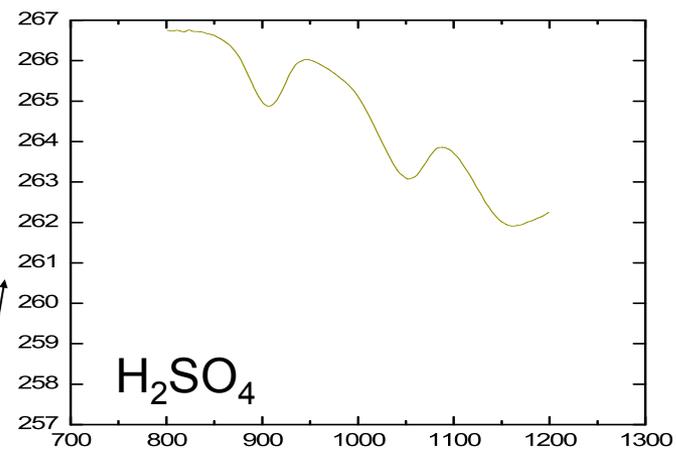
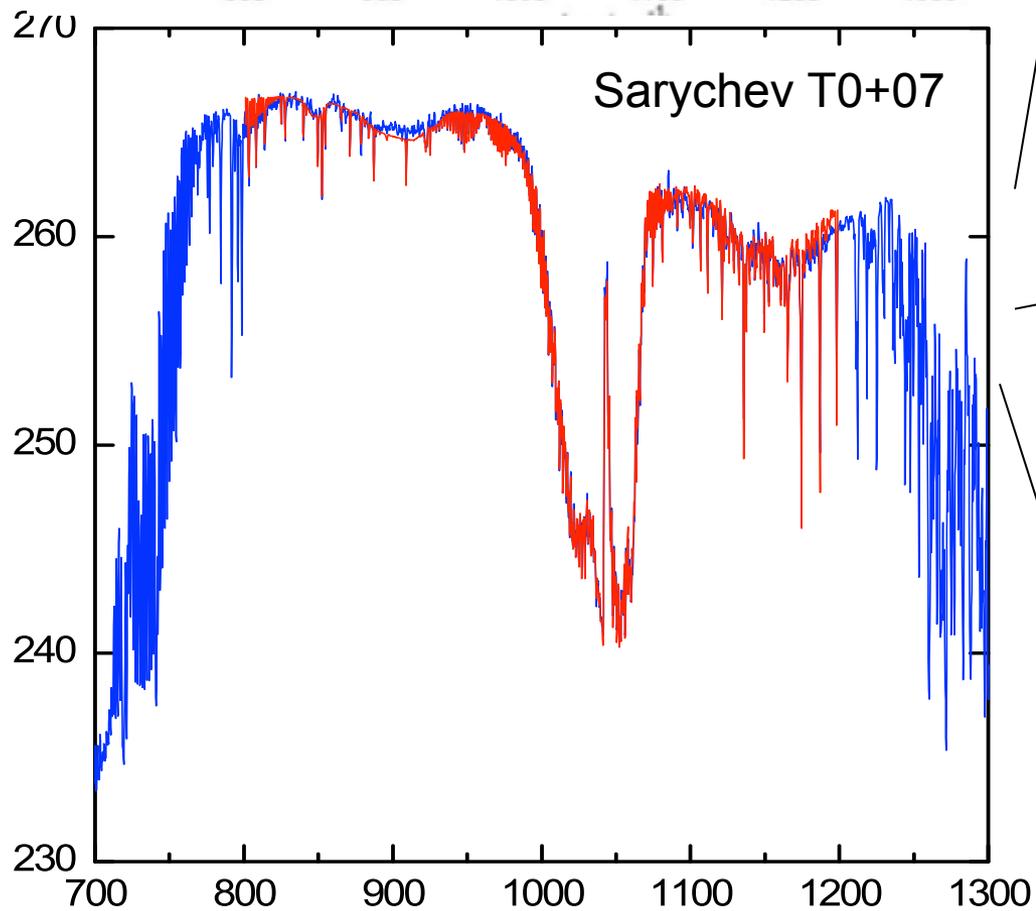
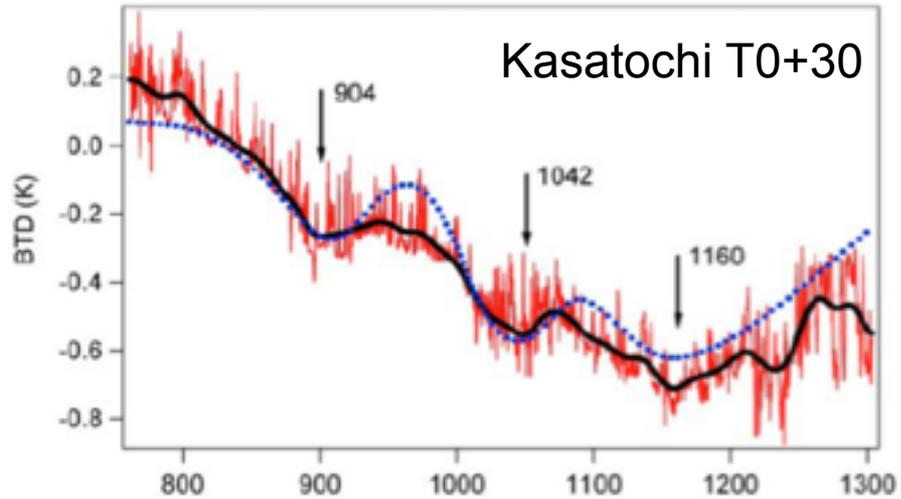


Boundary Layer: -0.2 W/m<sup>2</sup>

70 ± 10 Tg

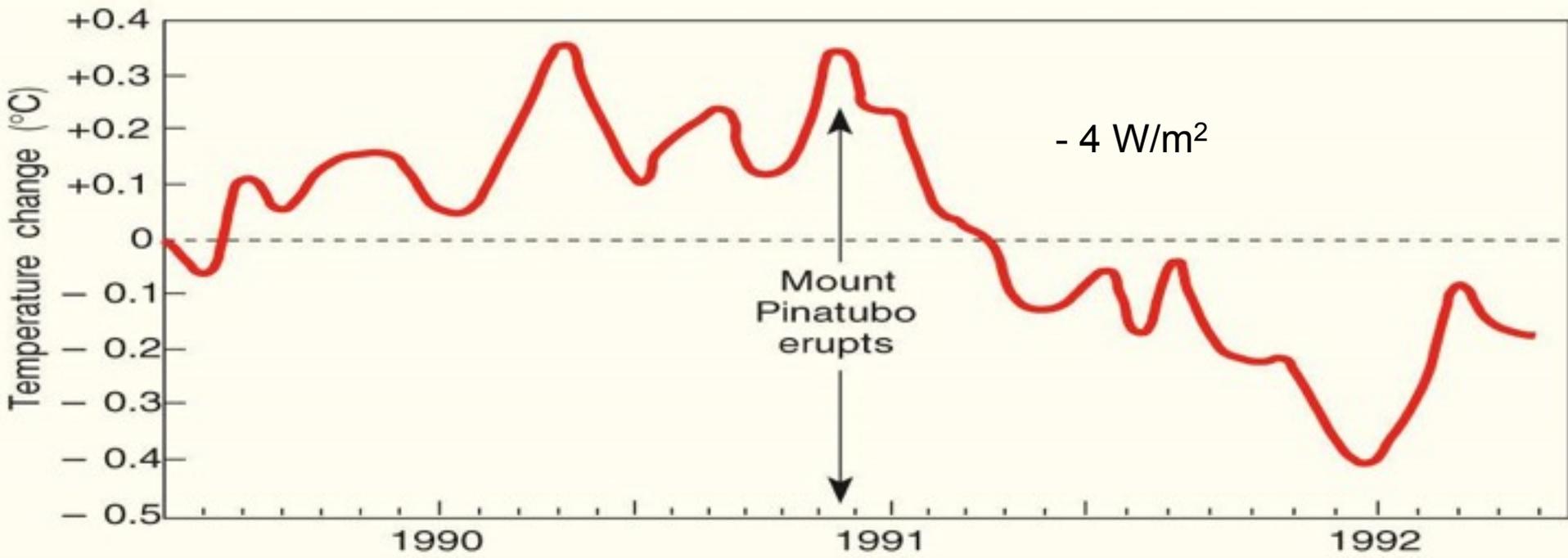
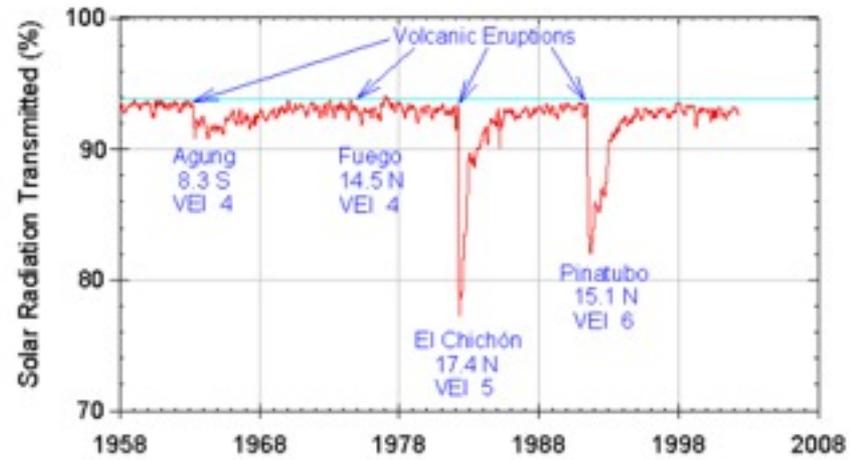


Sulphuric acid  
 Ammonium sulphate



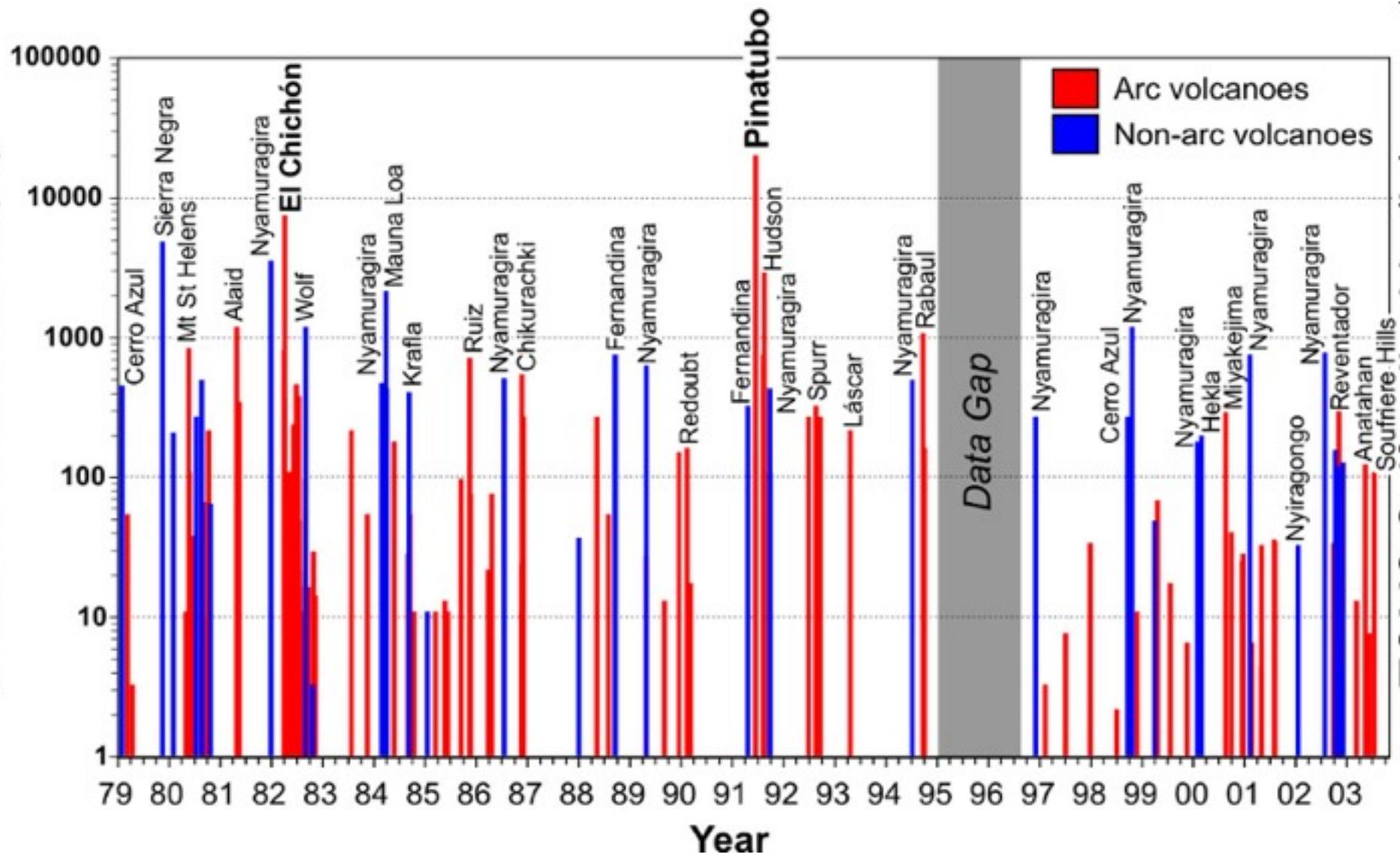


Mauna Loa Observatory Atmospheric Transmission



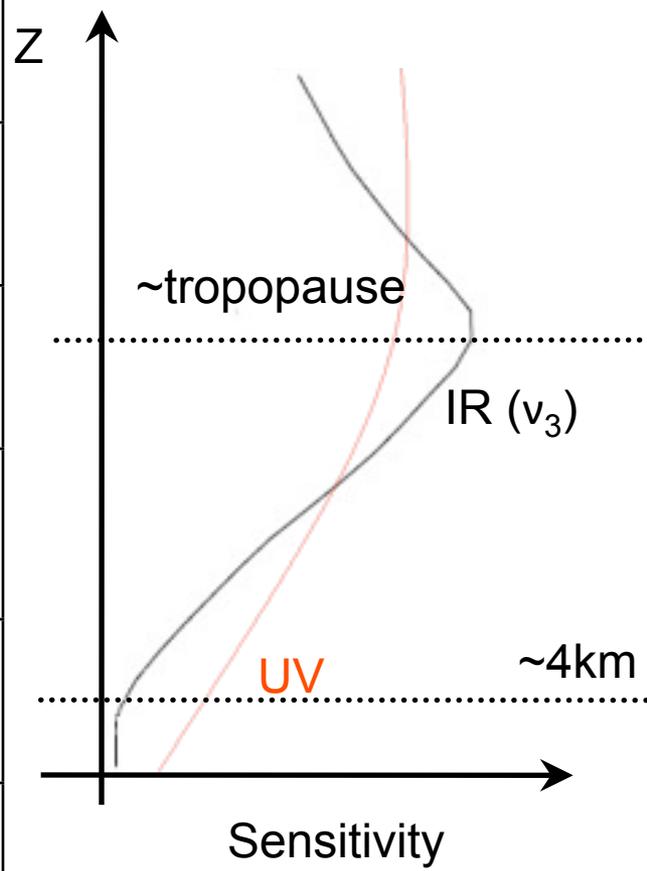
# TOMS Sulfur dioxide (kt) record

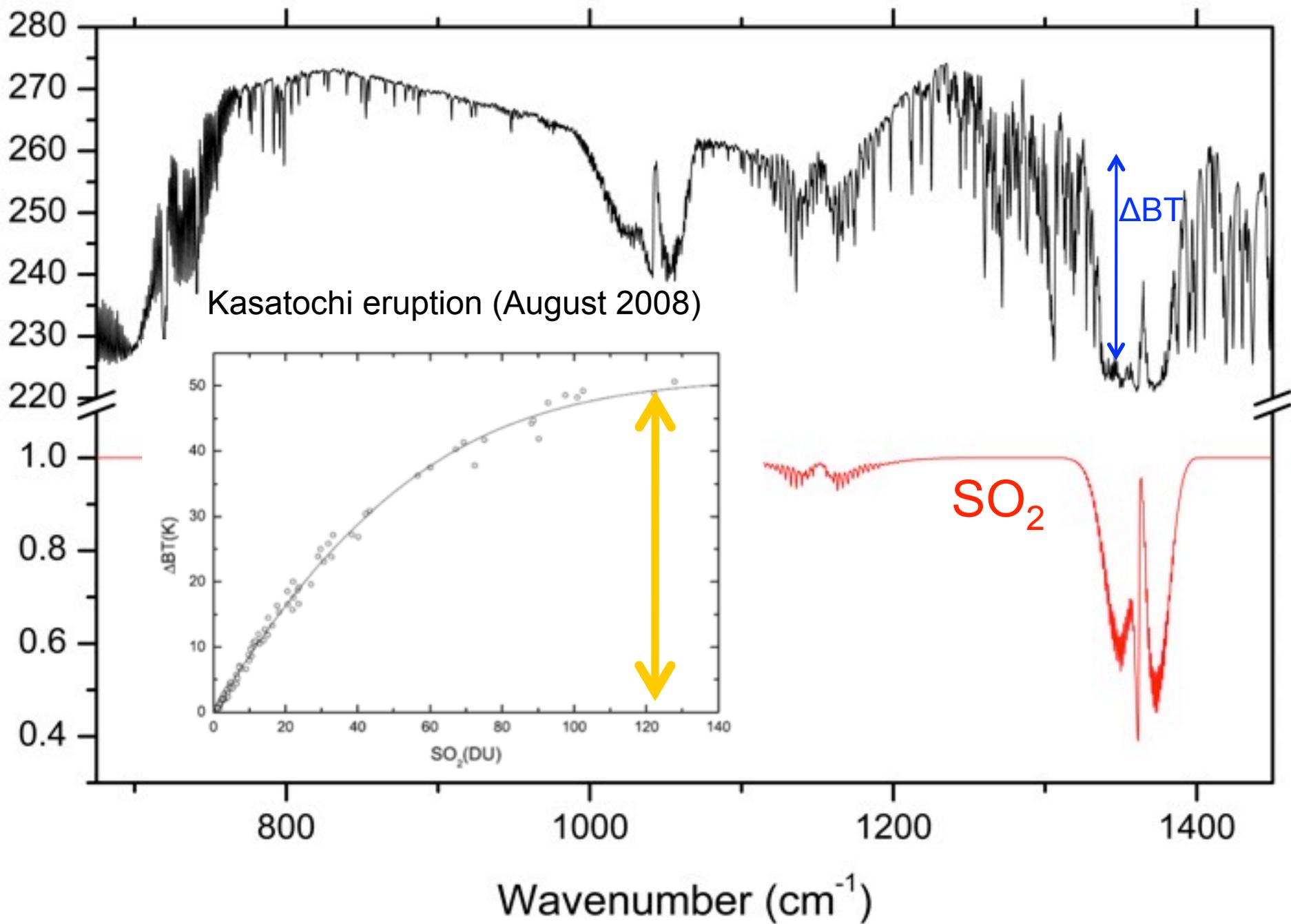
TOMS Sulfur dioxide (kt)



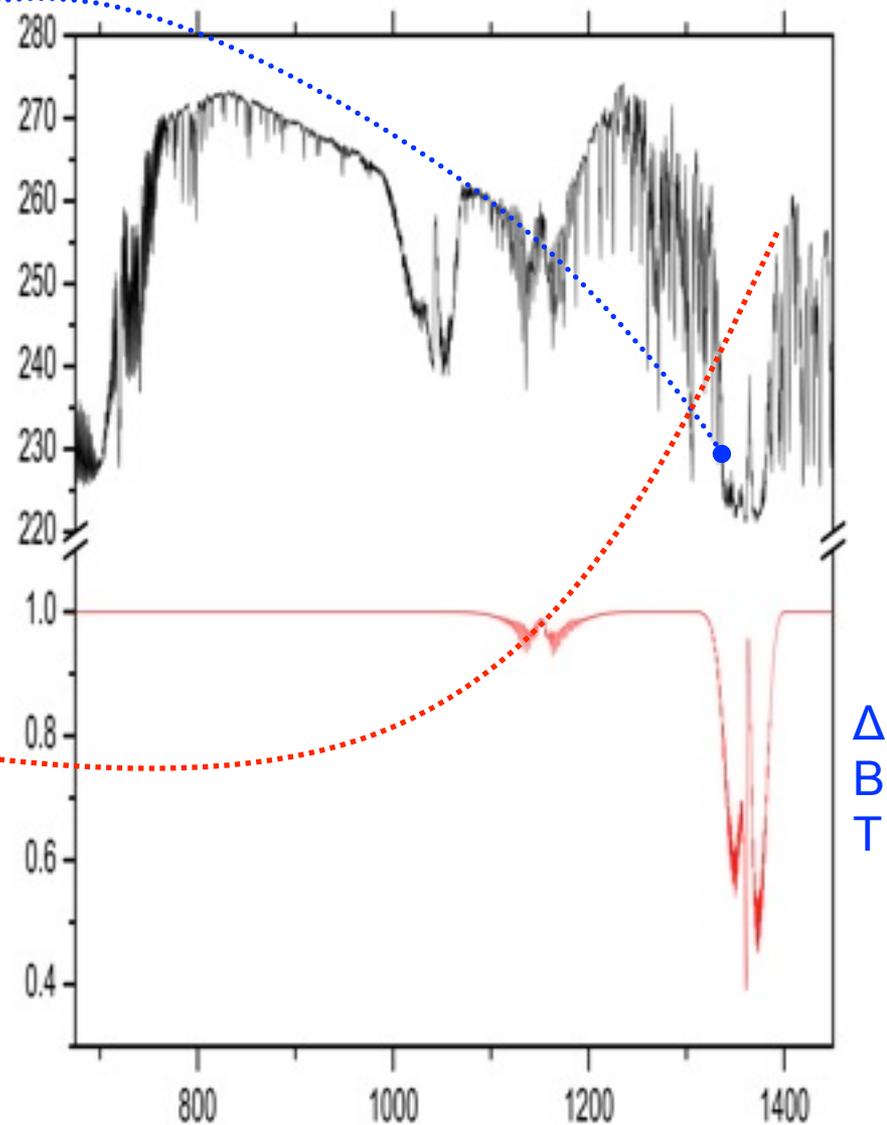
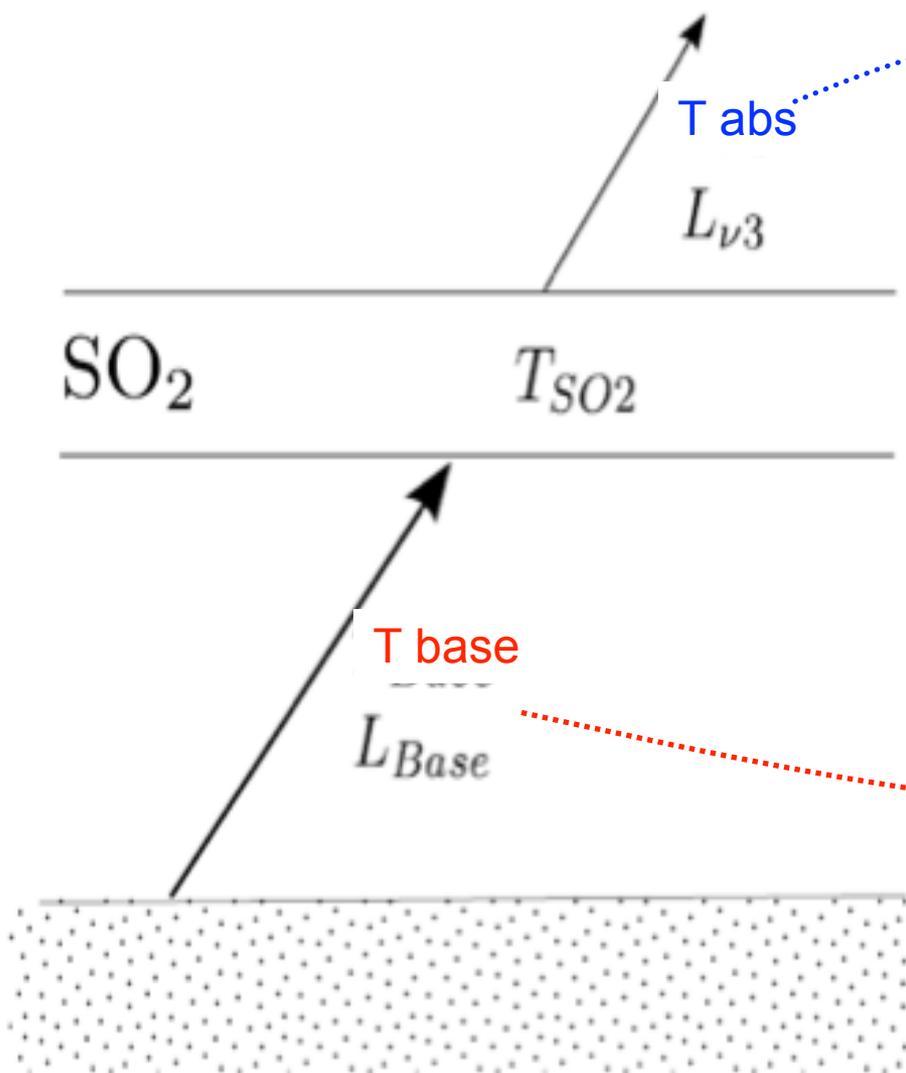
5DU detection limit / 10-30 % error !?

	TOMS UV	GOME2 UV	OMI UV	AIRS IR	IASI IR
Footprint	39x39	80x40	13x24	13x13	12x12
Channels	7	4096	780	2378	8460
Spectral resolution	1.1nm	0.3nm	0.45nm	1cm <sup>-1</sup>	0.5cm <sup>-1</sup>
Detection Limit – DU	5	0.25	0.25	0.25	0.25
Coverage	1	1	1	2	2





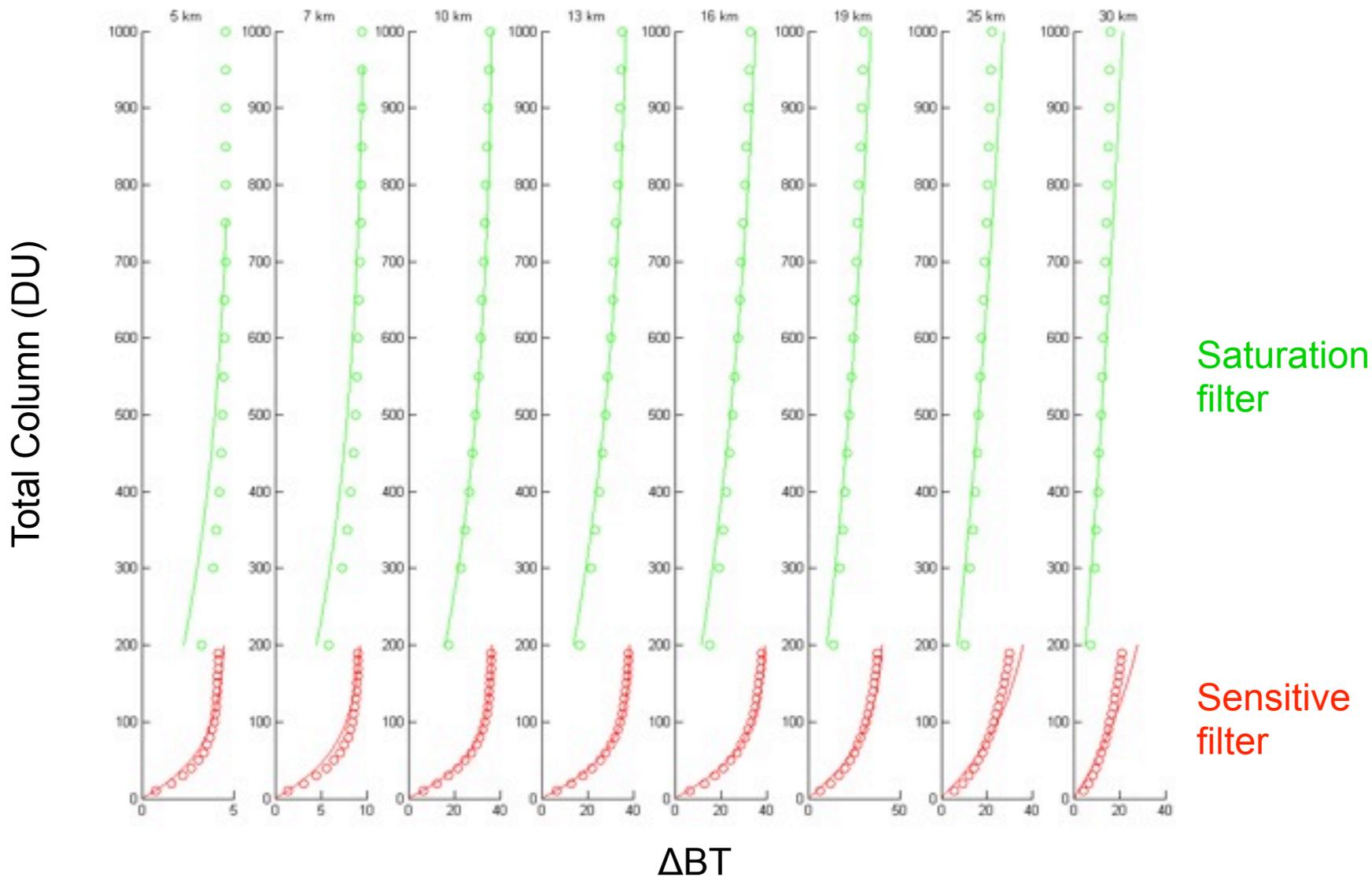
# Rapid SO<sub>2</sub> Retrieval Methodology



$$C = -\frac{\cos \theta}{c_1} \ln \left( \frac{(\exp(A/T_B) - 1)(\exp(A/T_{SO2}) - \exp(A/T_{\nu3}))}{(\exp(A/T_{\nu3}) - 1)(\exp(A/T_{SO2}) - \exp(A/T_B))} \right)$$

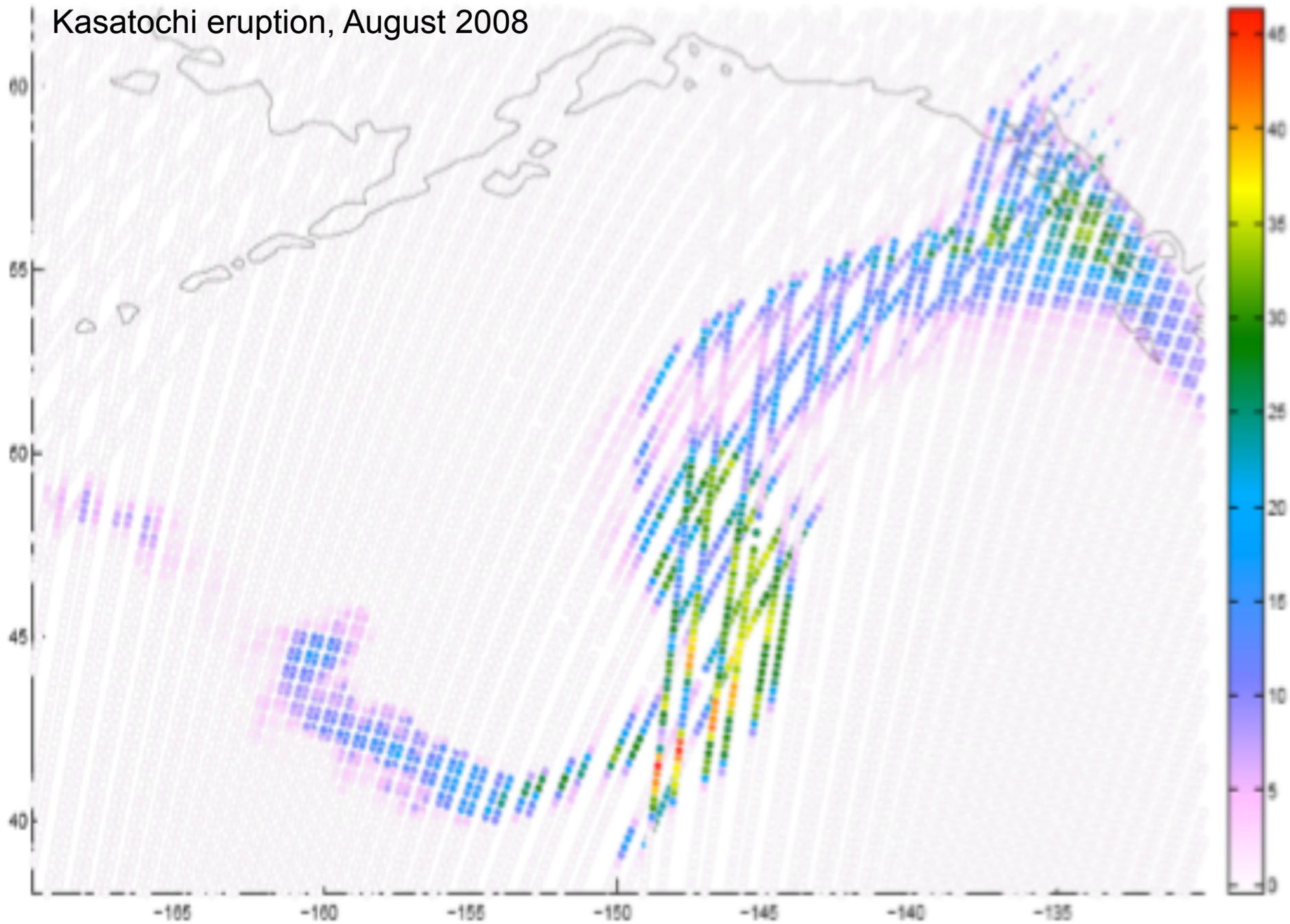
Wavenumber (cm<sup>-1</sup>)

# Rapid SO<sub>2</sub> Retrieval Methodology

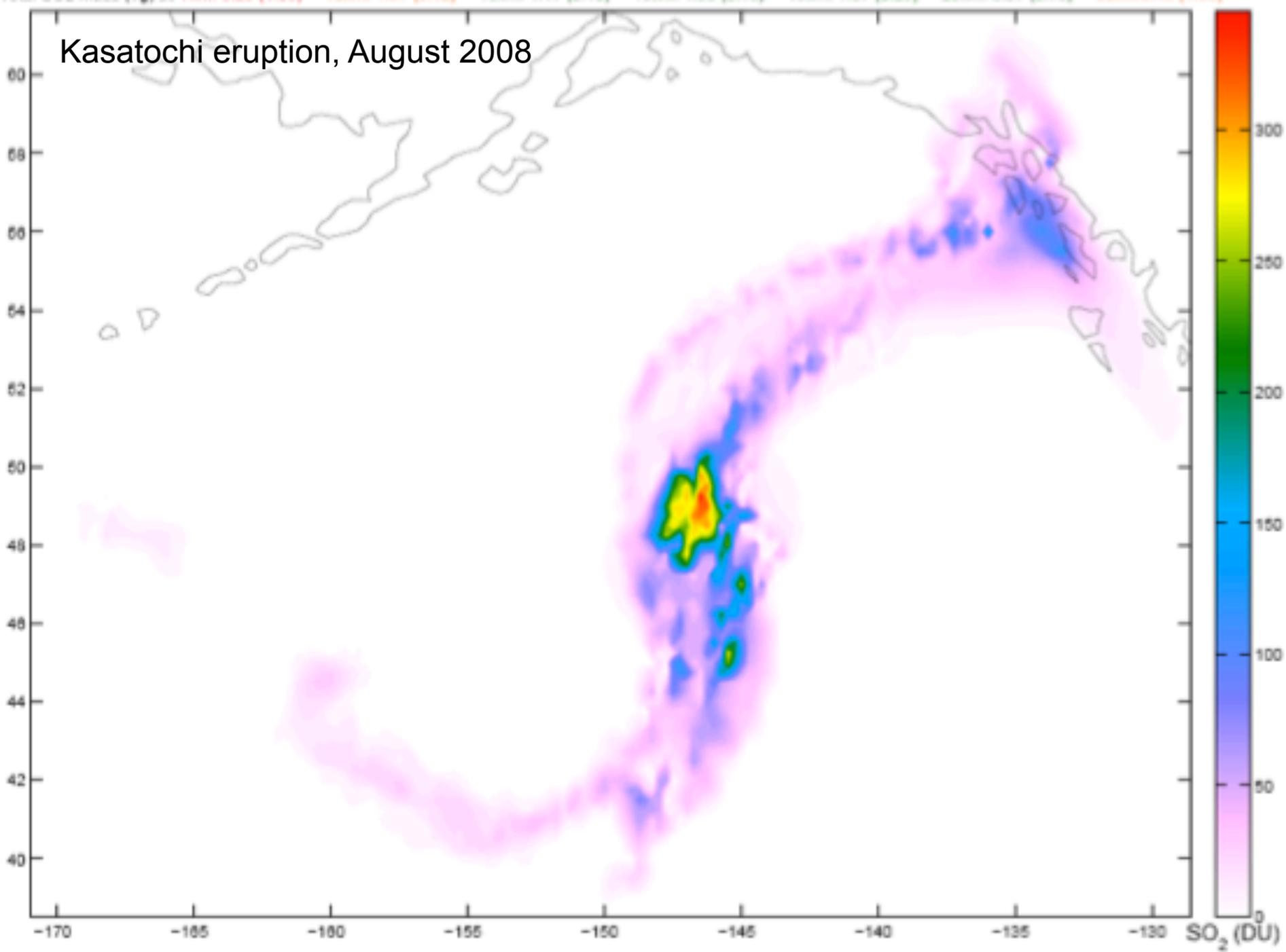


6% theoretical accuracy !

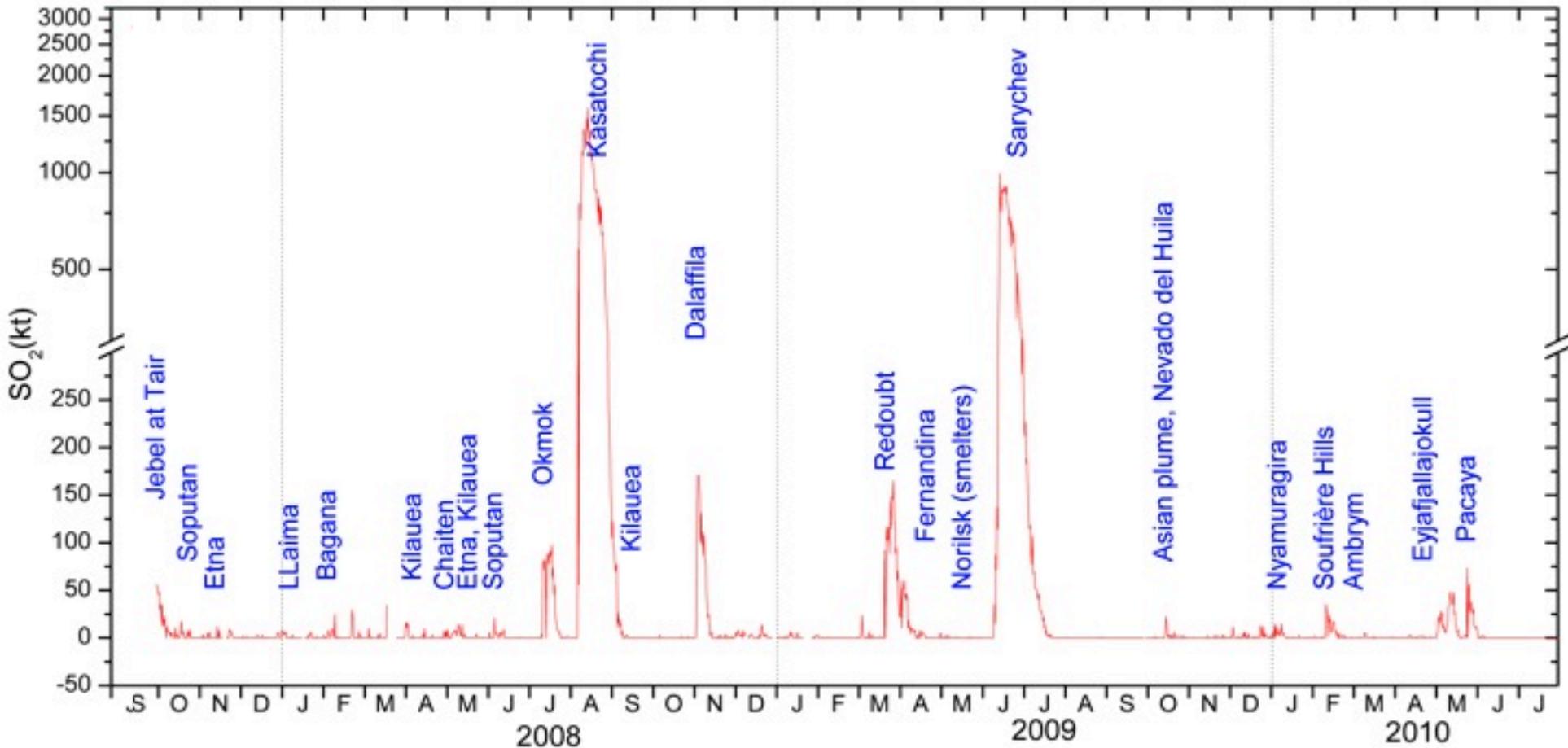
# Kasatochi eruption, August 2008



Kasatochi eruption, August 2008

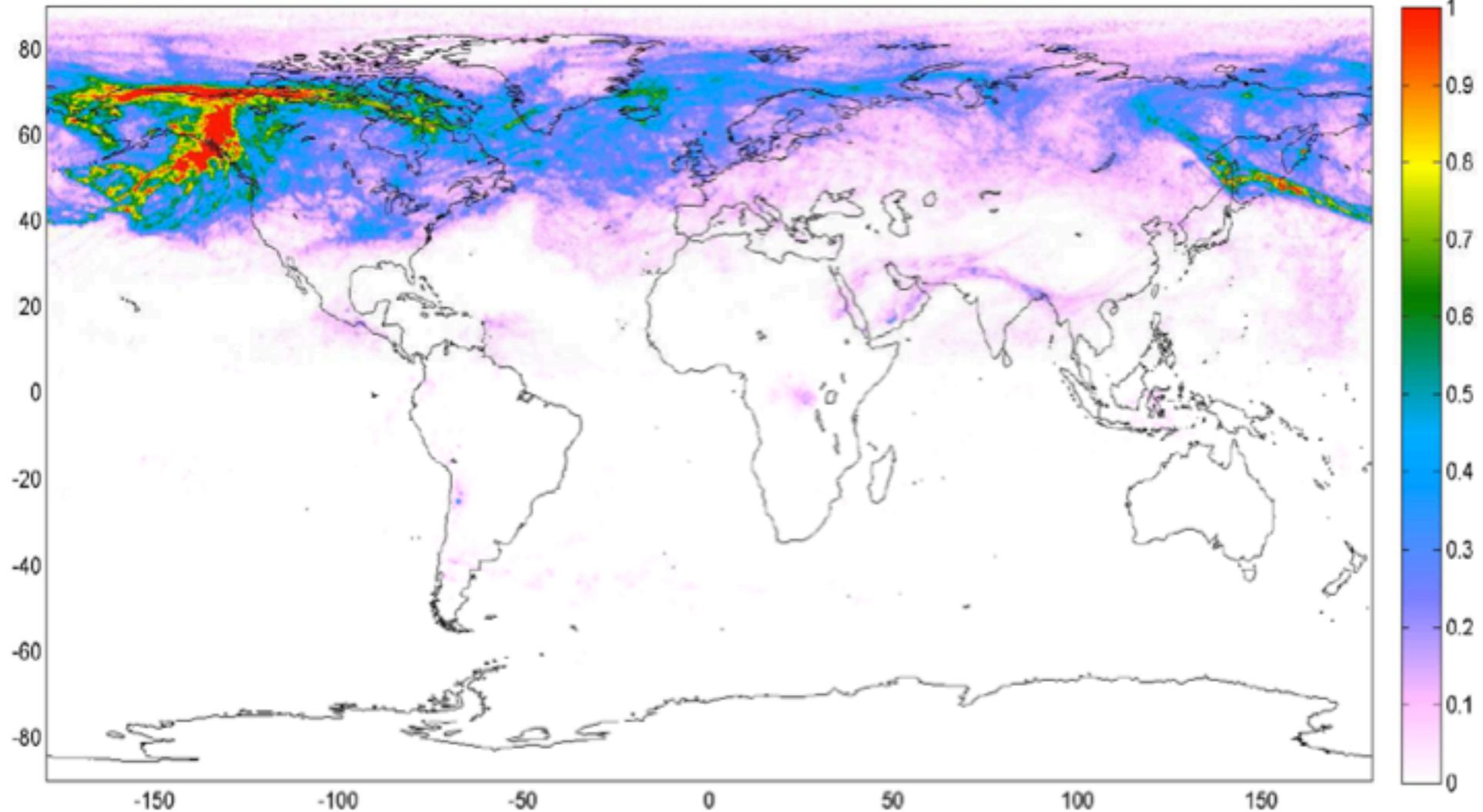


# 3 year timeseries of SO<sub>2</sub> from IASI



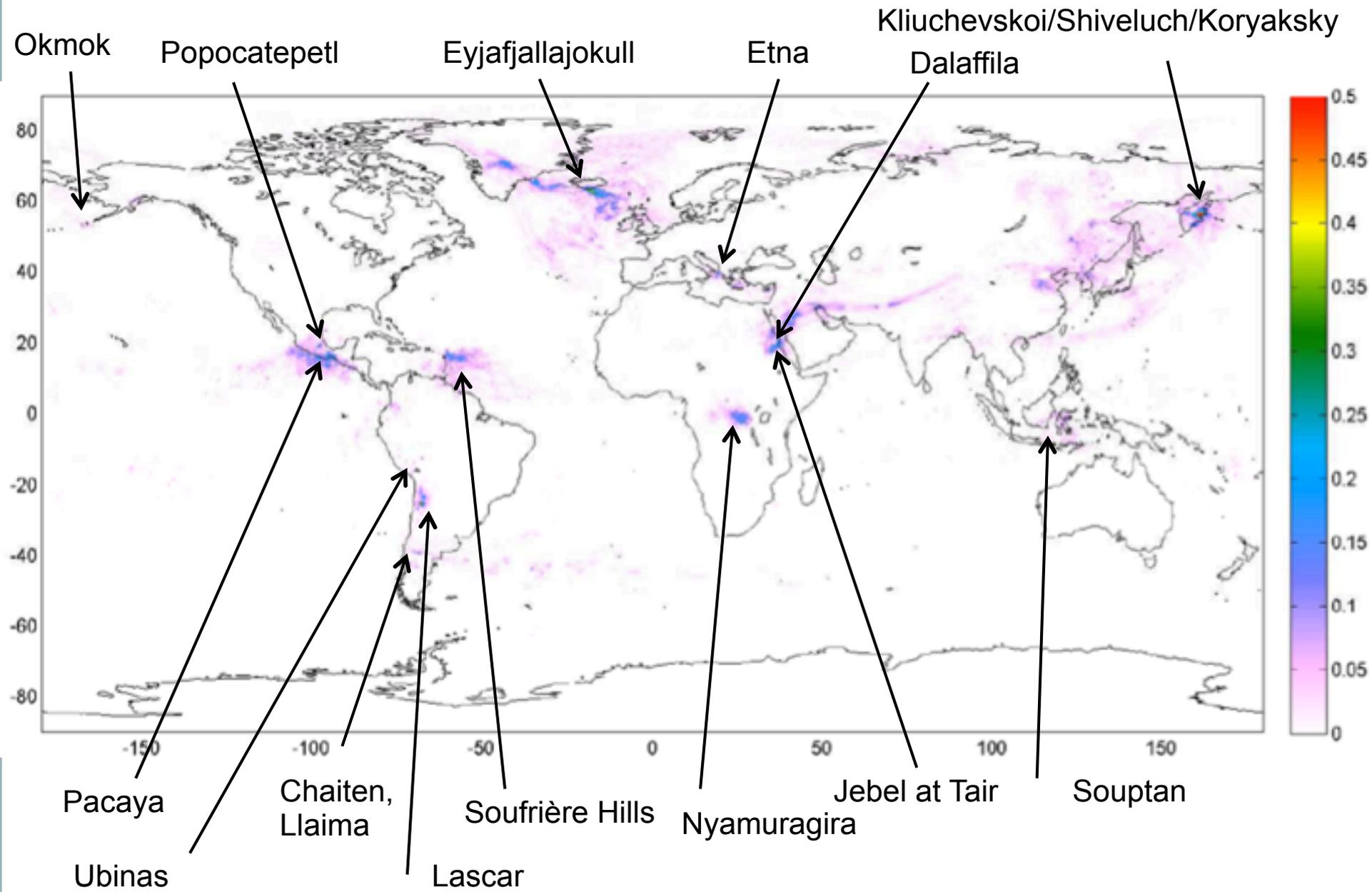
20 % error wrt optimal estimation!!

# 3 year global UTLS SO<sub>2</sub> from IASI

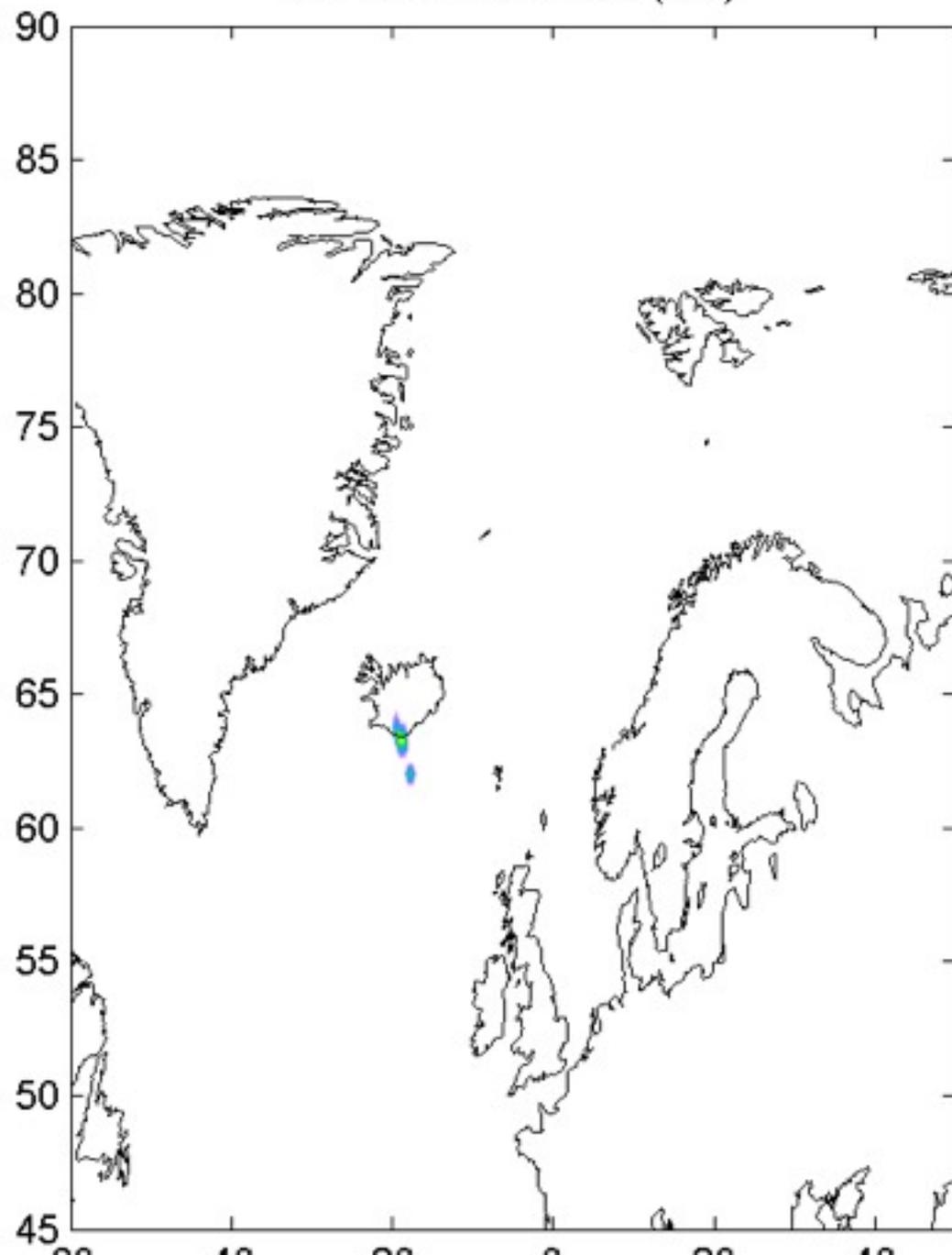


Up to 1 DU on average!

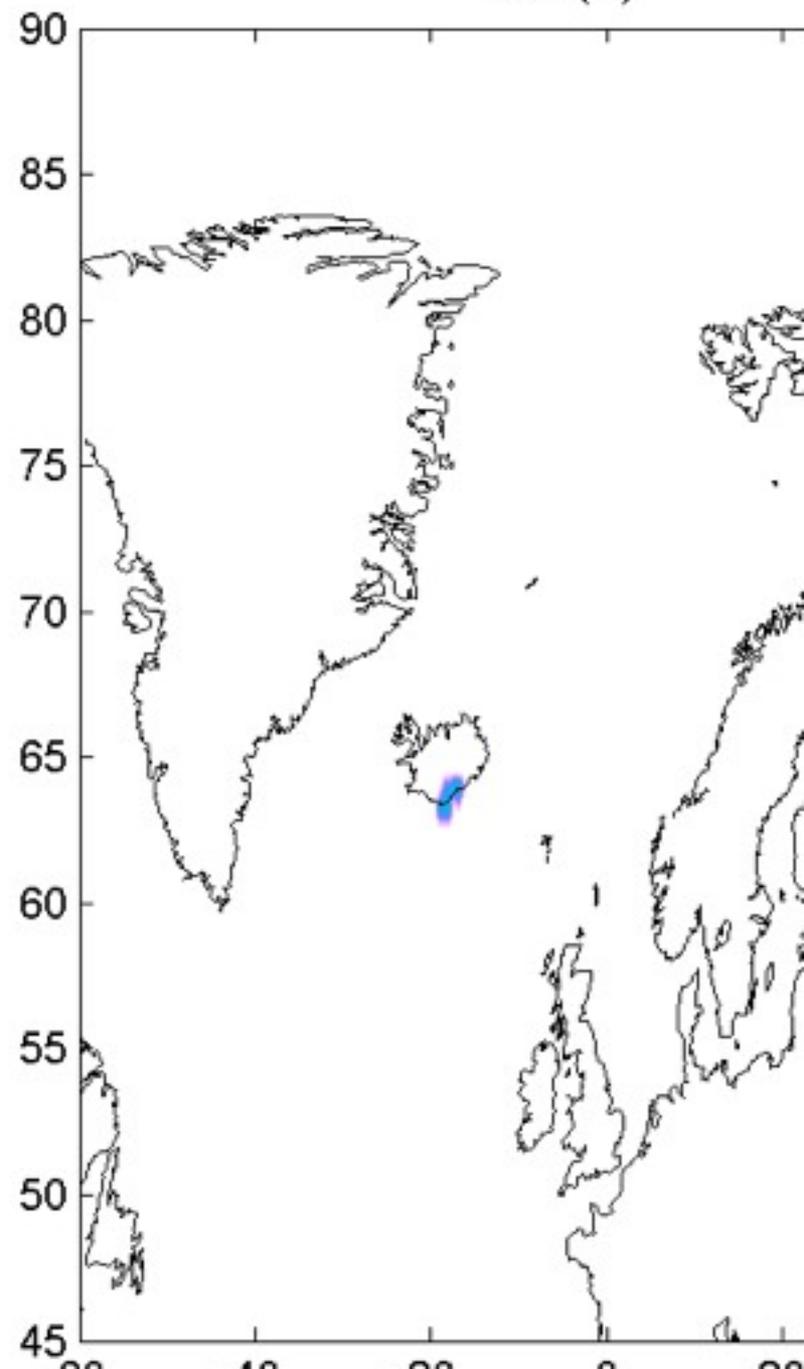
# 3 year global average SO<sub>2</sub> from IASI (without large eruptions)



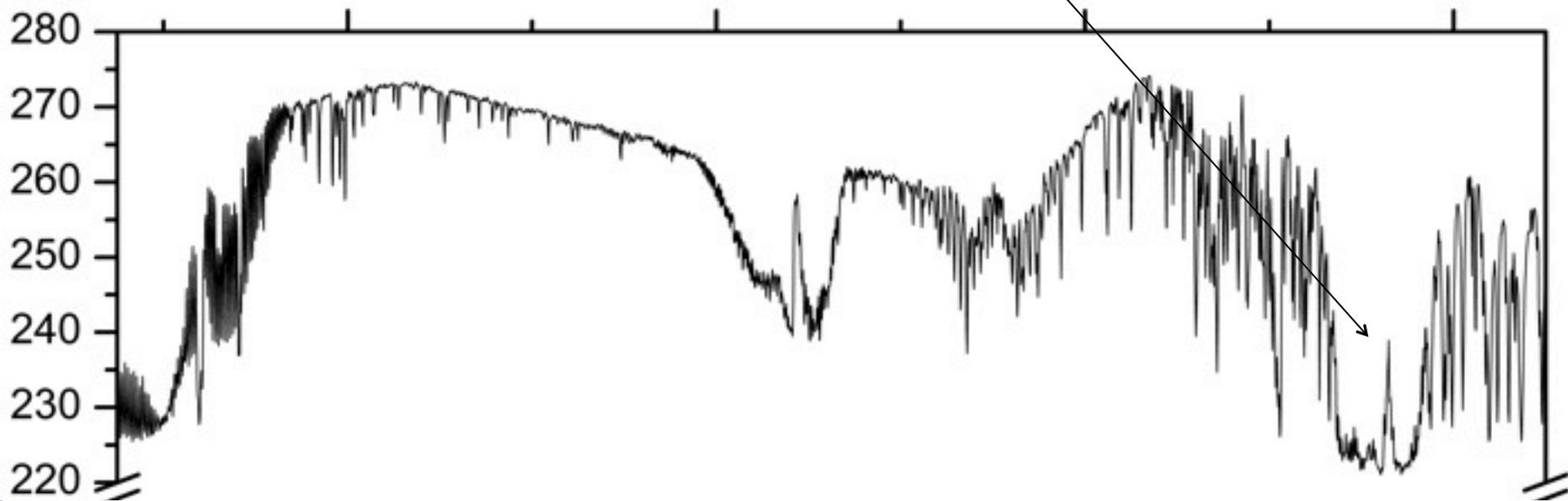
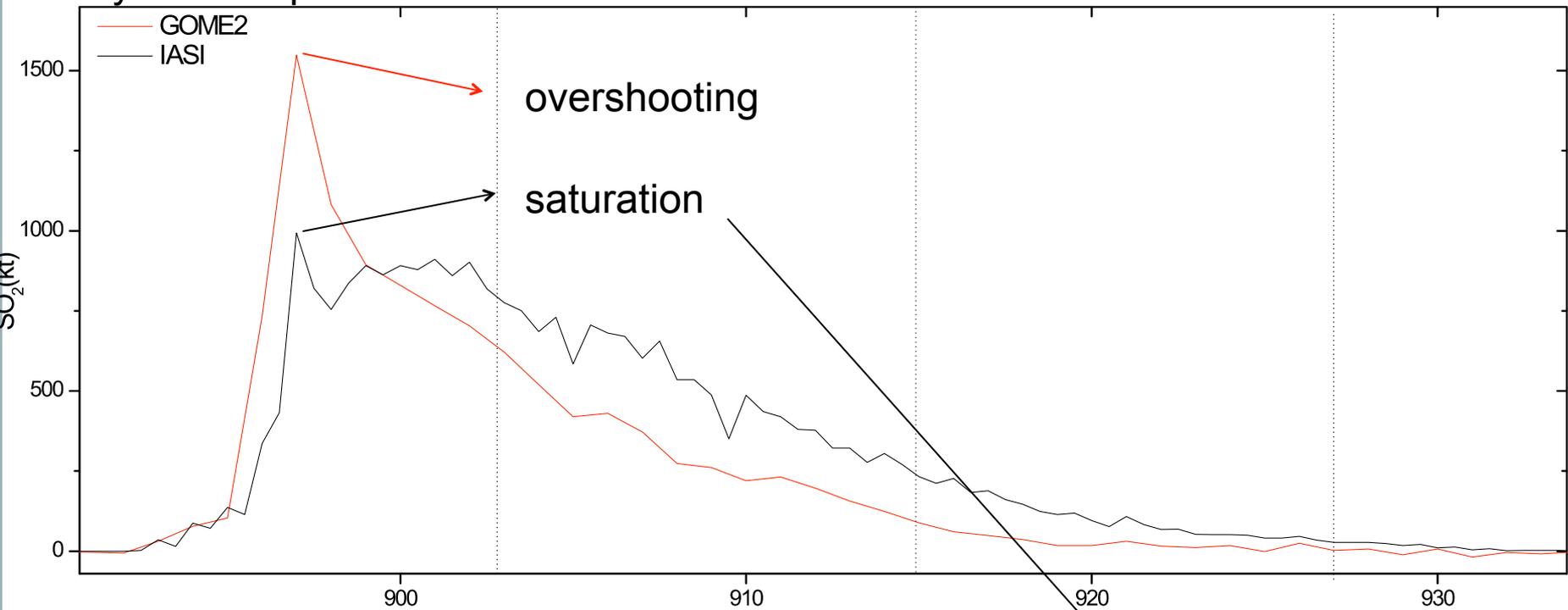
20100401AM - SO2 (DU)



Ash (K)



# Sarychev eruption



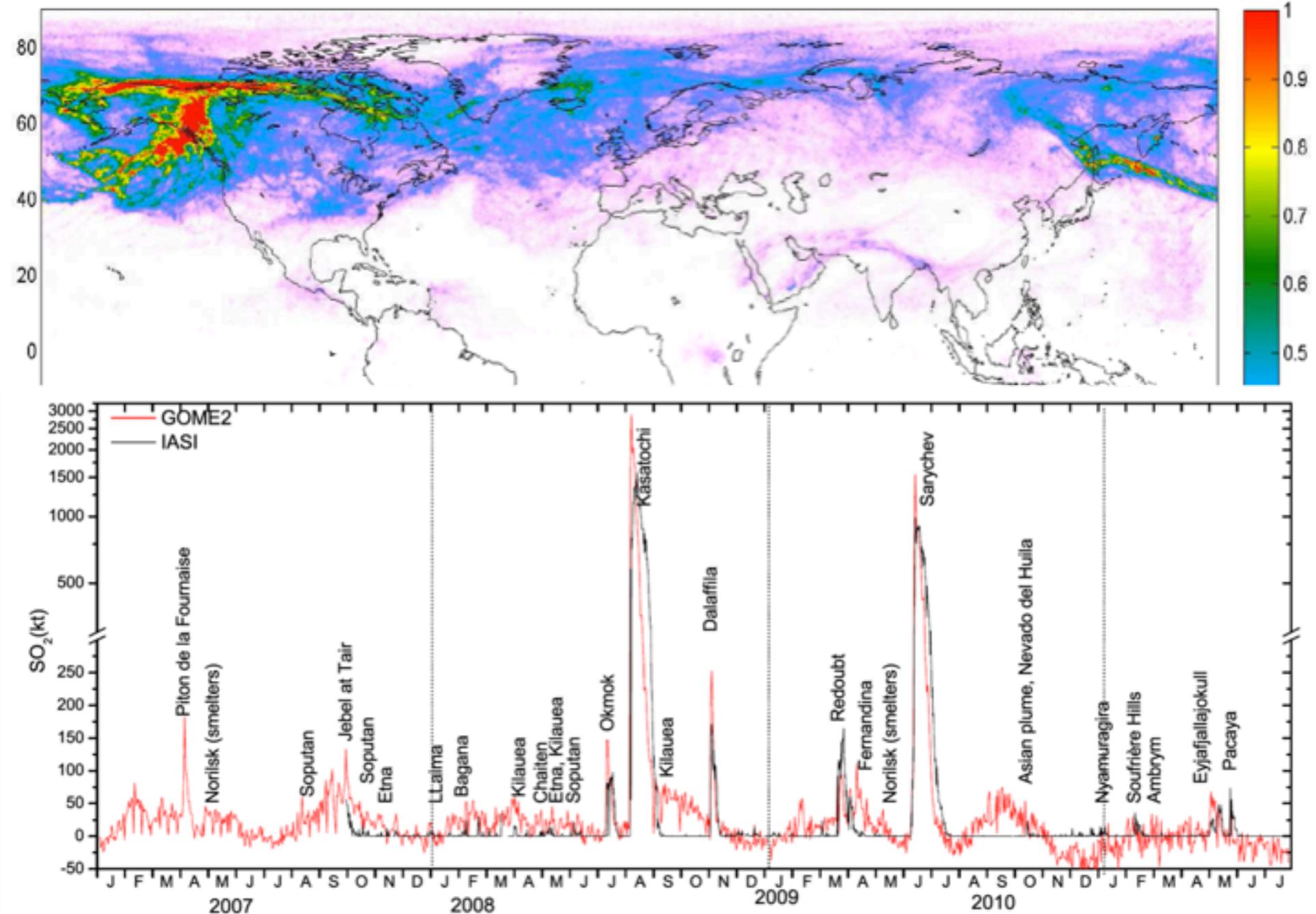


Figure 1: Timeseries 2007-2010 of SO<sub>2</sub> (kt) at 12km from IASI and GOME2.

# Measuring volcanic SO<sub>2</sub>

Explosive volcanism - effects on climate and life

Anthropogenic emissions - decreasing in the developed world but still rising in China/India/South Africa

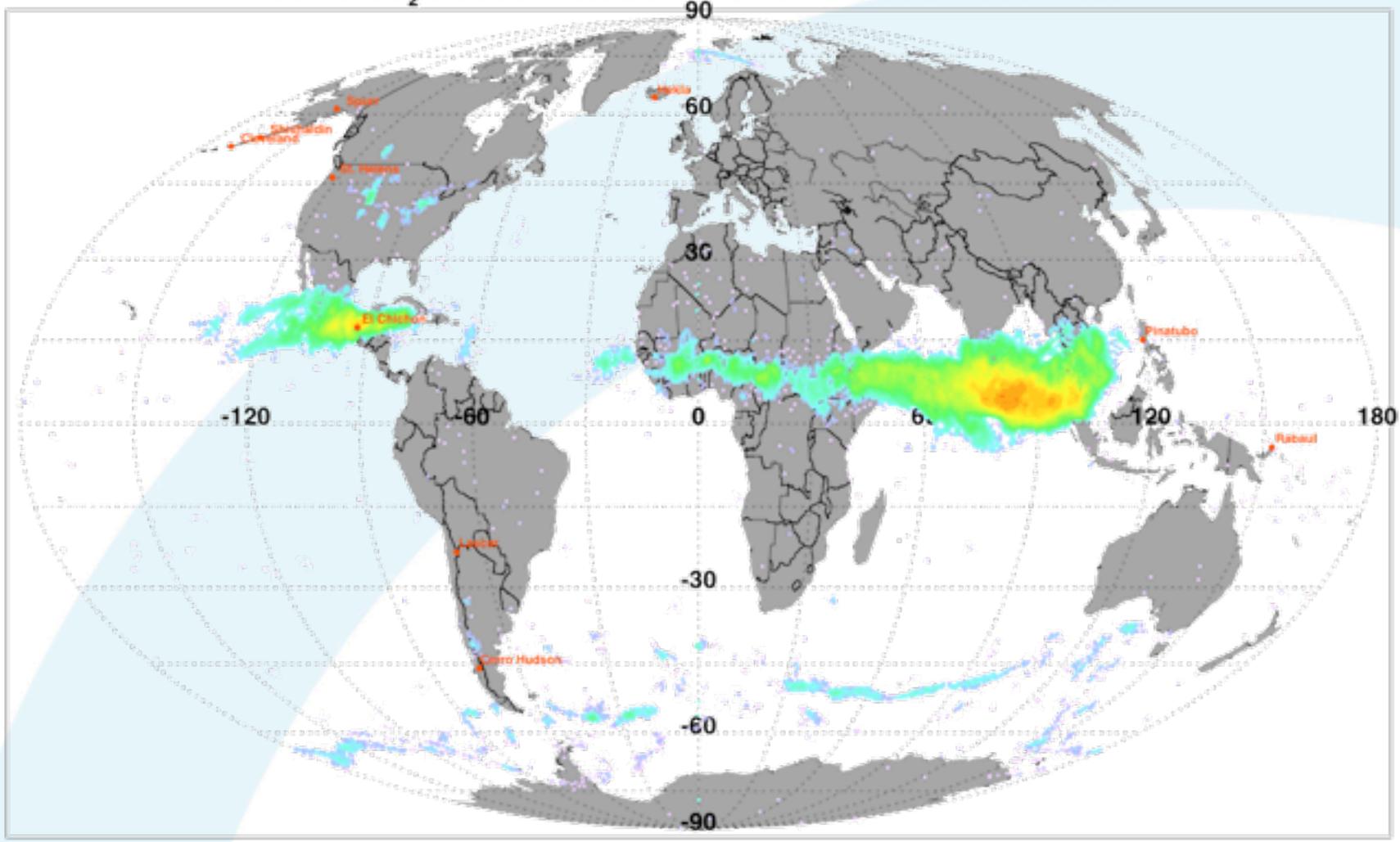
Geo-engineering (controversial suggestion to inject S into stratosphere)

Aviation hazard (?)

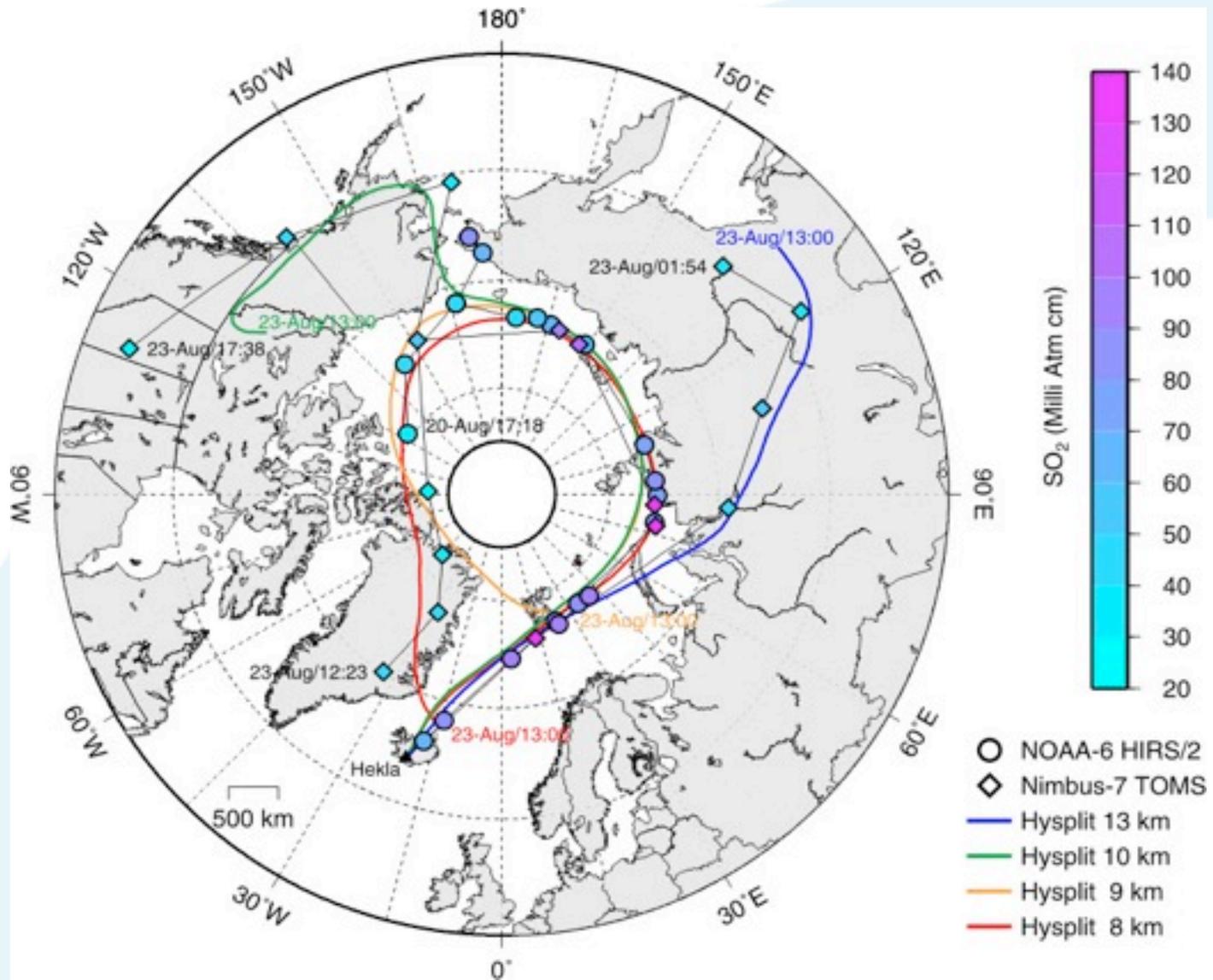
# SO<sub>2</sub> measurements from satellites

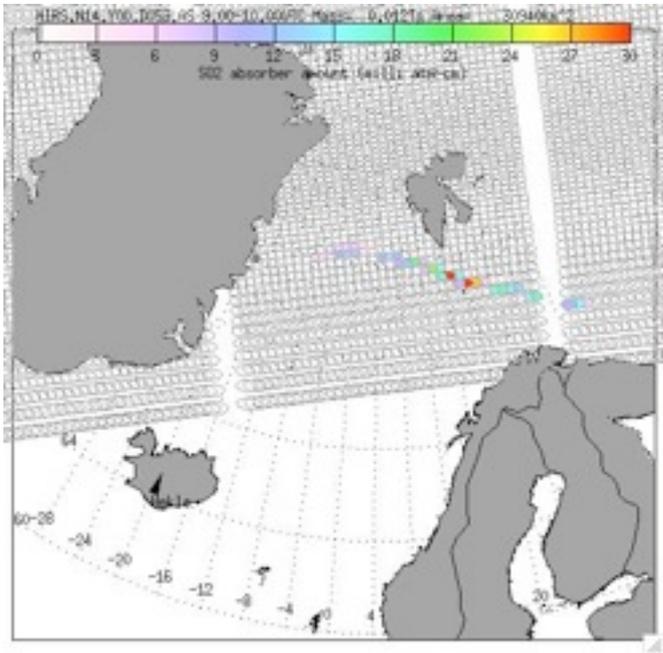
<b>INSTRUMENT</b>	<b>From</b>
TOMS	1979
TOVS	1979
MLS	1991
MODIS	1999
ASTER	1999
<i>ACE</i>	<i>2003</i>

# Some SO<sub>2</sub> clouds detected in TOVS/HIRS data 1980 - 2000



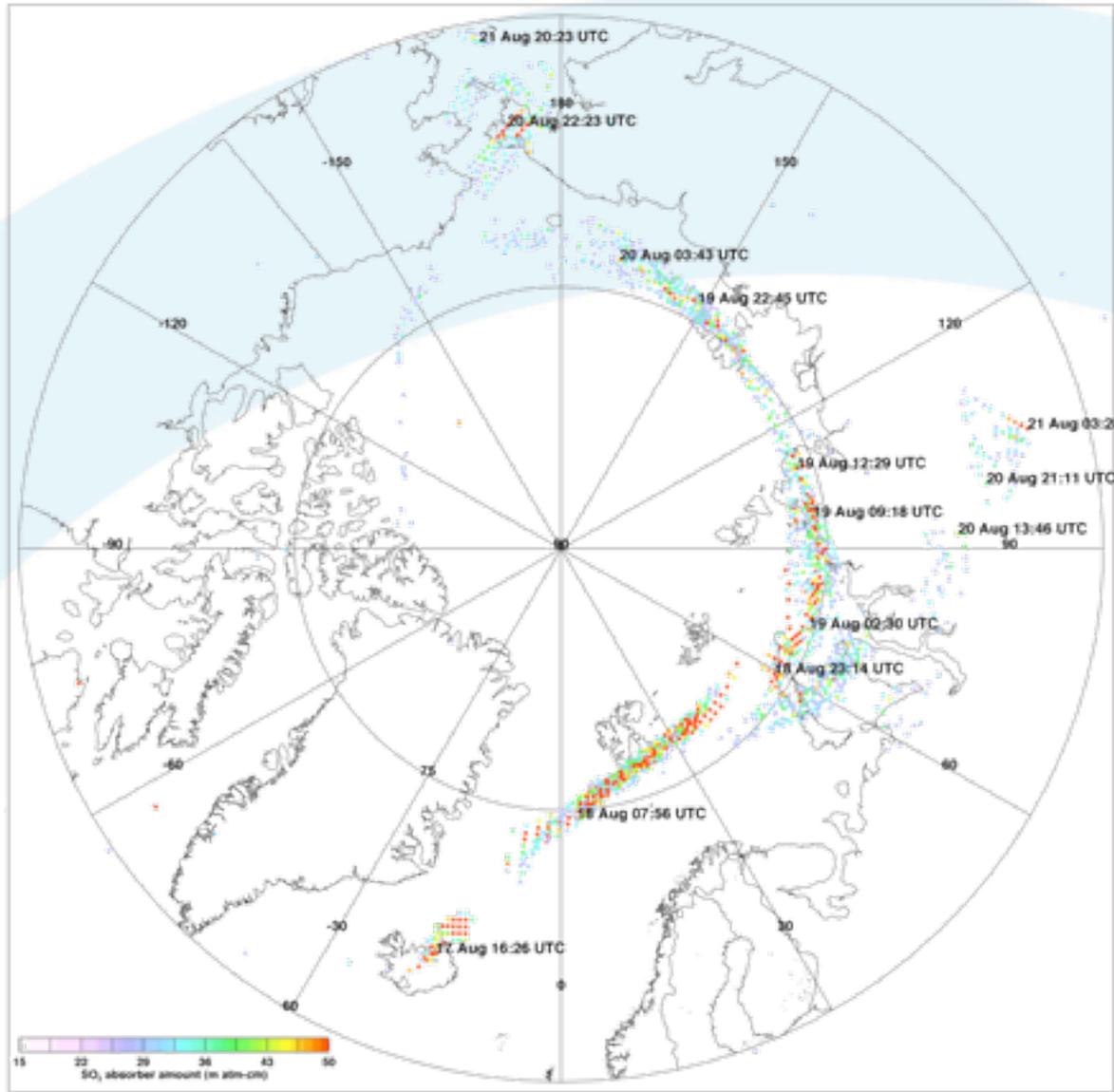
Carn, S., Prata, A. and S. Karlsdottir, 2008, Circumpolar transport of a volcanic cloud from Hekla (Iceland). *J. Geophys. Res.*, 113, D14311, doi:10.1029/2008JD009878.





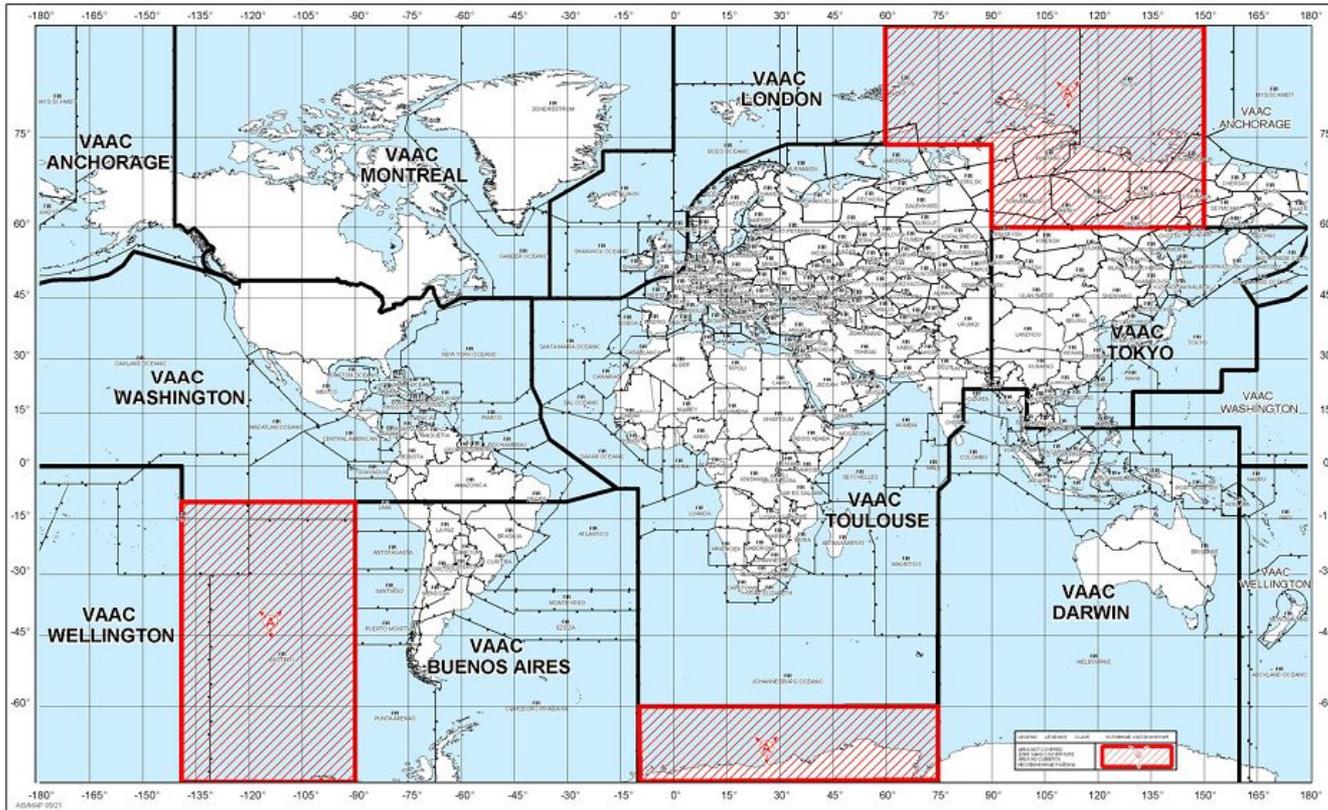
TOVS/ATOVS

*S-Watch* implementation at the Bureau of Met



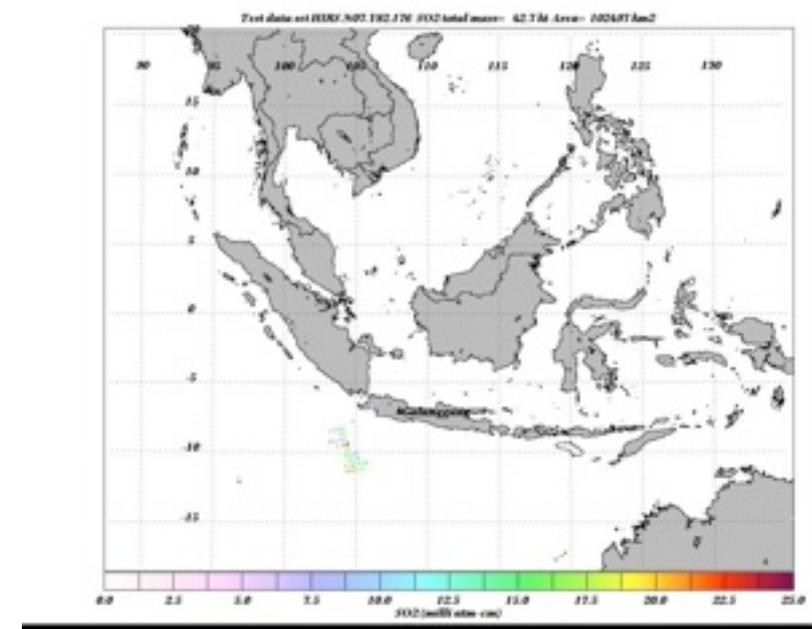
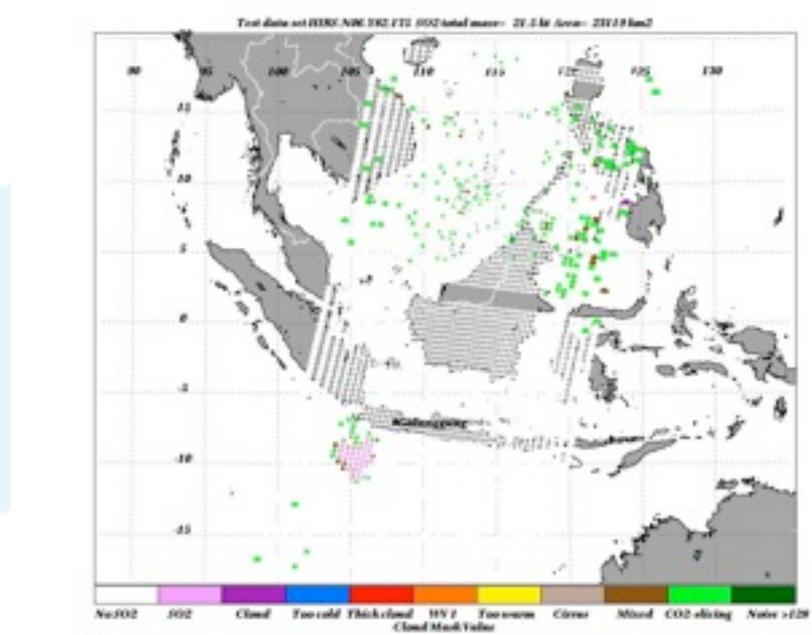
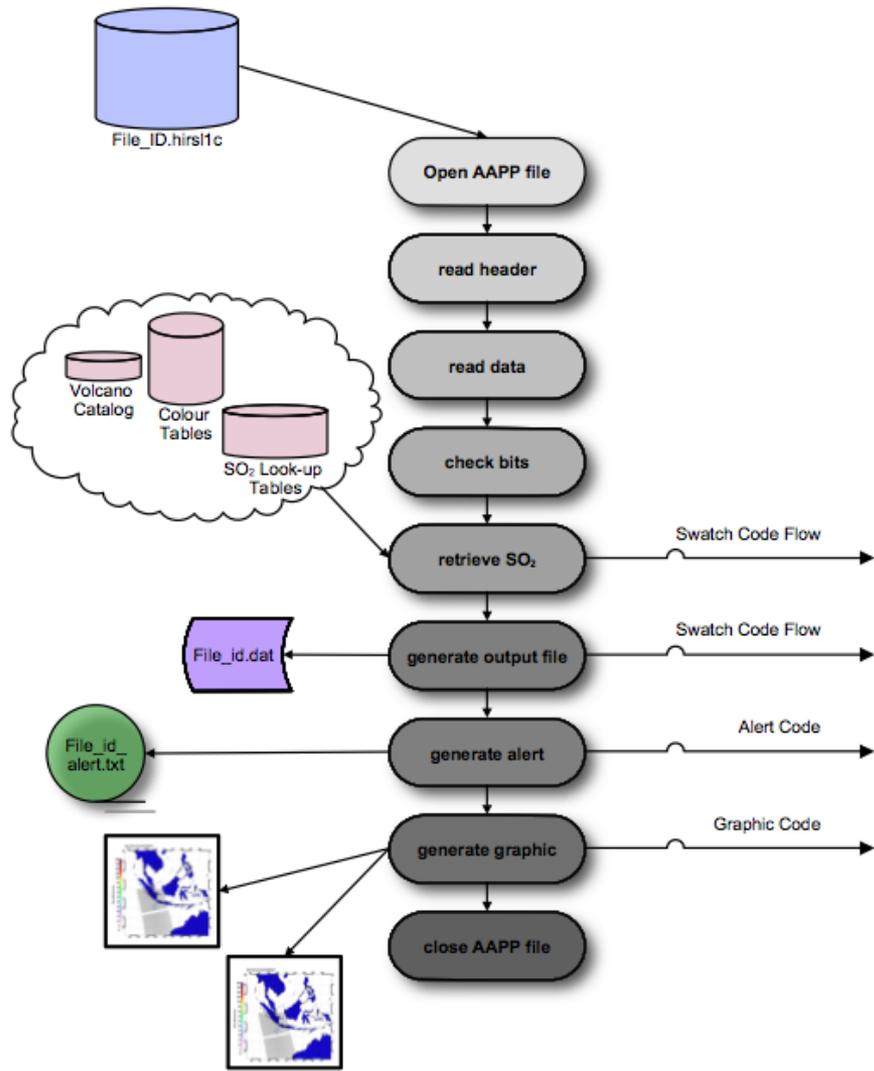
# VAAC Region - Swatch graphical coverage

CURRENT STATUS OF ICAO VOLCANIC ASH ADVISORY CENTRES (VAAC) - AREAS OF RESPONSIBILITY  
SITUATION ACTUELLE DES CENTRES OACI D'AVIS DE CENDRES VOLCANIQUES (VAAC) - ZONES DE RESPONSABILITÉ  
ESTADO ACTUAL DE LOS CENTROS DE AVISOS DE CENIZAS VOLCÁNICAS (VAAC) DE LA OACI - ÁREAS DE RESPONSABILIDAD  
СУЩЕСТВУЮЩЕЕ РАСПРЕДЕЛЕНИЕ КОНСУЛЬТАТИВНЫХ ЦЕНТРОВ ИКАО ПО ВУЛКАНИЧЕСКОМУ ПЕЛУ (ВААС) - РАЙОНЫ ОТВЕТСТВЕННОСТИ



VAAC1=[88,135,-18,20]  
VAAC2=[125,170,-35,20]  
VAACm=[88,170,-35,20]

# Swatch Processor



# I/O: The Data Output File

Label	Meaning
Lon	Longitude (decimal degrees)
Lat	Latitude (decimal degrees)
T67	BT 6.7 $\mu\text{m}$ (Kelvin)
T73	BT 7.3 $\mu\text{m}$ (Kelvin)
T83	BT 8.3 $\mu\text{m}$ (kelvin)
T11	BT 11 $\mu\text{m}$ (Kelvin)
SO2	SO <sub>2</sub> column (DU)
DT	BT 7.3 $\mu\text{m}$ – BT 7.3 $\mu\text{m}$ interpolated (Kelvin)
D83-11	BT 8.3 $\mu\text{m}$ – BT 11 $\mu\text{m}$ (Kelvin)
T134	BT 13.4 $\mu\text{m}$ (Kelvin)
T137	BT 13.7 $\mu\text{m}$ (Kelvin)
T140	BT 14.0 $\mu\text{m}$ (Kelvin)
T142	BT 14.2 $\mu\text{m}$ (Kelvin)
T145	BT 14.5 $\mu\text{m}$ (Kelvin)
T147	BT 14.7 $\mu\text{m}$ (Kelvin)
Trat	CO <sub>2</sub> slicing ratio
CLD	Cloud flag value
Smass	SO <sub>2</sub> pixel mass (kt)
Area	Pixel area (km <sup>2</sup> )

## Data Output File

Naming convention:

SO2ndddyymm.tt.tttt.dat

n=VAAC region (1,2 or m)

ddd=day of year (1-366)

yyyy=Year

tt.tttt=decimal time (hrs, UTC)

# Measuring SO<sub>2</sub> from AIRS

JOURNAL OF GEOPHYSICAL RESEARCH, VOL. 112, D20204, doi:10.1029/2006JD007955, 2007

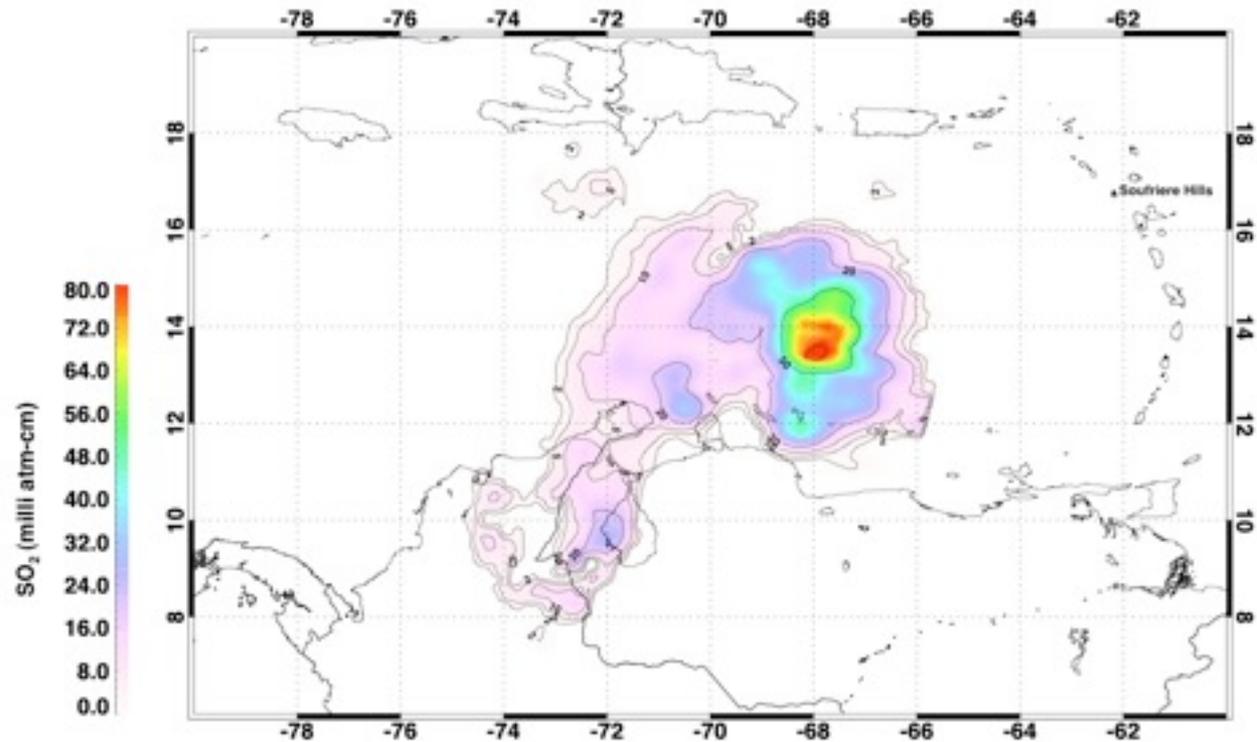


## Retrieval of volcanic SO<sub>2</sub> column abundance from Atmospheric Infrared Sounder data

A. J. Prata<sup>1</sup> and C. Bernardo<sup>2</sup>

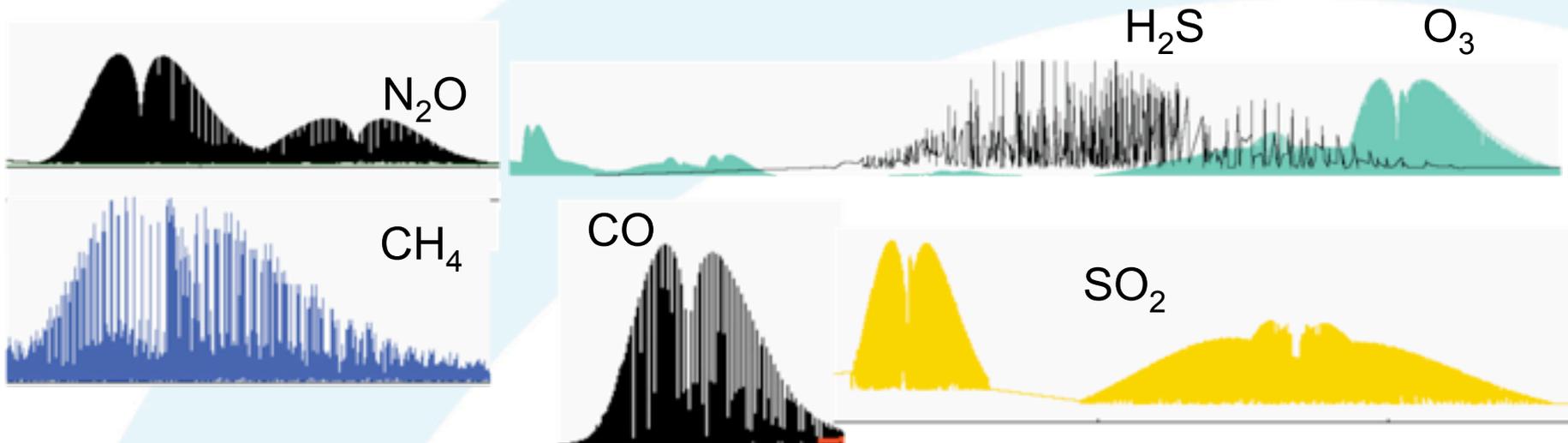
Received 23 August 2006; revised 25 June 2007; accepted 16 July 2007; published 19 October 2007.

<sup>1</sup>Norwegian Institute for Air Research, Kjeller, Norway  
<sup>2</sup>Auspace Ltd., Mitchell, ACT, Australia.



# Spectral matching

Many molecules have distinctive absorption bands



Using theoretically-derived spectral shapes, molecules can be identified by comparing (or matching) the observed shape with theory.

# Spectral matching (1)

Radiative transfer equation

$$I_\nu = I_{\nu,s} + \int_0^\infty B_\nu[T(z)] \left( \frac{\partial \tau_\nu[z, q_1(z), q_2(z) \dots q_n(z)]}{\partial z} \right) dz,$$

Choose absorption region and split integral into 3 parts:

$$\begin{aligned} I_\nu &\approx \int_0^{z_1} B_\nu[T(z)] \left( \frac{\partial \tau_\nu[z, q_2(z) \dots q_n(z)]}{\partial z} \right) dz && \text{Below gas layer} \\ &+ \int_{z_1}^{z_2} B_\nu[T(z)] \left( \frac{\partial \tau_\nu[z, q_1(z), q_2(z) \dots q_n(z)]}{\partial z} \right) dz && \text{In gas layer} \\ &+ \int_{z_2}^\infty B_\nu[T(z)] \left( \frac{\partial \tau_\nu[z, q_2(z) \dots q_n(z)]}{\partial z} \right) dz. && \text{Above gas layer} \end{aligned}$$

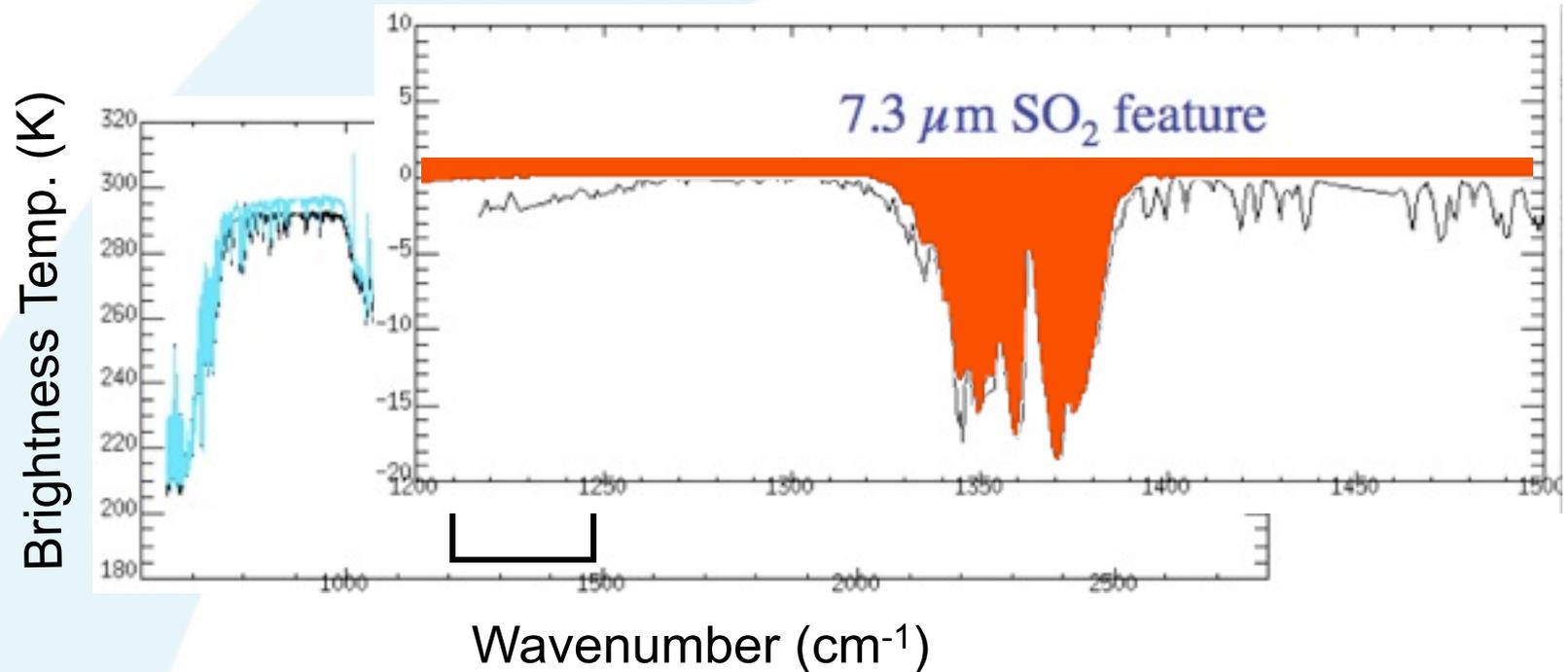
# Spectral matching (2)

$$I'_\nu = I_{\nu,0} \exp \left\{ - \int_0^\infty k_\nu(z) q(z) dz \right\}$$

Beer-Bougier-Lambert law

$$A_\nu = - \ln \left\{ \frac{I'_\nu}{I_{o,\nu}} \right\} = \int_0^\infty k_\nu(z) q(z) dz.$$

Absorbance spectrum



# Retrieval

Single Lorentz line

$$A\Delta\nu = \int_{\Delta\nu} \left\{ 1 - \exp\left[-\frac{1}{\pi} \int_{z_1}^{z_2} \frac{\alpha q(z)}{(\nu - \nu_0)^2 + \alpha^2} dz\right] \right\} d\nu,$$

$$A\Delta\nu = 2\pi\alpha\psi [L_0(\psi) + L_1(\psi)] \exp\{-\psi\},$$

$$\psi = \frac{Su}{2\pi\alpha},$$

For a band with temperature dependent line strengths and line broadening effects, a radiative transfer model must be used

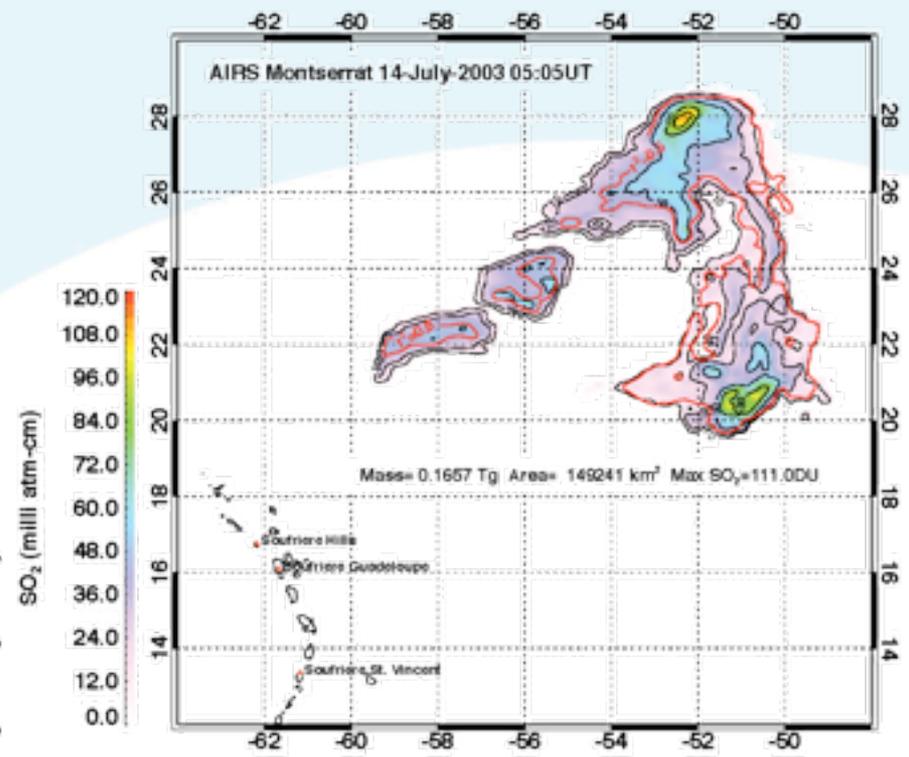
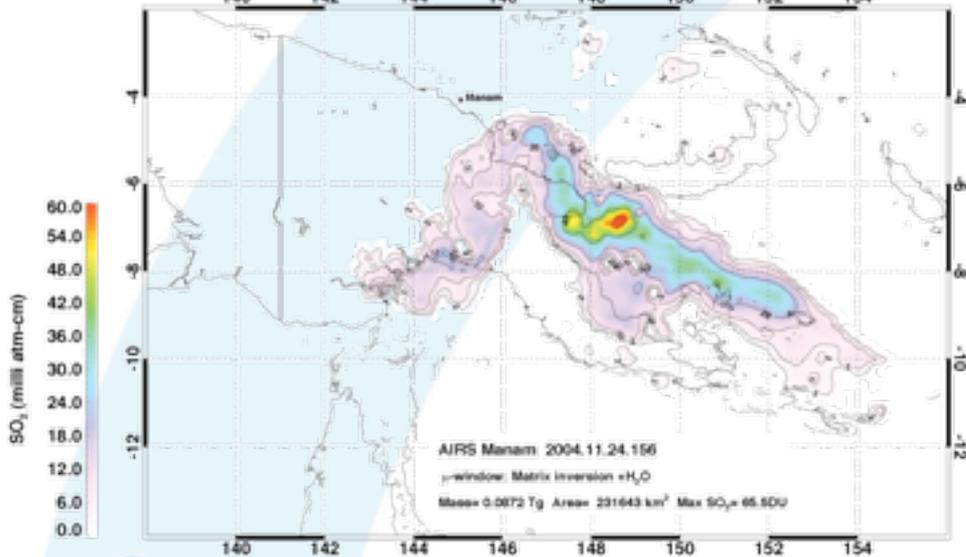
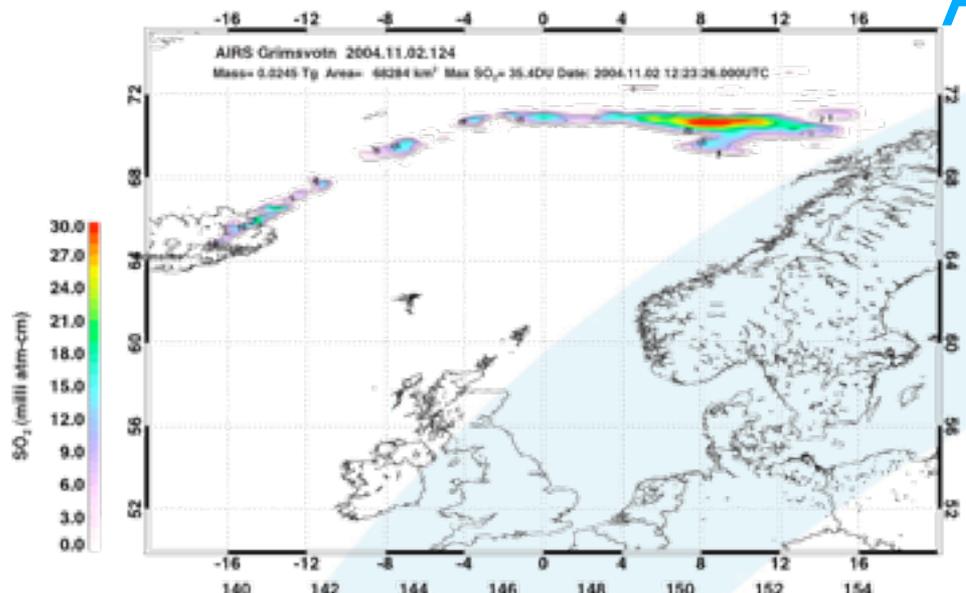
$$\vec{x} = (\mathbf{K}^T \mathbf{K})^{-1} \mathbf{K}^T \vec{y}. \quad \text{Least squares solution}$$

$\vec{x}$ : layer absorber amounts (N layers)

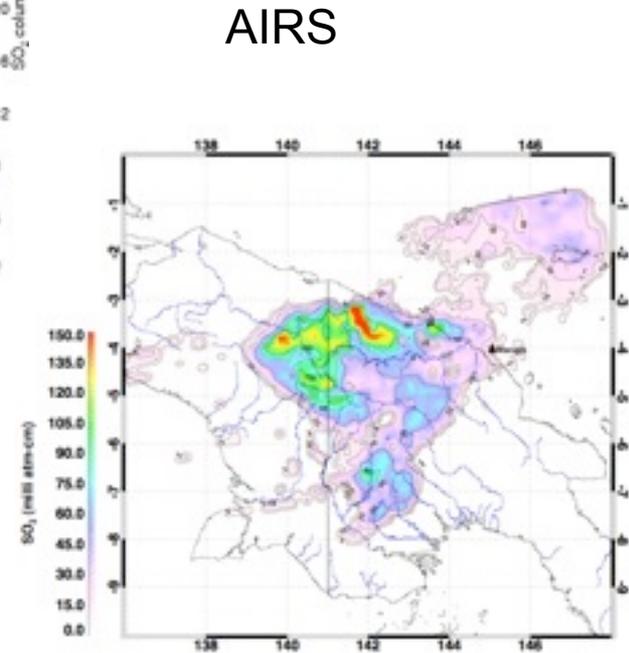
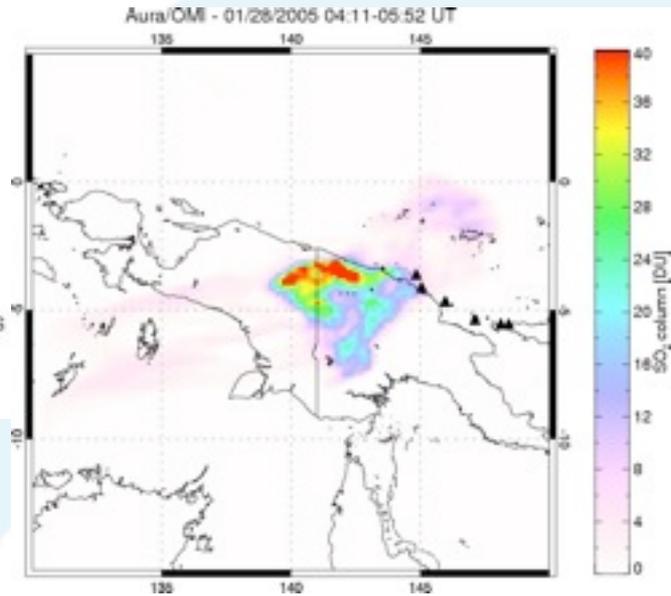
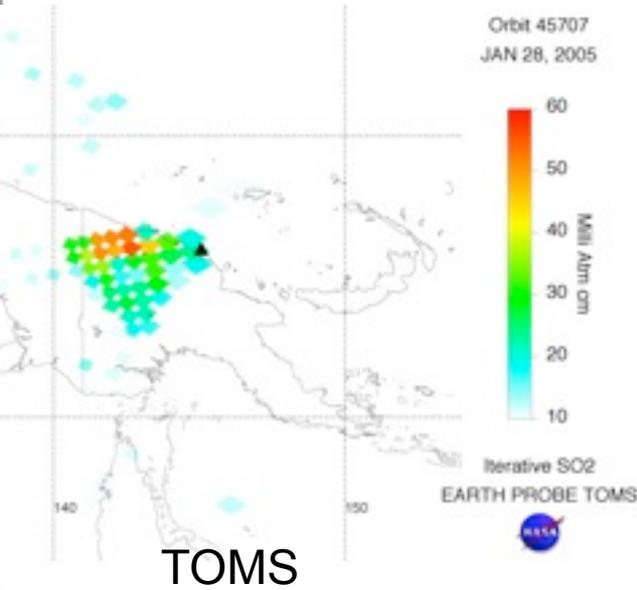
$\vec{y}$ : observed absorbances (M wavenumbers)

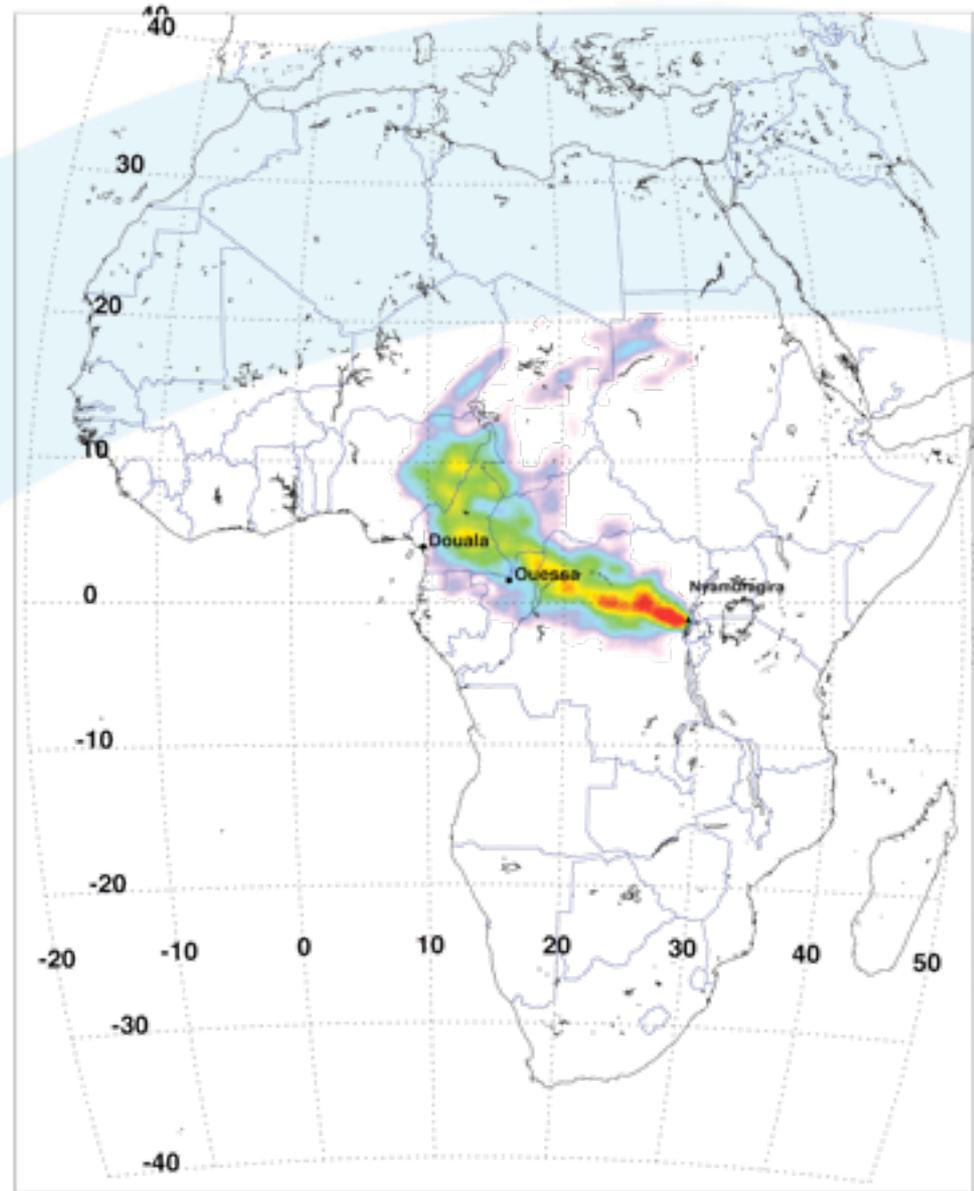
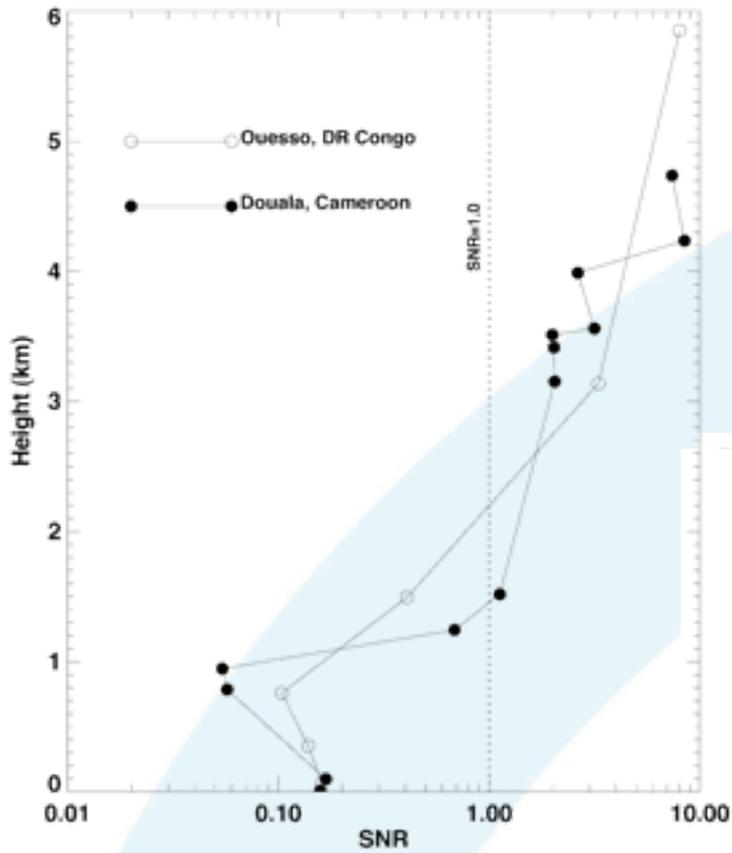
$\mathbf{K}$ : is an MxN matrix of absorption cross sections

# AIRS Volcanic SO<sub>2</sub>



# Manam Eruption Jan 28 2005





$$SNR = \frac{\sqrt{\frac{1}{\Delta\nu} \int_{\Delta\nu} \Delta S_{\nu}^2 d\nu}}{NE\Delta I}$$

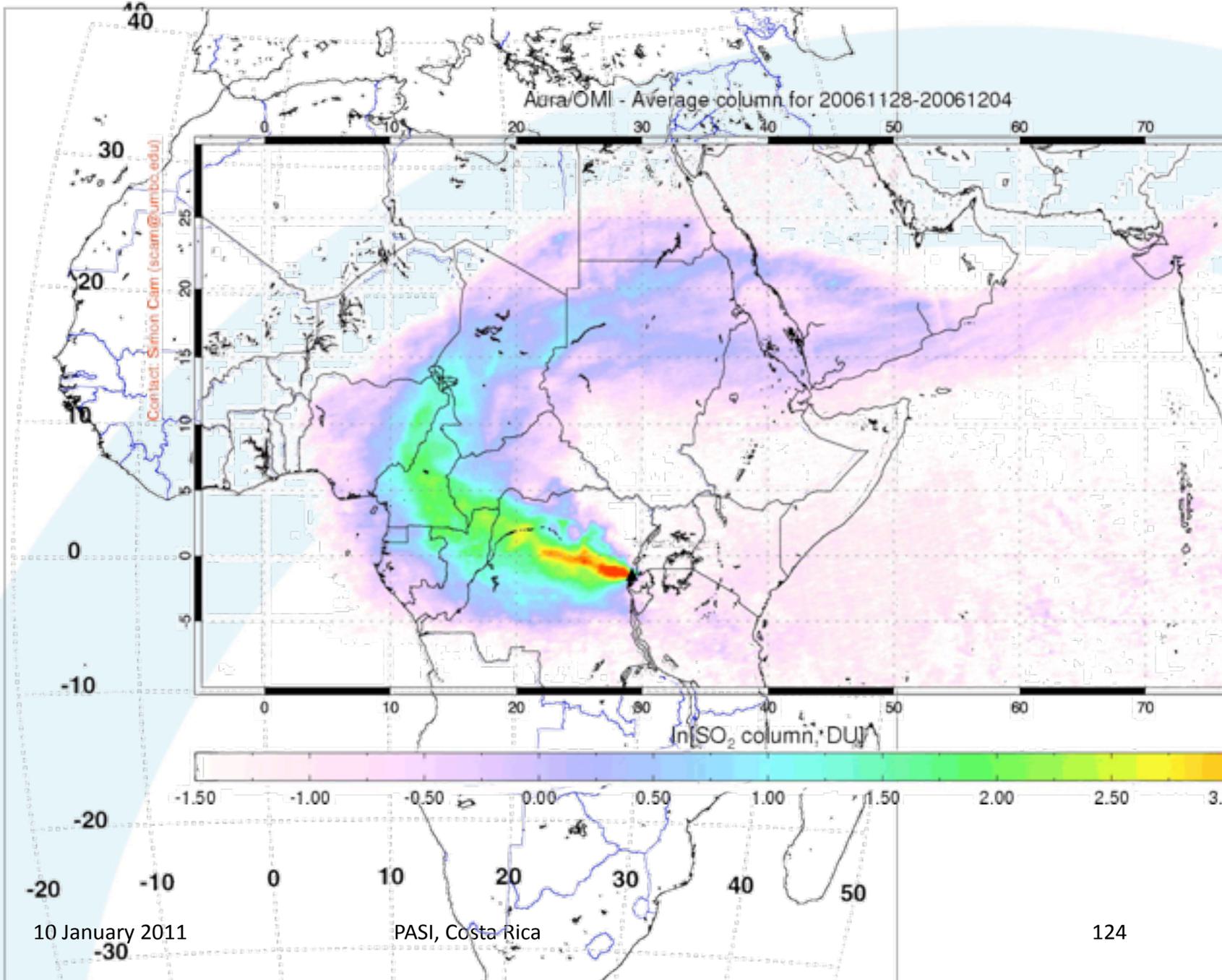
$$\Delta S_{\nu} = |I_{\nu,z'} - I_{\nu,0}|,$$

SO<sub>2</sub> (milli atm-cm)



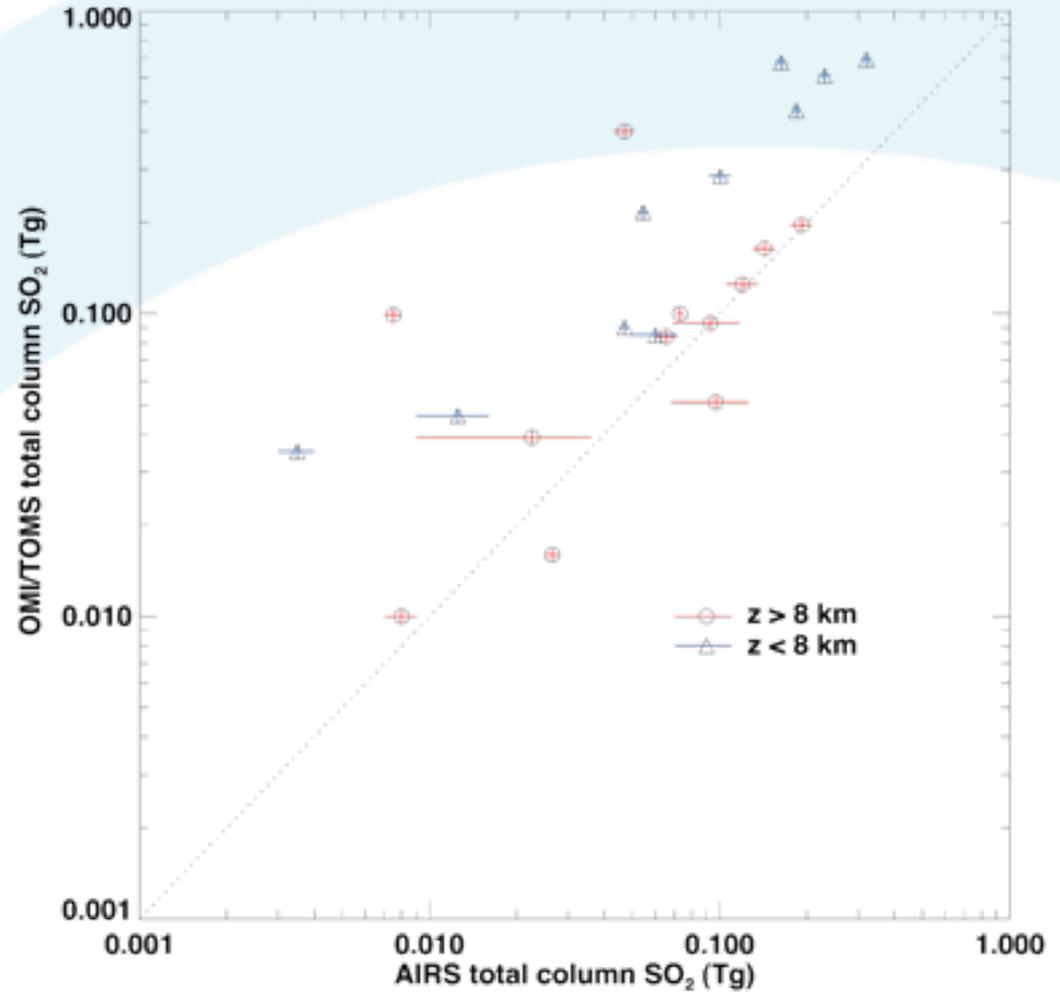
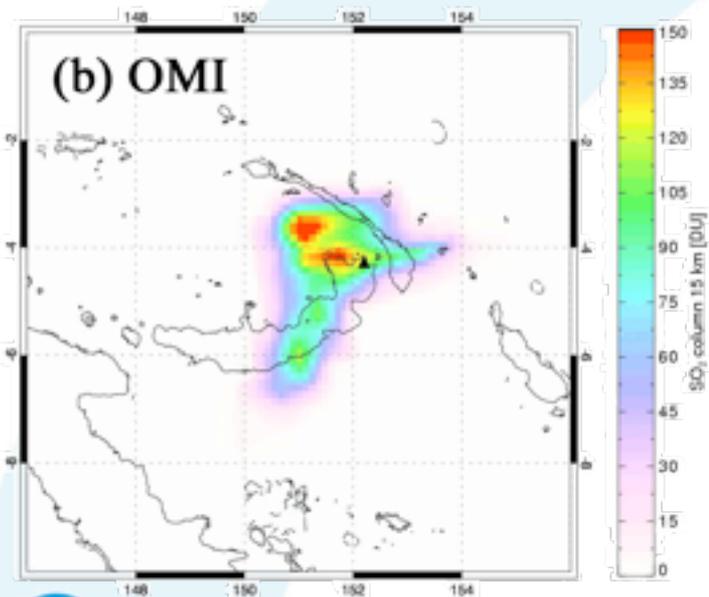
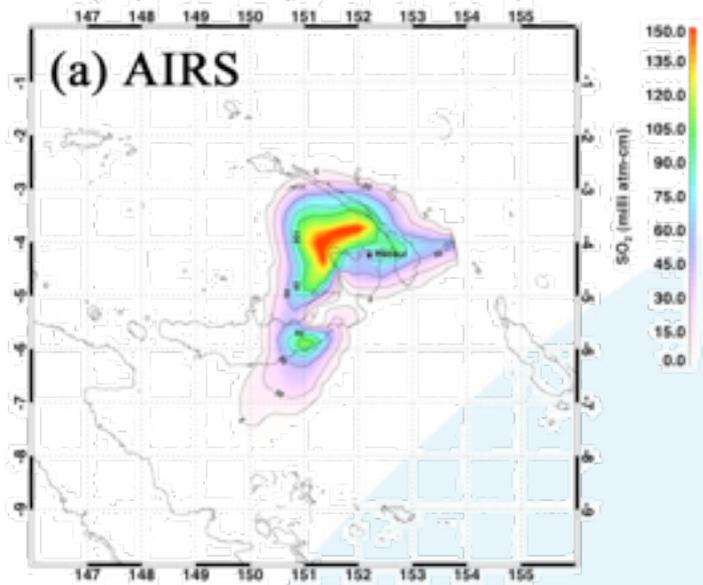
Contact: Simon Cam (scam@umbo.edu)

Aura/OMI - Average column for 20061128-20061204

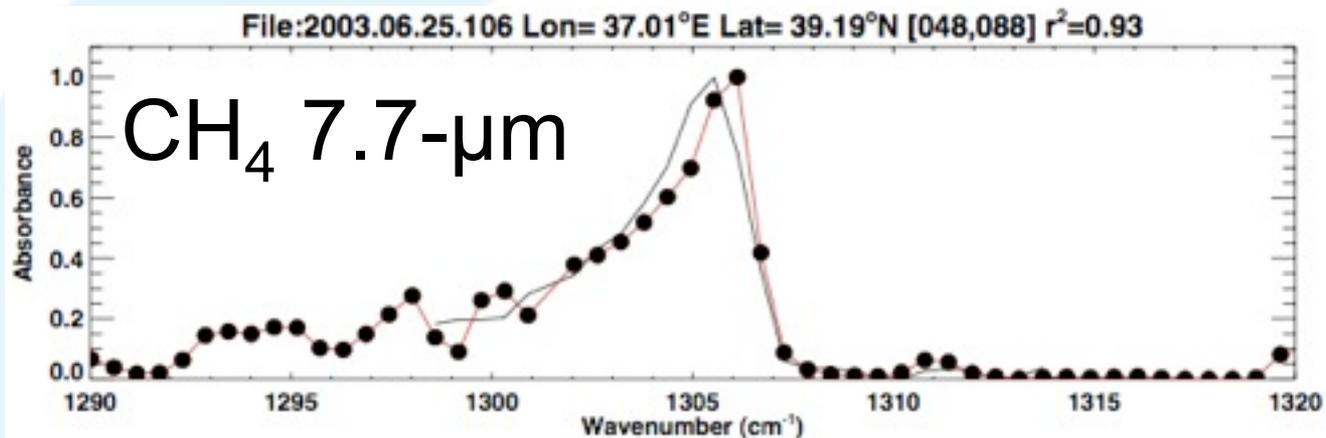
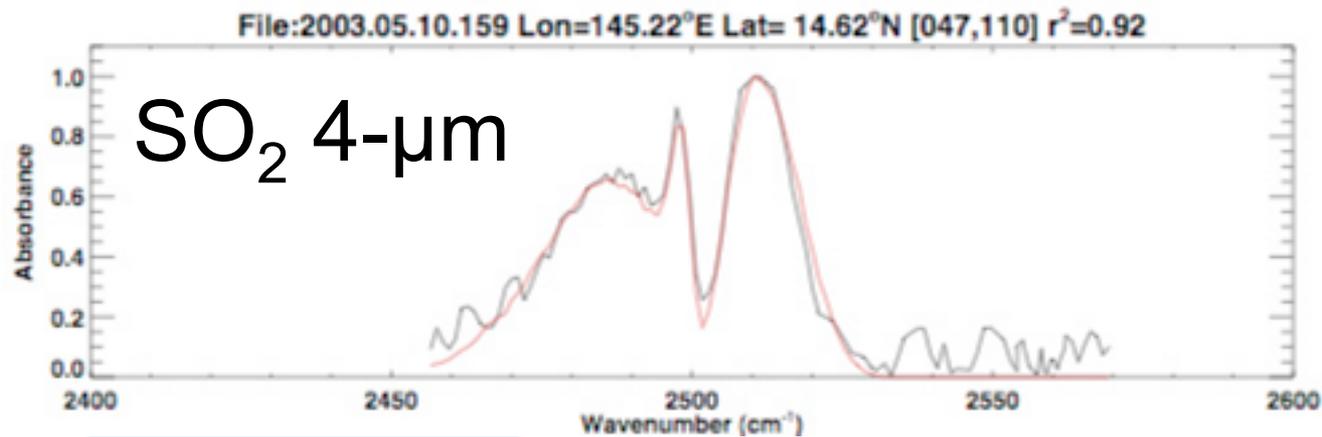


10 January 2011

PASI, Costa Rica



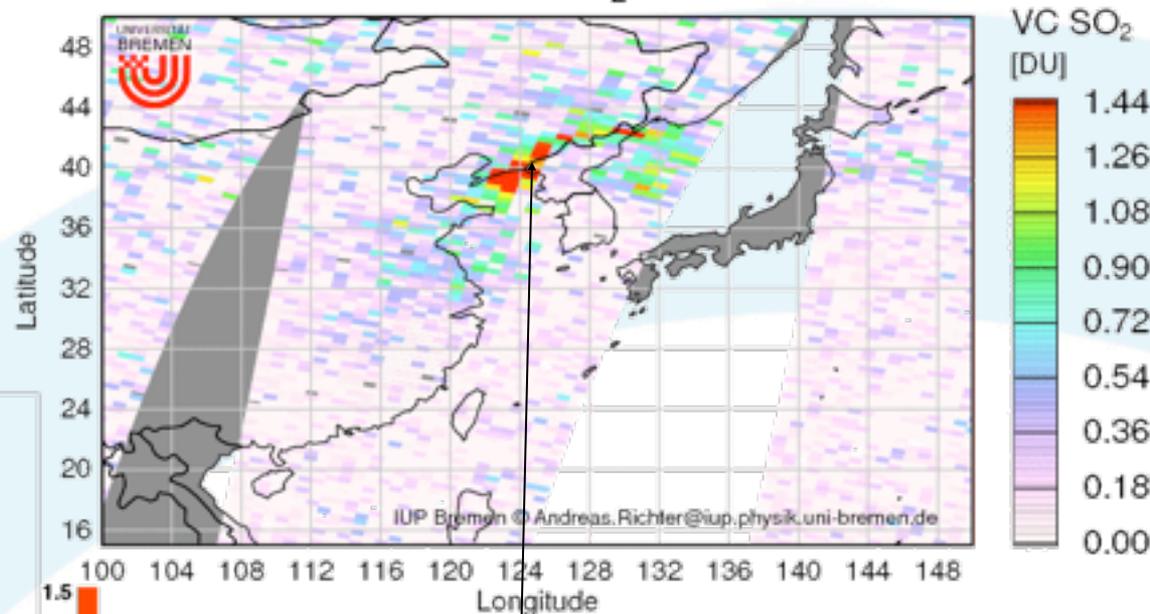
# Other spectral regions



H<sub>2</sub>S: 1332–1344 cm<sup>-1</sup>

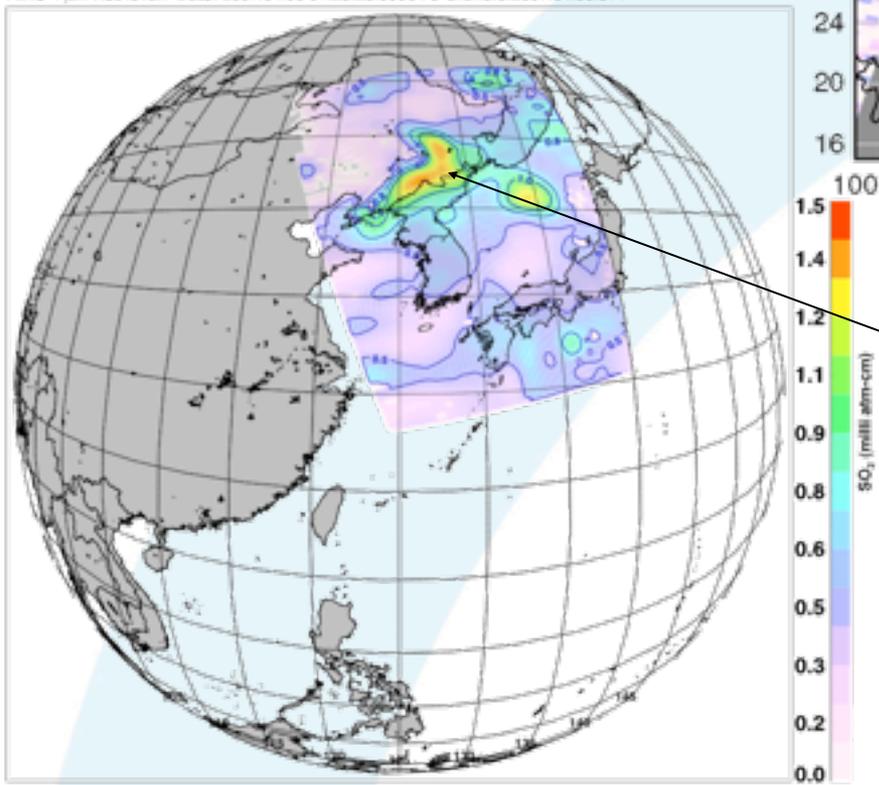


# METOP GOME2 SO<sub>2</sub>: 2007/04/06



AIRS 4 μm

AIRS 4 μm Retrieval: Date: 2007.04.06 04:23:25.000UTC Granule:2007.04.06.044



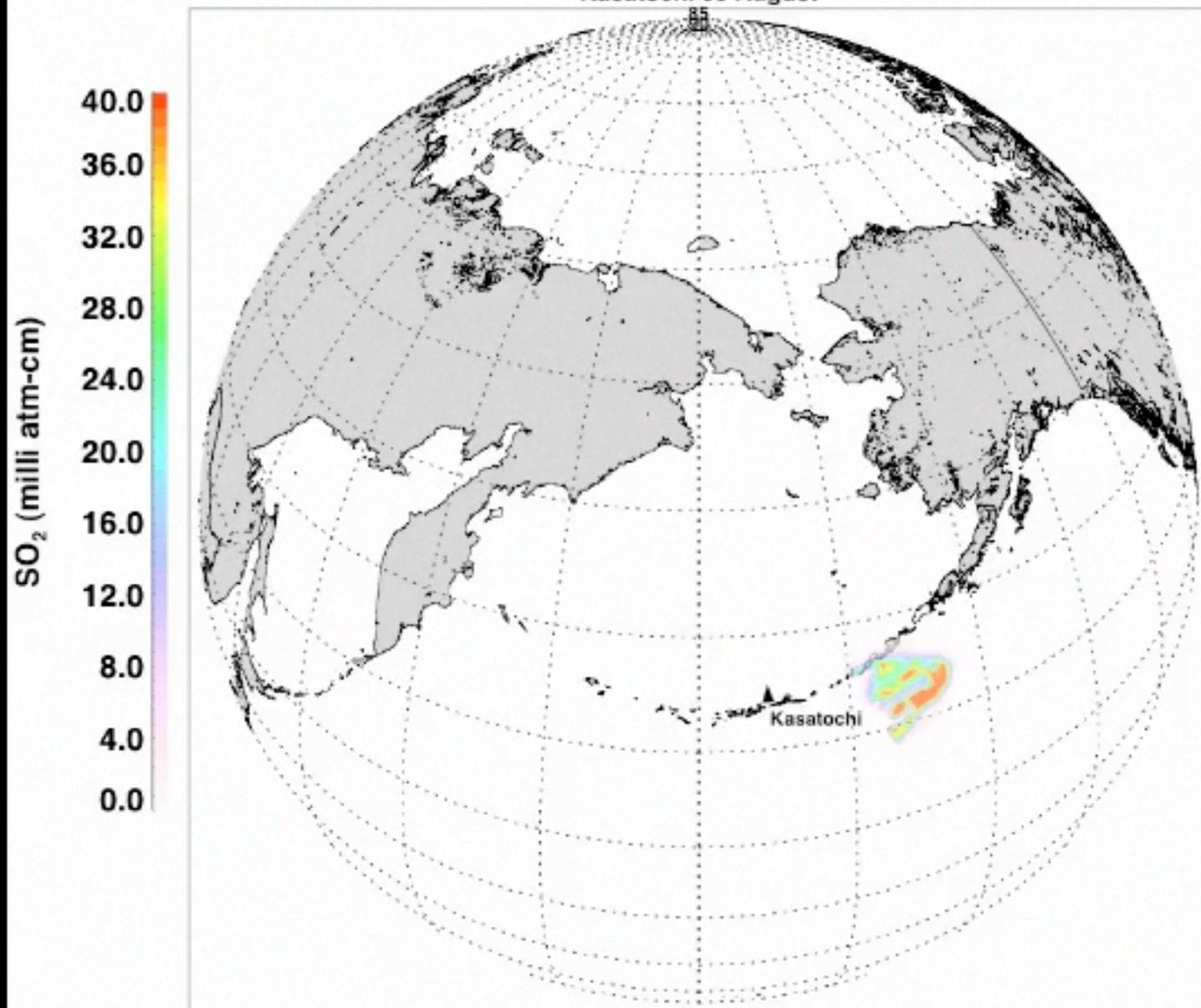
Anthropogenic SO<sub>2</sub> emissions

AIRS Lower Tropospheric SO<sub>2</sub>

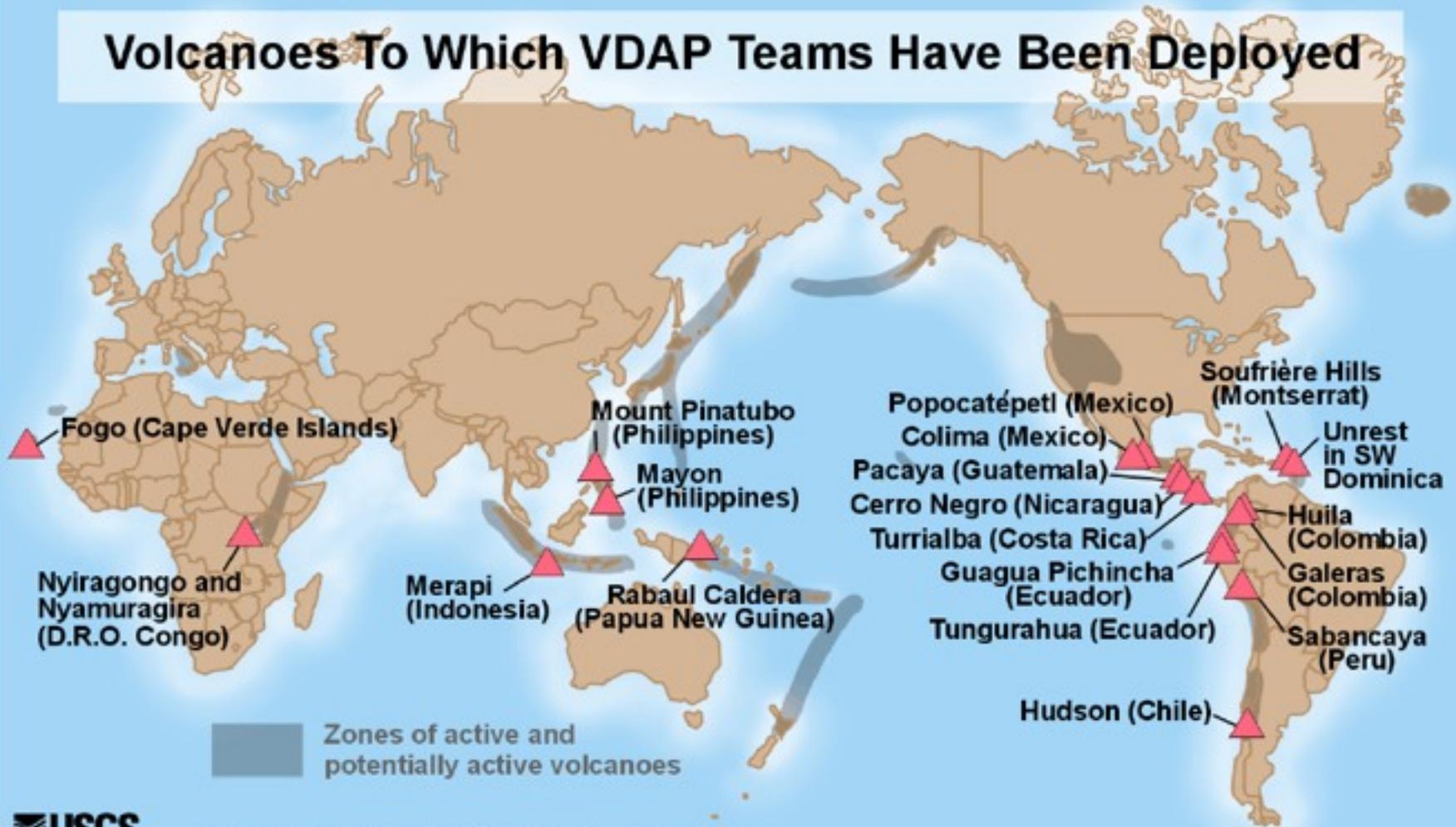
Produced by: Dr Fred Prata, NILU, (fred.prata@nilu.no)



Kasatochi 08 August

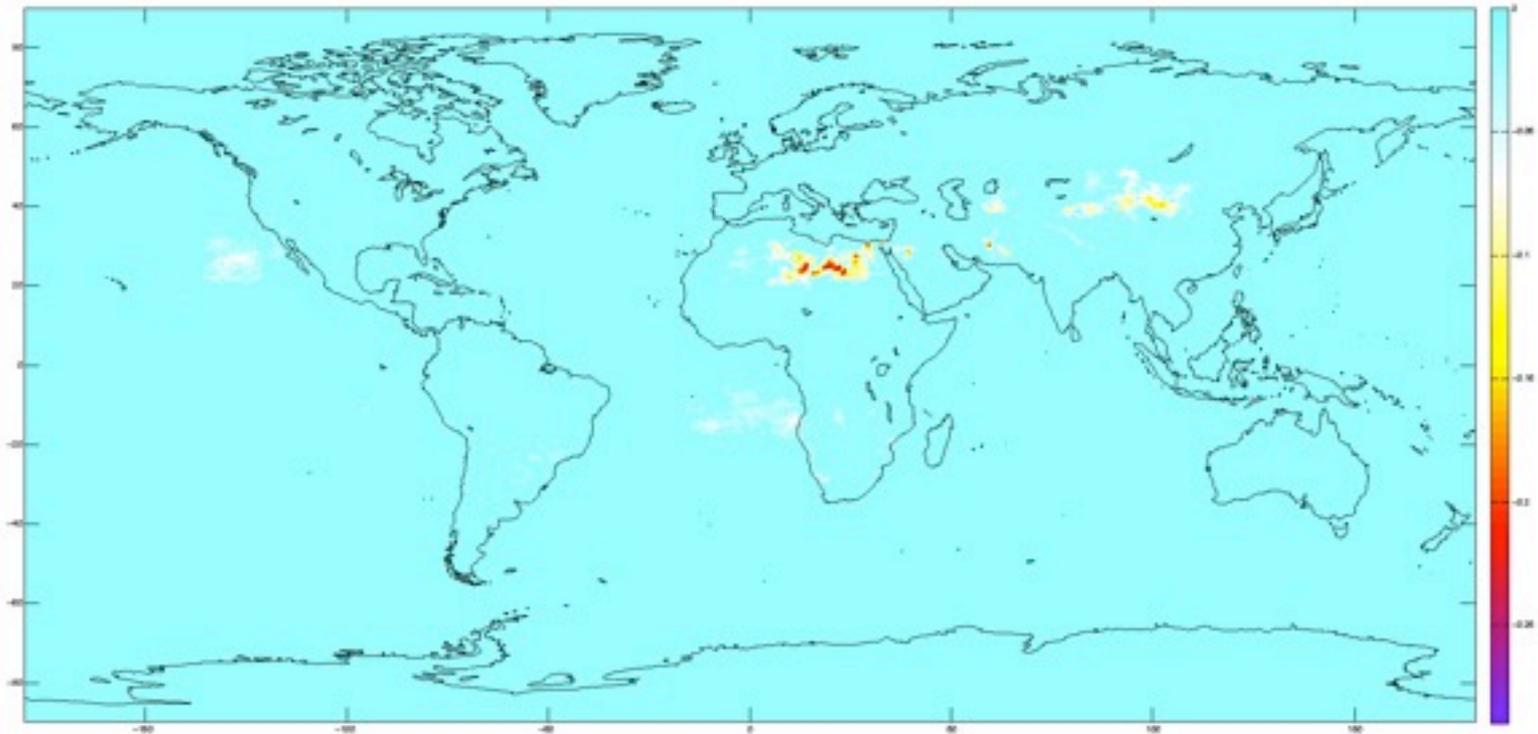


# Volcanoes To Which VDAP Teams Have Been Deployed



Updated 2/15/00 from Elvert, Miller, and others, 1997, USGS Fact Sheet 064-97

# IASI – Dust detection (based on the concavity index)



August 2008 pm.

(Courtesy Levien Clarisse)

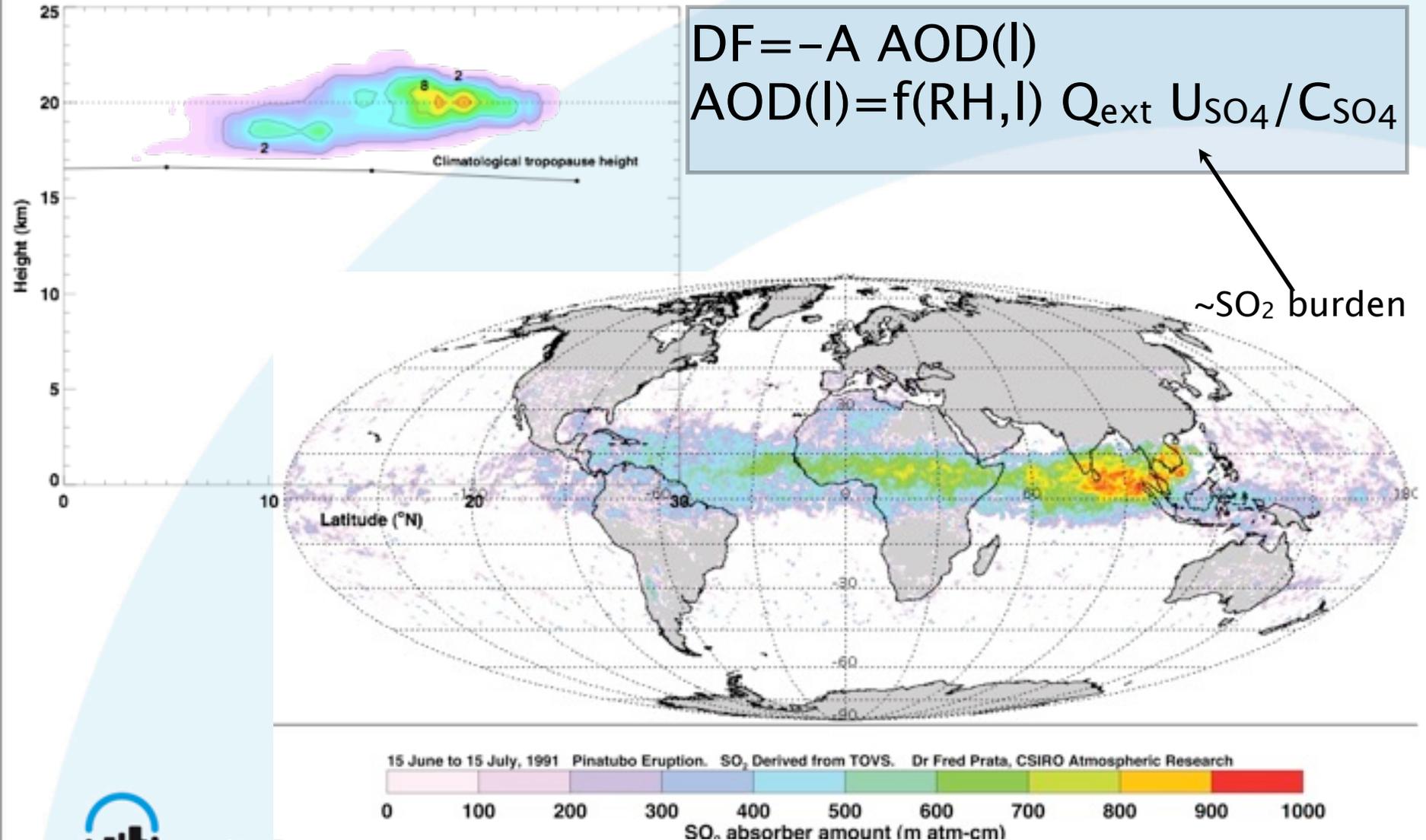
# Direct use of satellite SO<sub>2</sub> measurements

Climate sensitivity to volcanic forcing

$$DF = -A \text{ AOD}(l)$$

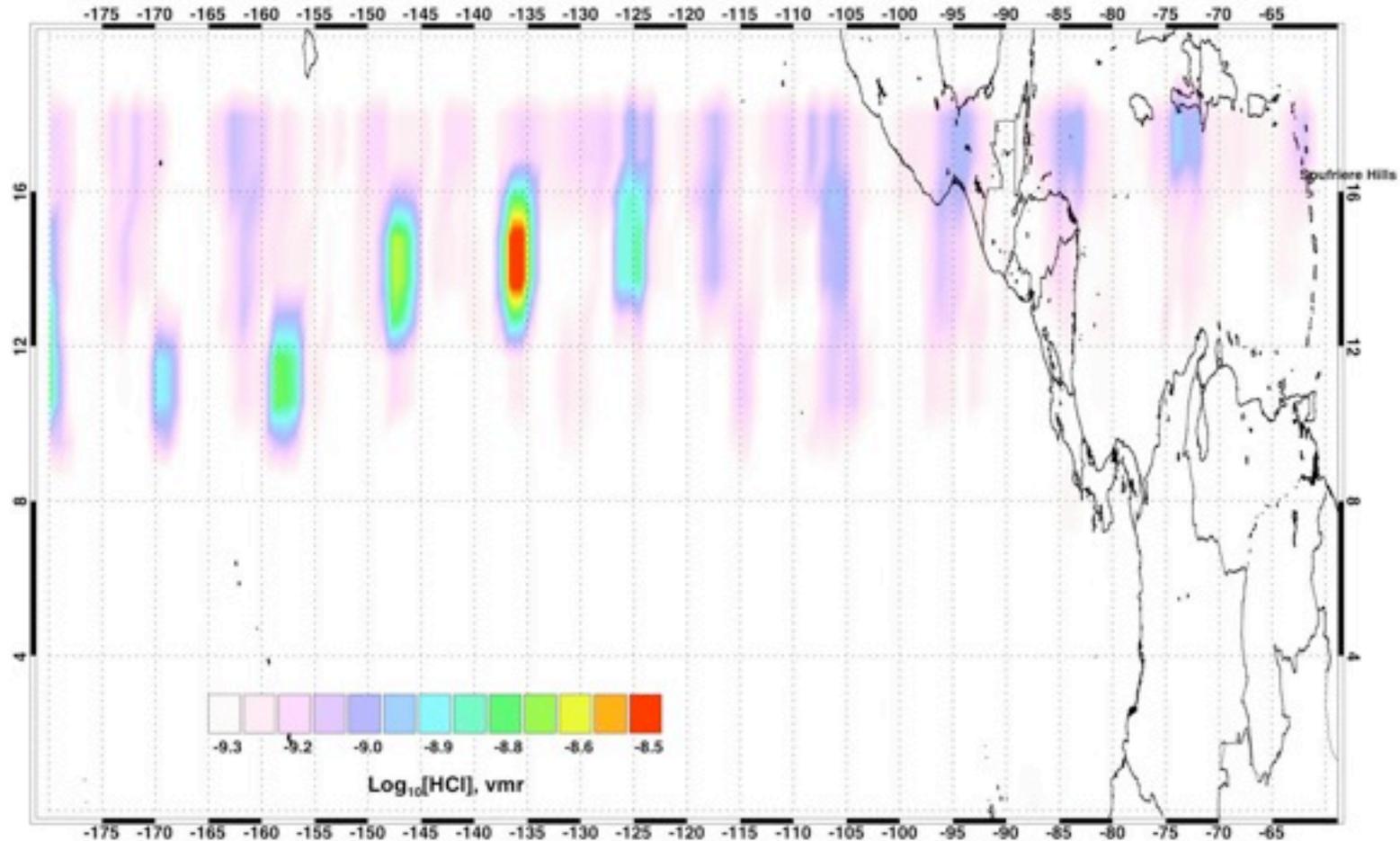
$$\text{AOD}(l) = f(\text{RH}, l) Q_{\text{ext}} U_{\text{SO}_4} / C_{\text{SO}_4}$$

~SO<sub>2</sub> burden

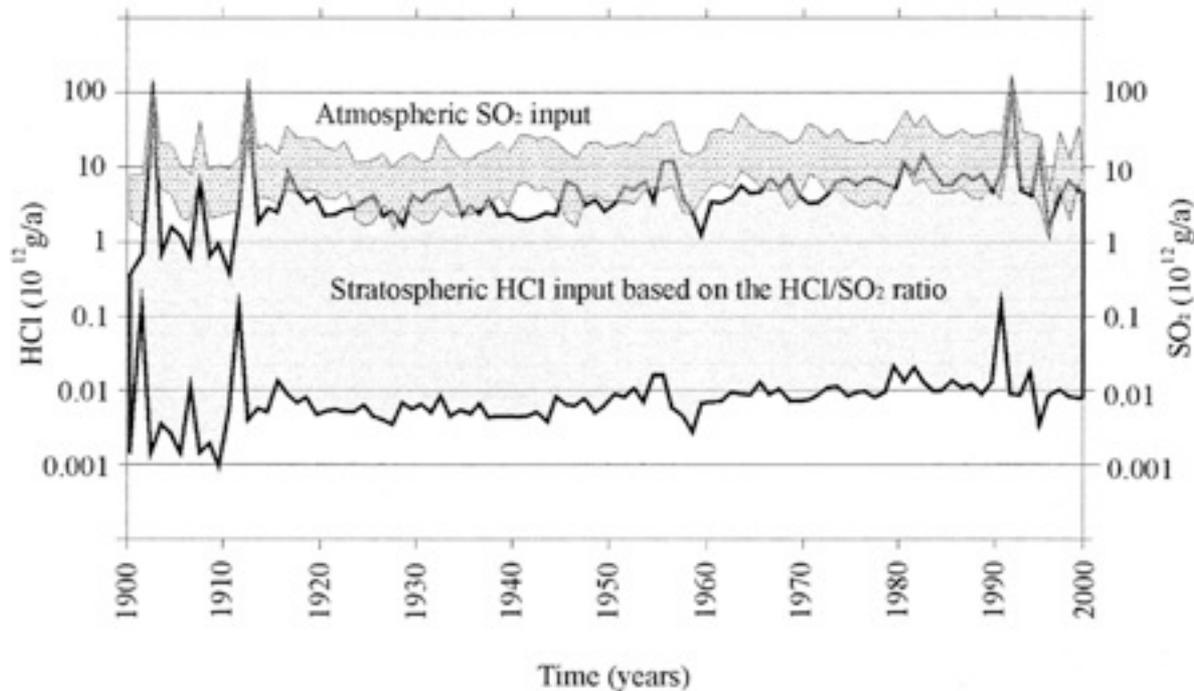


# Other gases - HCl

Cumulative MLS/Aura HCl at 68.13 hPa. 1-31 May, 2006. Soufriere Hills, Montserrat.



Prata, A. J., S. A. Carn, A. Stohl, and J. Kerkmann, 2007, Long range transport and fate of a stratospheric volcanic cloud from Soufriere Hills volcano, Montserrat, *Atmos. Chem. Phys.*, 7, 5093–5103.



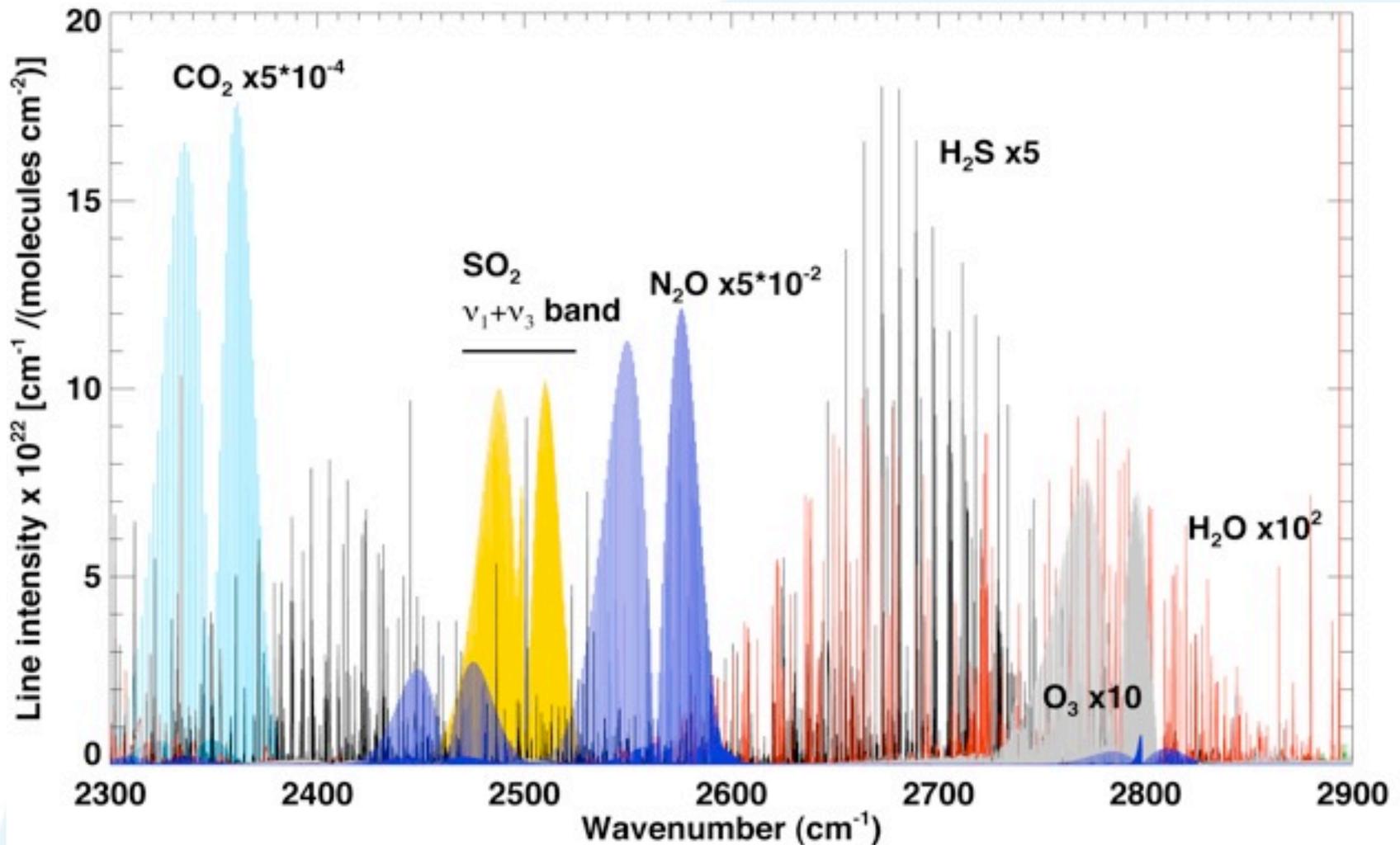
Assumes molar ratio of  $\sim 0.01$  (HCl:SO<sub>2</sub>)

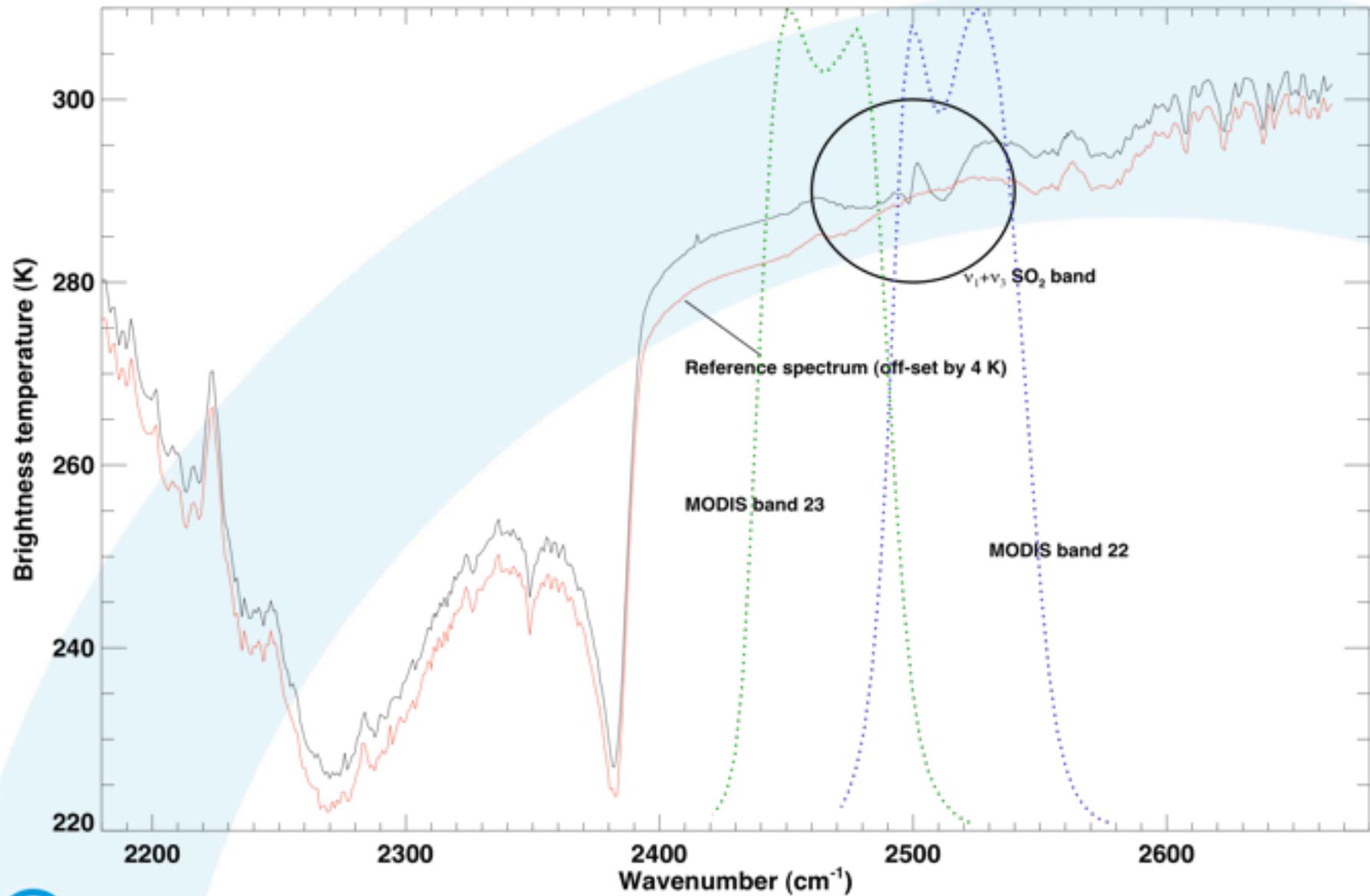
Prata *et al.*, (2007) find  $\sim 0.03$ - $0.1$  using MLS and AIRS data.

Annual global volcanic gas input into the atmosphere from 1972 to 2000 (this work)

	SO <sub>2</sub> (10 <sup>12</sup> g/yr)	H <sub>2</sub> S (10 <sup>12</sup> g/yr)	HCl (10 <sup>12</sup> g/yr)	HF (10 <sup>12</sup> g/yr)	HBr (10 <sup>9</sup> g/yr)
Atmosphere	15.0–21.0	1.5–37.1	1.2–170	0.7–8.6	2.6–43.2
Stratosphere	2.4–4.0	0.1–5.0	0.1–12	0.03–0.4	0.04–0.07
Stratosphere with washout effect	2.1–3.2 ( $\sim 25\%$ )	0.01–3.8 ( $\sim 25\%$ )	0.1–12 $\times 10^{-3}$ ( $\sim 99.99\%$ )	0.03–0.4 $\times 10^{-3}$ ( $\sim 99.99\%$ )	0.04–0.07 $\times 10^{-3}$ ( $\sim 99.99\%$ )

# Boundary layer SO<sub>2</sub> from AIRS



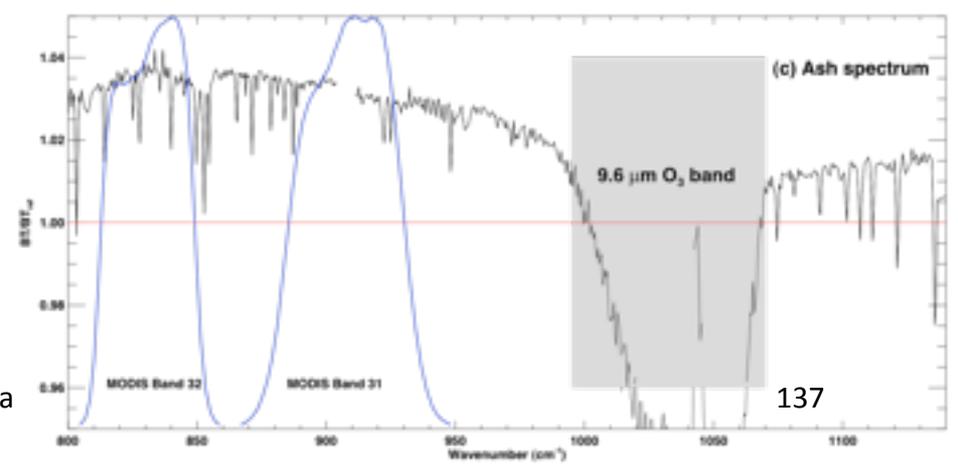
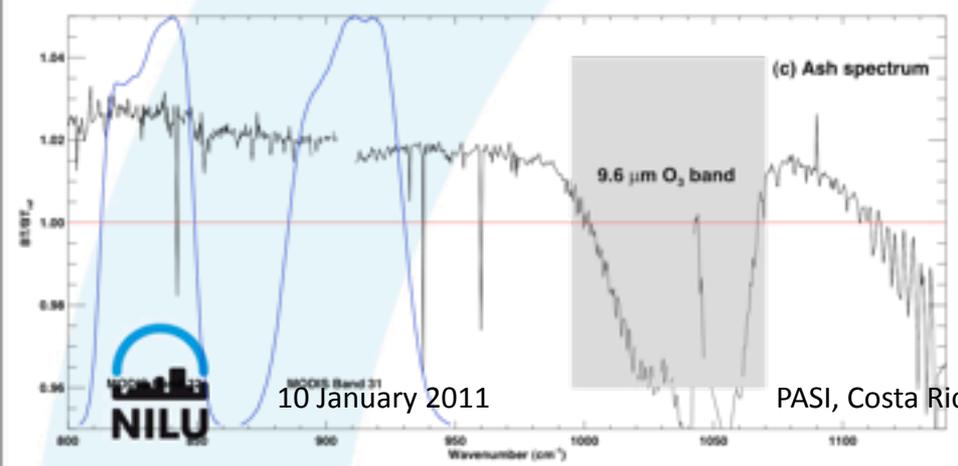
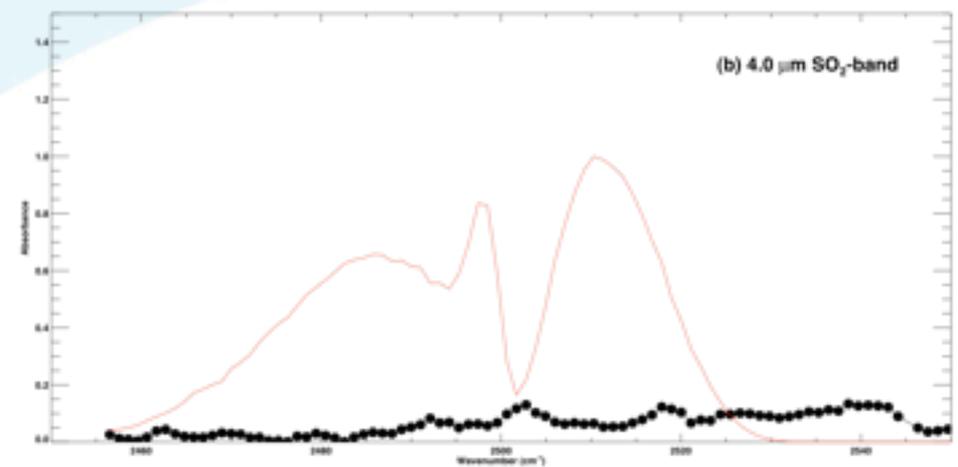
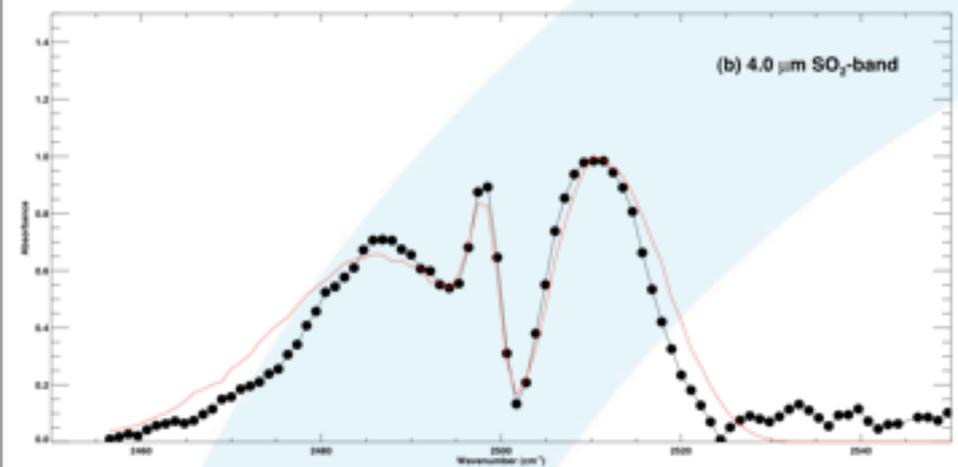
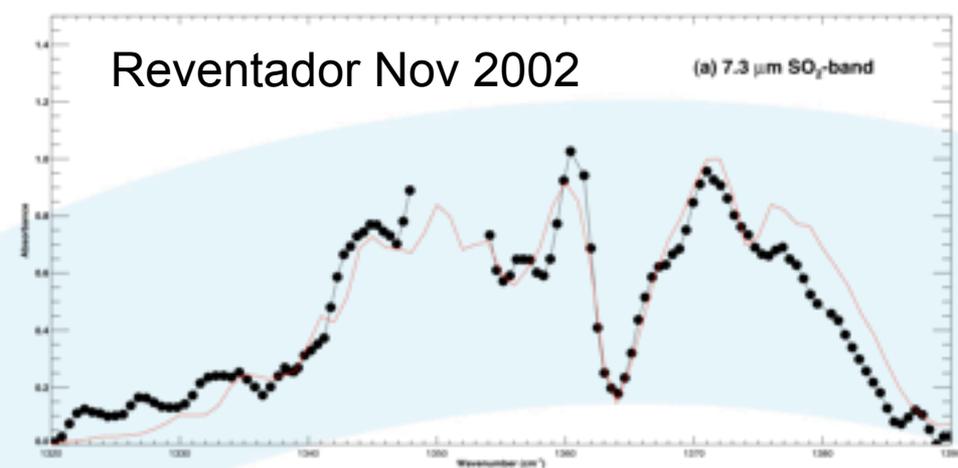
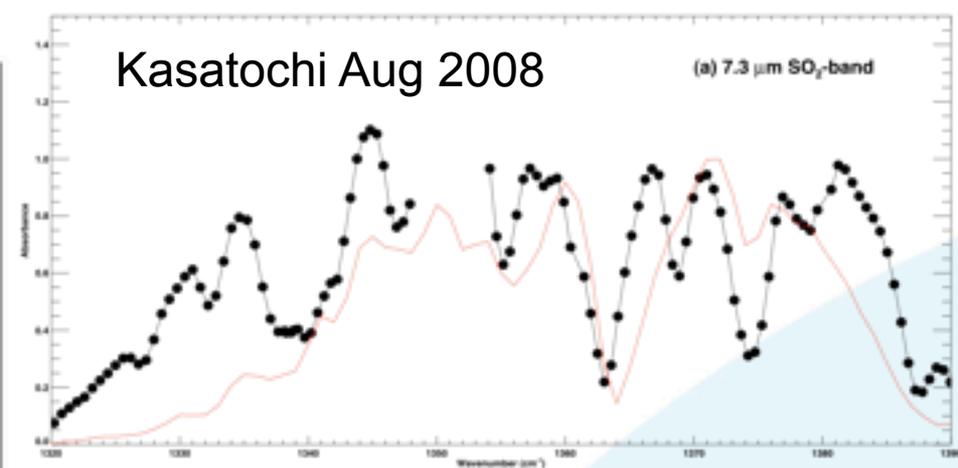


Kasatochi Aug 2008

(a) 7.3  $\mu\text{m}$  SO<sub>2</sub>-band

Reventador Nov 2002

(a) 7.3  $\mu\text{m}$  SO<sub>2</sub>-band



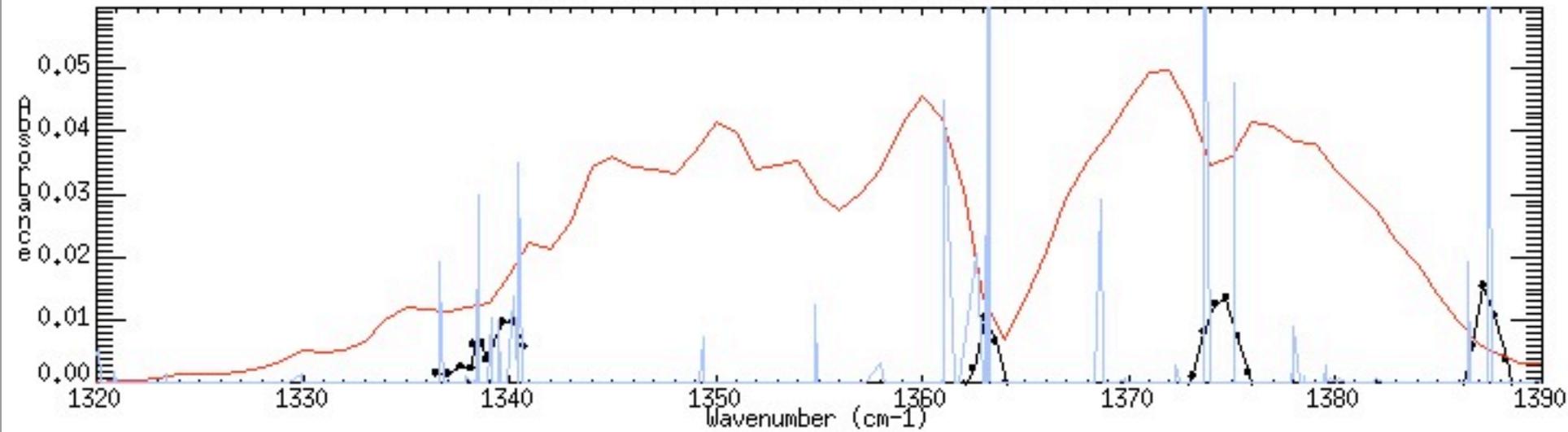
10 January 2011

PASI, Costa Rica

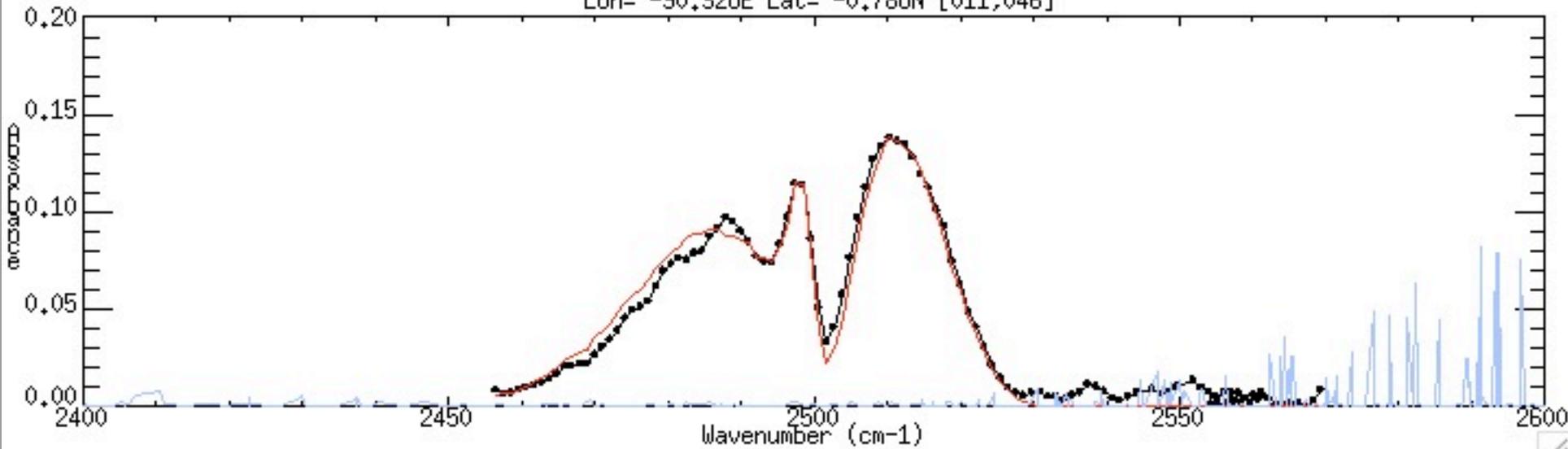
137

IDL 1

Lon= -90.320E Lat= -0.780N [011,046] r2=0.08

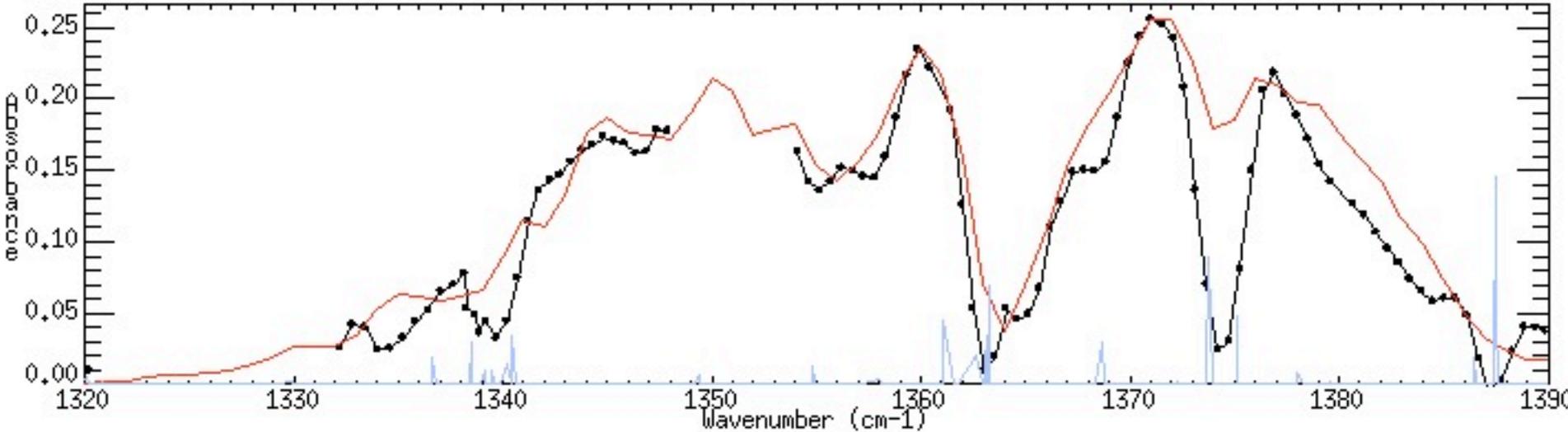


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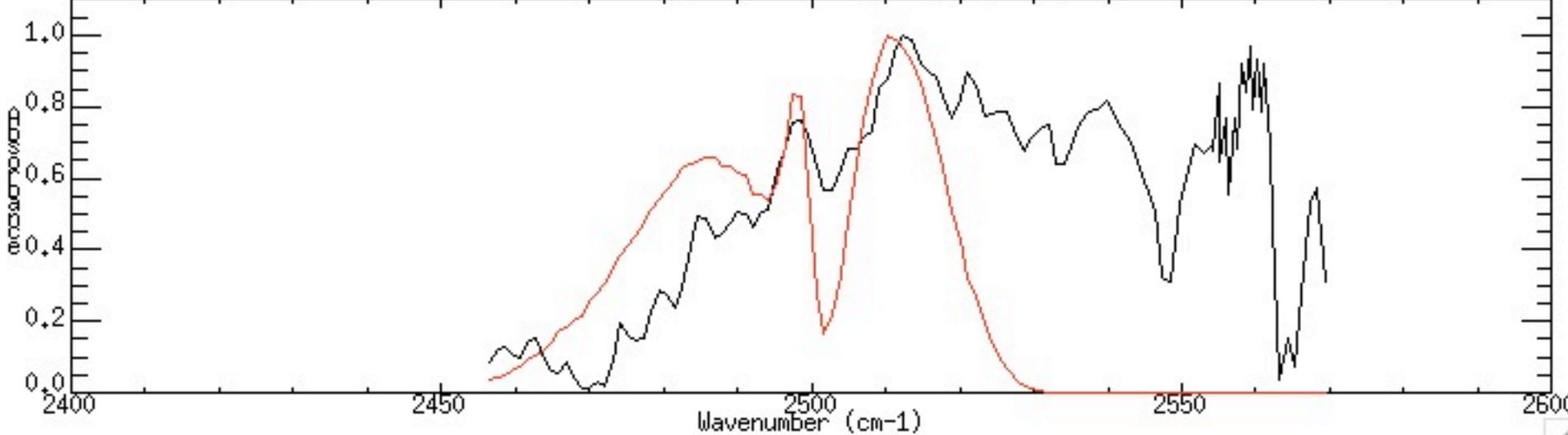


IDL 1

Lon= -92.76oE Lat= -0.48oN [002,050] r2=0.85

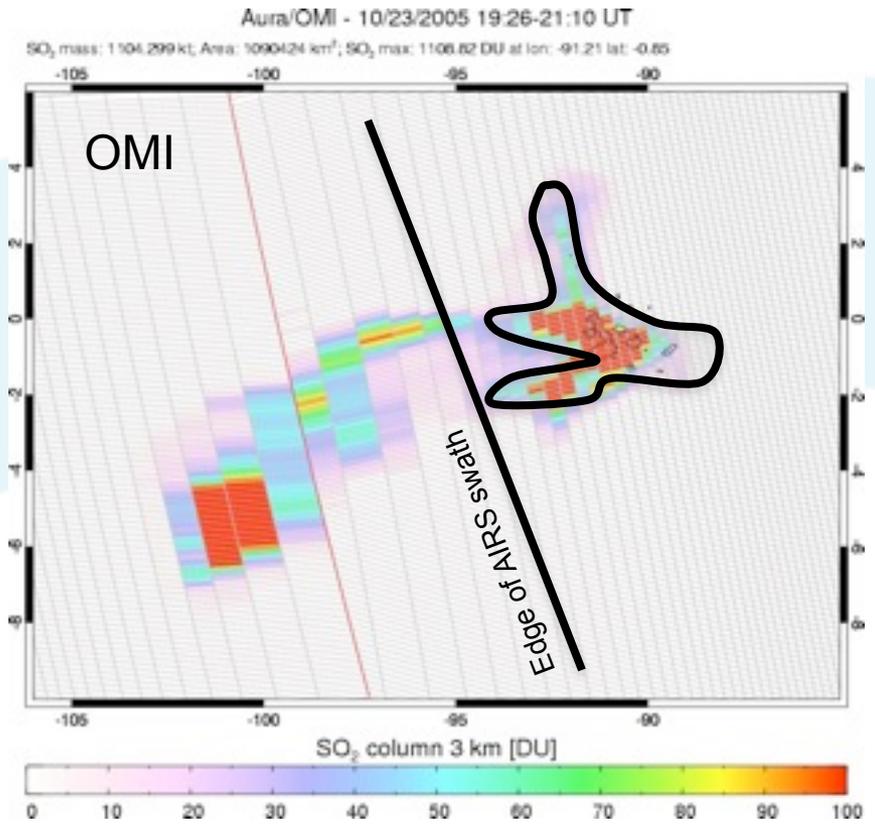
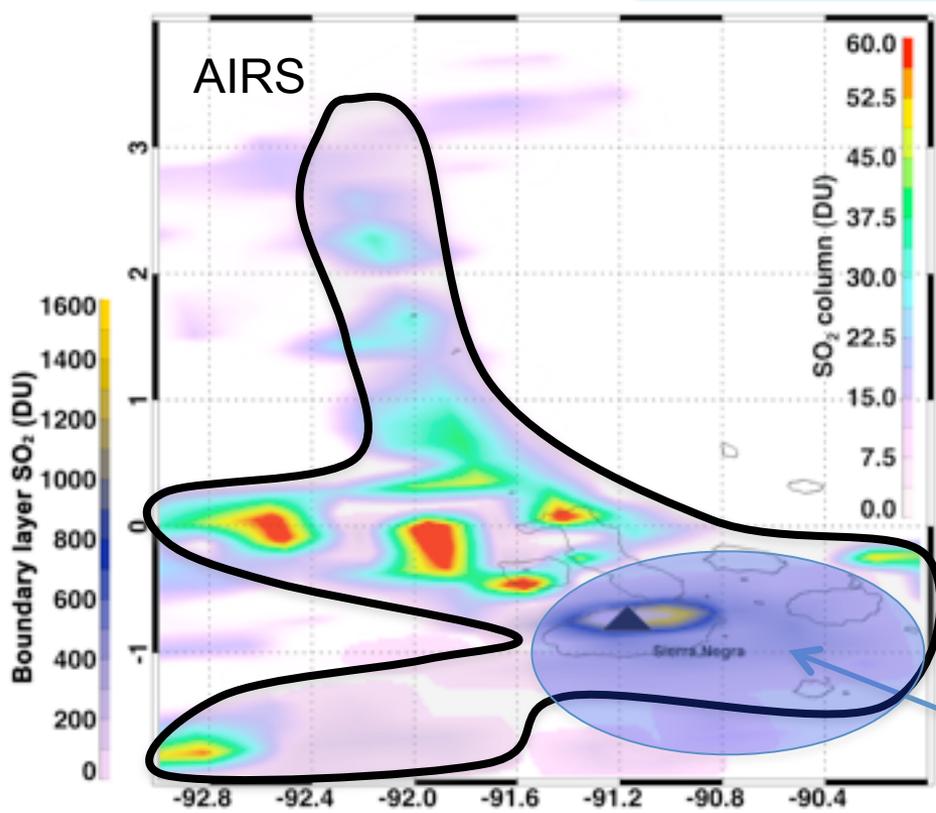


Lon= -92.76oE Lat= -0.48oN [002,050]



# Sierra Negra

23.10.2005



Boundary layer SO<sub>2</sub>?

# European volcano-related projects

Project	Lead (NILU involvement)	Funding (M€)	Funding Agency	Duration (yrs)
GlobVolcano	Italy (x)	0.8	ESA	3
Exupery	DLR (x)	0.3	ESA	3
SAVAA	NILU (✓)	0.5	ESA	3
MIA-VITA	BRGM-France (✓)	5	FP7	3
SAFER	French Consortium (✓)	23	FP7	5
PGP	Oslo U (✓)	5	NFR	5
NOVAC	Chalmers (x)	5	FP5	5

# Further information

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- **Prata, A. J.**, 2008, Satellite detection of hazardous volcanic clouds and the risk to global air traffic, *Nat. Hazards*, DOI 10.1007/s11069-008-9273-z.

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TOVS